# Solutions of the exercises of the book "Mechanics of Continuous Media: an Introduction" 

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## Preface

We taught mechanics of continuous media for two decades at EPFL (Swiss Institute of Technology in Lausanne) and we used a lot of various exercises to illustrate the subject.

We have decided to make available these solutions in order to help as much as possible the understanding of the lecture contents and mastering of the concepts. The reader will note that some solutions are sometimes simple to set up. Others are more elaborate, need longer development and are more difficult to tackle.

We will refer to the equations of the monograph "Mechanics of Continuous Media : an Introduction" published by EPFL Press, by prefixing their numbers with the bold character B for Book. The present book of solutions possesses its own numbering.

## Acknowledgements

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## Chapitre 1

## Cartesian tensors

## Solution 1.1

Through the relations (B1.52) and (B1.15), one has

$$
\delta_{p q}^{\prime}=c_{p i} c_{q j} \delta_{i j}=c_{p j} c_{q j}=\delta_{p q} .
$$

## Solution 1.2

By the Kronecker symbol properties, one obtains

$$
\begin{aligned}
\delta_{i j} \delta_{i k} \delta_{j k} & =\delta_{11} \delta_{11} \delta_{11}+\delta_{22} \delta_{22} \delta_{22}+\delta_{33} \delta_{33} \delta_{33}= \\
& =1+1+1 \\
& =3 .
\end{aligned}
$$

One can also demonstrate the relation as follows

$$
\delta_{i j} \delta_{i k} \delta_{j k}=\delta_{i j} \delta_{i j}=\delta_{i i}=3 .
$$

## Solution 1.3

For the first relation, one has

$$
\begin{gathered}
\varepsilon_{i j k} u_{i} u_{j}=\varepsilon_{123} u_{1} u_{2}+\varepsilon_{231} u_{2} u_{3}+\varepsilon_{312} u_{3} u_{1} \\
+\varepsilon_{132} u_{1} u_{3}+\varepsilon_{321} u_{3} u_{2}+\varepsilon_{213} u_{2} u_{1}=0 .
\end{gathered}
$$

For the second relation, as

$$
\begin{cases}\delta_{i j} \neq 0 & \text { if } i=j \\ \varepsilon_{i j k}=0 & \text { if } i=j\end{cases}
$$

it follows that

$$
\delta_{i j} \varepsilon_{i j k}=0 .
$$

## Solution 1.4

It is asked to demonstrate the following relation

$$
\begin{equation*}
\boldsymbol{t} \times(\boldsymbol{u} \times \boldsymbol{v})=(\boldsymbol{t} \cdot \boldsymbol{v}) \boldsymbol{u}-(\boldsymbol{t} \cdot \boldsymbol{u}) \boldsymbol{v} \tag{1.1}
\end{equation*}
$$

By definition of vector product (B1.31), one gets

$$
(\boldsymbol{u} \times \boldsymbol{v})_{k}=\varepsilon_{k l m} u_{l} v_{m}
$$

Therefore the left hand side of (1.1) yields

$$
\begin{equation*}
(\boldsymbol{t} \times(\boldsymbol{u} \times \boldsymbol{v}))_{i}=\varepsilon_{i j k} t_{j}\left(\varepsilon_{k l m} u_{l} v_{m}\right)=\varepsilon_{i j k} \varepsilon_{k l m} t_{j} u_{l} v_{m} \tag{1.2}
\end{equation*}
$$

By (B1.30), we know that

$$
\varepsilon_{i j k} \varepsilon_{k l m}=\delta_{i l} \delta_{j m}-\delta_{i m} \delta_{j l}
$$

Combining (1.2) and (B1.30), one obtains successively

$$
\begin{aligned}
\varepsilon_{i j k} \varepsilon_{k l m} t_{j} u_{l} v_{m} & =\left(\delta_{i l} \delta_{j m}-\delta_{i m} \delta_{j l}\right) t_{j} u_{l} v_{m} \\
& =\delta_{i l} \delta_{j m} t_{j} u_{l} v_{m}-\delta_{i m} \delta_{j l} t_{j} u_{l} v_{m} \\
& =u_{i} t_{m} v_{m}-v_{i} t_{j} u_{j} \\
& =(\boldsymbol{t} \cdot \boldsymbol{v}) \boldsymbol{u}-(\boldsymbol{t} \cdot \boldsymbol{u}) \boldsymbol{v}
\end{aligned}
$$

## Solution 1.5

Note that

$$
(\boldsymbol{a} \times \boldsymbol{b})_{i}=\varepsilon_{i j k} a_{j} b_{k} \quad \text { and } \quad(\boldsymbol{c} \times \boldsymbol{d})_{l}=\varepsilon_{l m n} c_{m} d_{n}
$$

Then one writes

$$
((\boldsymbol{a} \times \boldsymbol{b}) \times(\boldsymbol{c} \times \boldsymbol{d}))_{r}=\varepsilon_{r i l}\left(\varepsilon_{i j k} a_{j} b_{k}\right)\left(\varepsilon_{l m n} c_{m} d_{n}\right)
$$

Let us note that $\varepsilon_{r i l}=\varepsilon_{i l r}=\varepsilon_{l r i}$. Consequently

$$
\varepsilon_{l r i}\left(\varepsilon_{i j k} a_{j} b_{k}\right)\left(\varepsilon_{l m n} c_{m} d_{n}\right)=\varepsilon_{l r i} \varepsilon_{l m n} \varepsilon_{i j k} a_{j} b_{k} c_{m} d_{n}
$$

Using identity $(\mathbf{B} 1.30), \varepsilon_{l r i} \varepsilon_{l m n}=\delta_{r m} \delta_{i n}-\delta_{r n} \delta_{i m}$, one obtains

$$
\begin{aligned}
\varepsilon_{r i l}\left(\varepsilon_{i j k} a_{j} b_{k}\right)\left(\varepsilon_{l m n} c_{m} d_{n}\right) & =\left(\delta_{r m} \delta_{i n}-\delta_{r n} \delta_{i m}\right) \varepsilon_{i j k} a_{j} b_{k} c_{m} d_{n} \\
& =\varepsilon_{i j k} \delta_{r m} \delta_{i n} a_{j} b_{k} c_{m} d_{n}-\varepsilon_{i j k} \delta_{r n} \delta_{i m} a_{j} b_{k} c_{m} d_{n} \\
& =\varepsilon_{i j k} a_{j} b_{k} c_{r} d_{i}-\varepsilon_{i j k} a_{j} b_{k} c_{i} d_{r} \\
& =\varepsilon_{j k i} b_{k} d_{i} a_{j} c_{r}-\varepsilon_{j k i} b_{k} c_{i} a_{j} d_{r} \\
& =(\boldsymbol{a} \cdot(\boldsymbol{b} \times \boldsymbol{d})) c_{r}-(\boldsymbol{a} \cdot(\boldsymbol{b} \times \boldsymbol{c})) d_{r}
\end{aligned}
$$

Finally

$$
(a \times b) \times(c \times d)=(a \cdot(b \times d)) c-(a \cdot(b \times c)) d
$$

## Solution 1.6

Identity (B1.230)
Expressing it in index notation one has

$$
\begin{aligned}
\boldsymbol{\nabla} \cdot(\boldsymbol{a} \times \boldsymbol{b}) & =\frac{\partial}{\partial x_{j}}\left(\varepsilon_{j k l} a_{k} b_{l}\right) \\
& =\varepsilon_{j k l} a_{k} \frac{\partial b_{l}}{\partial x_{j}}+\varepsilon_{j k l} b_{l} \frac{\partial a_{k}}{\partial x_{j}} \\
& =-\varepsilon_{k j l} a_{k} \frac{\partial b_{l}}{\partial x_{j}}+\varepsilon_{l j k} b_{l} \frac{\partial a_{k}}{\partial x_{j}} \\
& =-\boldsymbol{a} \cdot(\boldsymbol{\nabla} \times \boldsymbol{b})+(\boldsymbol{\nabla} \times \boldsymbol{a}) \cdot \boldsymbol{b} .
\end{aligned}
$$

Identity (B1.231)
By definition, the expression $(\boldsymbol{a} \cdot \boldsymbol{\nabla}) \boldsymbol{b}$ with index notation is written as

$$
((\boldsymbol{a} \cdot \boldsymbol{\nabla}) \boldsymbol{b})_{i}=a_{j} \frac{\partial b_{i}}{\partial x_{j}} .
$$

With the curl definition (B1.177), one has

$$
(\nabla \times \boldsymbol{b})_{i}=(\operatorname{curl} \boldsymbol{b})_{i}=\varepsilon_{i j k} \frac{\partial b_{k}}{\partial x_{j}} .
$$

As a consequence,

$$
(\boldsymbol{a} \times(\nabla \times \boldsymbol{b}))_{m}=\varepsilon_{m n i} \varepsilon_{i j k} a_{n} \frac{\partial b_{k}}{\partial x_{j}} .
$$

With (B1.30), one has

$$
\varepsilon_{m n i} \varepsilon_{i j k}=\varepsilon_{i m n} \varepsilon_{i j k}=\delta_{m j} \delta_{n k}-\delta_{m k} \delta_{n j} .
$$

The third term of the right hand side of ( $\mathbf{B} 1.231$ ) becomes

$$
\begin{aligned}
(\boldsymbol{a} \times(\nabla \times \boldsymbol{b}))_{m} & =\left(\delta_{m j} \delta_{n k}-\delta_{m k} \delta_{n j}\right) a_{n} \frac{\partial b_{k}}{\partial x_{j}} \\
& =\delta_{m j} \delta_{n k} a_{n} \frac{\partial b_{k}}{\partial x_{j}}-\delta_{m k} \delta_{n j} a_{n} \frac{\partial b_{k}}{\partial x_{j}} \\
& =a_{k} \frac{\partial b_{k}}{\partial x_{m}}-a_{j} \frac{\partial b_{m}}{\partial x_{j}} .
\end{aligned}
$$

Accordingly, one finds

$$
\begin{aligned}
a_{j} \frac{\partial b_{i}}{\partial x_{j}}+b_{j} \frac{\partial a_{i}}{\partial x_{j}}+a_{n} \frac{\partial b_{n}}{\partial x_{i}}-a_{j} \frac{\partial b_{i}}{\partial x_{j}}+b_{n} \frac{\partial a_{n}}{\partial x_{i}}-b_{j} \frac{\partial a_{i}}{\partial x_{j}} & =a_{n} \frac{\partial b_{n}}{\partial x_{i}}+b_{n} \frac{\partial a_{n}}{\partial x_{i}} \\
=\frac{\partial\left(a_{n} b_{n}\right)}{\partial x_{i}} & =(\boldsymbol{\nabla}(\boldsymbol{a} \cdot \boldsymbol{b}))_{i}
\end{aligned}
$$

## Identity (B1.232)

The right hand side of ( $\mathbf{B} 1.232$ ) is written as

$$
b_{j} \frac{\partial a_{i}}{\partial x_{j}}-a_{j} \frac{\partial b_{i}}{\partial x_{j}}+a_{i} \frac{\partial b_{j}}{\partial x_{j}}-b_{i} \frac{\partial a_{j}}{\partial x_{j}}=\frac{\partial\left(a_{i} b_{j}\right)}{\partial x_{j}}-\frac{\partial\left(a_{j} b_{i}\right)}{\partial x_{j}} .
$$

The left hand side gives

$$
\begin{equation*}
\varepsilon_{m n i} \frac{\partial\left(\varepsilon_{i j k} a_{j} b_{k}\right)}{\partial x_{n}}=\varepsilon_{m n i} \varepsilon_{i j k} \frac{\partial\left(a_{j} b_{k}\right)}{\partial x_{n}} . \tag{1.3}
\end{equation*}
$$

By (B1.30), one has

$$
\varepsilon_{m n i} \varepsilon_{i j k}=\varepsilon_{i m n} \varepsilon_{i j k}=\delta_{m j} \delta_{n k}-\delta_{m k} \delta_{n j} .
$$

Relation (1.3) becomes

$$
\left(\delta_{m j} \delta_{n k}-\delta_{m k} \delta_{n j}\right) \frac{\partial\left(a_{j} b_{k}\right)}{\partial x_{n}}=\frac{\partial\left(a_{m} b_{n}\right)}{\partial x_{n}}-\frac{\partial\left(a_{n} b_{m}\right)}{\partial x_{n}} .
$$

Identity (B1.233)
We successively obtain

$$
\begin{aligned}
\frac{\partial}{\partial x_{j}}\left(a_{i} b_{j}\right) & =a_{i} \frac{\partial b_{j}}{\partial x_{j}}+b_{j} \frac{\partial a_{i}}{\partial x_{j}} \\
& =\boldsymbol{a} \operatorname{div} \boldsymbol{b}+(\boldsymbol{\nabla} \boldsymbol{a}) \boldsymbol{b} .
\end{aligned}
$$

## Solution 1.7

Identity (B1.234)
One writes successively

$$
\begin{aligned}
(\operatorname{curl}(\Phi \boldsymbol{a}))_{i} & =\varepsilon_{i j k} \frac{\partial\left(\Phi a_{k}\right)}{\partial x_{j}} \\
& =\varepsilon_{i j k} \frac{\partial \Phi}{\partial x_{j}} a_{k}+\varepsilon_{i j k} \Phi \frac{\partial a_{k}}{\partial x_{j}} \\
& =\nabla \Phi \times \boldsymbol{a}+\Phi \mathbf{c u r l} \boldsymbol{a} \\
& =-\boldsymbol{a} \times \nabla \Phi+\Phi \operatorname{curl} \boldsymbol{a} .
\end{aligned}
$$

Identity (B1.235)
One has

$$
\begin{aligned}
(\boldsymbol{\nabla}(\Phi \boldsymbol{a}))_{i} & =\frac{\partial \Phi a_{i}}{\partial x_{j}} \\
& =\Phi \frac{\partial a_{i}}{\partial x_{j}}+a_{i} \frac{\partial \Phi}{\partial x_{j}} \\
& =\Phi \boldsymbol{\nabla} \boldsymbol{a}+\boldsymbol{a} \otimes \boldsymbol{\nabla} \Phi .
\end{aligned}
$$

## Identity (B1.236)

The $j$ component of gradient of $\Phi$ yields

$$
(\nabla \Phi)_{j}=\frac{\partial \Phi}{\partial x_{j}} .
$$

Therefore

$$
\left(\left(\nabla^{2}(\nabla \Phi)\right)_{j}=\frac{\partial^{2}}{\partial x_{i} \partial x_{i}}\left(\frac{\partial \Phi}{\partial x_{j}}\right)=\frac{\partial}{\partial x_{j}}\left(\frac{\partial^{2} \Phi}{\partial x_{i} \partial x_{i}}\right)=\left(\nabla\left(\nabla^{2} \Phi\right)\right)_{j} .\right.
$$

Thus

$$
\nabla^{2}(\nabla \Phi)=\boldsymbol{\nabla}\left(\nabla^{2} \Phi\right) .
$$

## Identity (B1.237)

The $l$ component of the vector corresponding to the left hand side is such that

$$
\begin{aligned}
\left(\boldsymbol{\nabla} \times\left(\nabla^{2} \boldsymbol{a}\right)\right)_{l} & =\left(\varepsilon_{i j k} \frac{\partial}{\partial x_{j}}\left(\frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{m}}\right)\right)_{l}=\left(\varepsilon_{i j k} \frac{\partial^{2}}{\partial x_{m} \partial x_{m}}\left(\frac{\partial a_{k}}{\partial x_{j}}\right)\right)_{l} \\
& =\left(\frac{\partial^{2}}{\partial x_{m} \partial x_{m}}\left(\varepsilon_{i j k} \frac{\partial a_{k}}{\partial x_{j}}\right)\right)_{l}=\left(\nabla^{2}(\boldsymbol{\nabla} \times \boldsymbol{a})\right)_{l} .
\end{aligned}
$$

## Identity (B1.238)

First indexed equality

$$
(\Delta \boldsymbol{a})_{k}=\frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{m}}=\frac{\partial}{\partial x_{m}}\left(\frac{\partial a_{k}}{\partial x_{m}}\right)=(\boldsymbol{\nabla} \cdot(\boldsymbol{\nabla} \boldsymbol{a}))_{k} .
$$

Second equality

$$
\begin{aligned}
(\boldsymbol{\nabla} \times \boldsymbol{a})_{i} & =\varepsilon_{i j k} \frac{\partial a_{k}}{\partial x_{j}} \\
(\boldsymbol{\nabla} \times(\boldsymbol{\nabla} \times \boldsymbol{a}))_{l} & =\varepsilon_{l m i} \frac{\partial}{\partial x_{m}}\left(\varepsilon_{i j k} \frac{\partial a_{k}}{\partial x_{j}}\right)=\varepsilon_{l m i} \varepsilon_{i j k} \frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{j}} .
\end{aligned}
$$

Note that by (B1.30), one has

$$
\varepsilon_{l m i} \varepsilon_{i j k}=\varepsilon_{i l m} \varepsilon_{i j k}=\delta_{l j} \delta_{m k}-\delta_{l k} \delta_{m j} .
$$

Therefore, one obtains

$$
\begin{aligned}
\varepsilon_{l m i} \varepsilon_{i j k} \frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{j}} & =\left(\delta_{l j} \delta_{m k}-\delta_{l k} \delta_{m j}\right) \frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{j}} \\
& =\delta_{l j} \delta_{m k} \frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{j}}-\delta_{l k} \delta_{m j} \frac{\partial^{2} a_{k}}{\partial x_{m} \partial x_{j}}
\end{aligned}
$$

and then

$$
\begin{aligned}
(\nabla \times(\nabla \times \boldsymbol{a}))_{l} & =\frac{\partial^{2} a_{m}}{\partial x_{m} \partial x_{l}}-\frac{\partial^{2} a_{l}}{\partial x_{j} \partial x_{j}} \\
& =\frac{\partial}{\partial x_{l}}\left(\frac{\partial a_{m}}{\partial x_{m}}\right)-\frac{\partial^{2} a_{l}}{\partial x_{j} \partial x_{j}}=\frac{\partial}{\partial x_{l}}(\boldsymbol{\nabla} \cdot \boldsymbol{a})-\Delta a_{l} \\
& =\left(\boldsymbol{\nabla}(\boldsymbol{\nabla} \cdot \boldsymbol{a})-\nabla^{2} \boldsymbol{a}\right)_{l}
\end{aligned}
$$

This last relation is valid for each component of the vector function $\boldsymbol{a}$

$$
\Delta a=\nabla(\nabla \cdot a)-\operatorname{curl} \operatorname{curl} a .
$$

## Solution 1.8

Identity (B1.239)
The $i$ component of the vector corresponding to the left hand side is such that

$$
\begin{aligned}
(\nabla(\boldsymbol{a} \cdot \boldsymbol{x}))_{i} & =\left(\frac{\partial a_{j}}{\partial x_{i}} x_{j}+a_{j} \delta_{i j}\right) \\
& =\left(\frac{\partial a_{j}}{\partial x_{i}} x_{j}+a_{i}\right)=\left(\boldsymbol{a}+(\nabla \boldsymbol{a})^{T} \boldsymbol{x}\right)_{i} .
\end{aligned}
$$

Thus

$$
\nabla(\boldsymbol{a} \cdot \boldsymbol{x})=\boldsymbol{a}+(\nabla \boldsymbol{a})^{T} \boldsymbol{x} .
$$

## Identity (B1.240)

The Laplacian is written

$$
\nabla^{2}(\boldsymbol{a} \cdot \boldsymbol{x})=\frac{\partial^{2}\left(a_{i} x_{i}\right)}{\partial x_{j} \partial x_{j}}=\frac{\partial}{\partial x_{j}}\left(\frac{\partial\left(a_{i} x_{i}\right)}{\partial x_{j}}\right) .
$$

Carrying out the algebra, one has

$$
\begin{aligned}
\frac{\partial}{\partial x_{j}}\left(\frac{\partial\left(a_{i} x_{i}\right)}{\partial x_{j}}\right) & =\frac{\partial}{\partial x_{j}}\left(\frac{\partial a_{i}}{\partial x_{j}} x_{i}+a_{i} \frac{\partial x_{i}}{\partial x_{j}}\right)=\frac{\partial}{\partial x_{j}}\left(\frac{\partial a_{i}}{\partial x_{j}} x_{i}+a_{i} \delta_{i j}\right) \\
& =\frac{\partial^{2} a_{i}}{\partial x_{j} \partial x_{j}} x_{i}+\frac{\partial a_{i}}{\partial x_{j}} \frac{\partial x_{i}}{\partial x_{j}}+\delta_{i j} \frac{\partial a_{i}}{\partial x_{j}} \\
& =\frac{\partial^{2} a_{i}}{\partial x_{j} \partial x_{j}} x_{i}+\frac{\partial a_{i}}{\partial x_{j}} \delta_{i j}+\delta_{i j} \frac{\partial a_{i}}{\partial x_{j}}=\frac{\partial^{2} a_{i}}{\partial x_{j} \partial x_{j}} x_{i}+2 \frac{\partial a_{i}}{\partial x_{j}} \delta_{i j} \\
& =\left(\nabla^{2} \boldsymbol{a}\right)_{i}+2(\operatorname{div} \boldsymbol{a})_{i} .
\end{aligned}
$$

Thus

$$
\nabla^{2}(\boldsymbol{a} \cdot \boldsymbol{x})=2 \operatorname{div} \boldsymbol{a}+\boldsymbol{x} \cdot\left(\nabla^{2} \boldsymbol{a}\right) .
$$

Identity (B1.241)
The Laplacian is written

$$
\begin{aligned}
\left(\nabla^{2}(\Phi \boldsymbol{x})\right)_{i} & =\frac{\partial^{2}\left(\Phi x_{i}\right)}{\partial x_{j} \partial x_{j}}=\frac{\partial}{\partial x_{j}}\left(\frac{\partial\left(\Phi x_{i}\right)}{\partial x_{j}}\right)=\frac{\partial}{\partial x_{j}}\left(\frac{\partial \Phi}{\partial x_{j}} x_{i}+\Phi \frac{\partial x_{i}}{\partial x_{j}}\right) \\
& =\frac{\partial^{2} \Phi}{\partial x_{j} \partial x_{j}} x_{i}+\frac{\partial \Phi}{\partial x_{j}} \delta_{i j}+\frac{\partial \Phi}{\partial x_{j}} \delta_{i j}=\frac{\partial^{2} \Phi}{\partial x_{j} \partial x_{j}} x_{i}+2 \frac{\partial \Phi}{\partial x_{i}}
\end{aligned}
$$

Thus

$$
\begin{aligned}
\left(\nabla^{2}(\Phi \boldsymbol{x})\right)_{i} & =\left(\boldsymbol{x} \nabla^{2} \Phi\right)_{i}+2(\nabla \phi)_{i} \\
\nabla^{2}(\Phi \boldsymbol{x}) & =2 \nabla \Phi+\boldsymbol{x} \nabla^{2} \Phi .
\end{aligned}
$$

## Solution 1.9

One writes with (B1.52)

$$
\begin{aligned}
L_{i j}^{\prime} & =\boldsymbol{e}_{i}^{\prime} \cdot \boldsymbol{L} \boldsymbol{e}_{j}^{\prime}= \\
& =\left(c_{i k} \boldsymbol{e}_{k}\right) \cdot \boldsymbol{L}\left(c_{j l} \boldsymbol{e}_{l}\right)= \\
& =c_{i k} c_{j l} \boldsymbol{e}_{k} \cdot \boldsymbol{L} \boldsymbol{e}_{l}= \\
& =c_{i k} c_{j l} L_{k l} .
\end{aligned}
$$

Moreover, by (B1.15)

$$
c_{i k} c_{i l}=\delta_{k l}
$$

One has

$$
\begin{aligned}
\operatorname{tr}\left(\boldsymbol{L}^{\prime}\right) & =L_{i i}^{\prime} \\
& =c_{i k} c_{i l} L_{k l} \\
& =\delta_{k l} L_{k l} \\
& =L_{k k} \\
& =\operatorname{tr}(\boldsymbol{L}) .
\end{aligned}
$$

## Solution (1.10)

Identity (B1.69)
One has successively

$$
\begin{aligned}
\boldsymbol{u} \cdot \boldsymbol{L}^{T} \boldsymbol{v} & =u_{i}\left(\boldsymbol{L}^{T}\right)_{i m} v_{m}=u_{i} L_{m i} v_{m}=L_{m i} u_{i} v_{m}=(\boldsymbol{L} \boldsymbol{u} \cdot \boldsymbol{v}) \\
\boldsymbol{u} \cdot \boldsymbol{L}^{T} \boldsymbol{v} & =u_{i}\left(\boldsymbol{L}^{T}\right)_{i m} v_{m}=u_{i} L_{m i} v_{m}=v_{m} L_{m i} u_{i}=(\boldsymbol{v} \cdot \boldsymbol{L} \boldsymbol{u}) .
\end{aligned}
$$

Identity (B1.71)
By definition of the scalar product of two tensors, one obtains

$$
((\boldsymbol{a} \otimes \boldsymbol{b}) \boldsymbol{L})_{i j}=(\boldsymbol{a} \otimes \boldsymbol{b})_{i k} L_{k j}=a_{i} b_{k} L_{k j}=a_{i}\left(\boldsymbol{L}^{T}\right)_{j k} b_{k}=\left(\boldsymbol{a} \otimes \boldsymbol{L}^{T} \boldsymbol{b}\right)_{i j} .
$$

## Solution 1.11

$$
\begin{aligned}
& I_{1}(\boldsymbol{L})=L_{11}+L_{22}+L_{33}=\operatorname{tr} \boldsymbol{L} \\
I_{2}(\boldsymbol{L}) & =L_{11} L_{22}-L_{21} L_{12}+L_{22} L_{33}-L_{23} L_{32}+L_{11} L_{33}-L_{13} L_{31}= \\
& =\left(L_{11} L_{22}+L_{22} L_{33}+L_{11} L_{33}\right)-\left(L_{21} L_{12}+L_{23} L_{32}+L_{13} L_{31}\right)= \\
& \left.=\frac{1}{2}\left(L_{11}+L_{22}+L_{33}\right)^{2}-\frac{1}{2}\left(L_{11}^{2}+L_{22}^{2}+L_{33}^{2}\right)\right] \\
& -\left(L_{21} L_{12}+L_{23} L_{32}+L_{13} L_{31}\right)= \\
& =\frac{1}{2}\left(L_{11}+L_{22}+L_{33}\right)^{2} \\
& -\frac{1}{2}\left(L_{11}^{2}+L_{22}^{2}+L_{33}^{2}+2 L_{21} L_{12}+2 L_{23} L_{32}+2 L_{13} L_{31}\right)= \\
& =\frac{1}{2}(\operatorname{tr} \boldsymbol{L})^{2}-\frac{1}{2}(\operatorname{tr}(\boldsymbol{L} \boldsymbol{L})) \\
& =\frac{1}{2}\left((\operatorname{tr} \boldsymbol{L})^{2}-(\operatorname{tr}(\boldsymbol{L} \boldsymbol{L}))\right) \\
& =\frac{1}{2}\left(L_{i i} L_{j j}-L_{i j} L_{j i}\right) \\
& I_{3}(\boldsymbol{L})=\operatorname{det}\left(\begin{array}{ccc}
L_{11} & L_{12} & L_{13} \\
L_{21} & L_{22} & L_{23} \\
L_{31} & L_{32} & L_{33}
\end{array}\right)=\varepsilon_{i j k} L_{i 1} L_{j 2} L_{k 3}=\operatorname{det} \boldsymbol{L} .
\end{aligned}
$$

## Solution 1.12

With definition (B1.159), one finds

$$
\begin{aligned}
\boldsymbol{\nabla}\left(A_{j k} x_{j} x_{k}\right) & =A_{j k} \boldsymbol{\nabla}\left(x_{j} x_{k}\right) \\
& =A_{j k}\left(x_{k} \boldsymbol{\nabla} x_{j}+x_{j} \nabla x_{k}\right) \\
& =A_{j k}\left(x_{k} \delta_{i j}+x_{j} \delta_{i k}\right) \\
& =A_{i k} x_{k}+A_{j i} x_{j} \\
& =\left(A_{i j}+A_{j i}\right) x_{j} \boldsymbol{e}_{i}
\end{aligned}
$$

## Solution 1.13

The tensor $D_{i j}$ is decomposed into the sum of a symmetric tensor and of an antisymmetric tensor. One has

$$
D_{i j}=D_{i j}^{S}+D_{i j}^{A}
$$

with the relations

$$
D_{i j}^{S}=D_{j i}^{S}
$$

and

$$
D_{i j}^{A}=-D_{j i}^{A} .
$$

Henceforth, one calculates

$$
\begin{aligned}
D_{i j} x_{i} x_{j} & =\left(D_{i j}^{S}+D_{i j}^{A}\right) x_{i} x_{j} \\
& =D_{i j}^{S} x_{i} x_{j}+D_{i j}^{A} x_{i} x_{j} .
\end{aligned}
$$

One uses the fact that the scalar product of the antisymmetric tensor $D_{i j}^{A}$ and of the symmetric tensor $x_{i} x_{j}$ vanishes (cfr. example $\mathbf{B} 1.7$ ) to obtain

$$
D_{i j} x_{i} x_{j}=D_{i j}^{S} x_{i} x_{j} .
$$

## Solution 1.14

As $\boldsymbol{Q}$ is orthogonal, relation (B1.92) gives

$$
Q u \cdot Q v=u \cdot v
$$

By (B1.69), one has

$$
\boldsymbol{L} \boldsymbol{u} \cdot \boldsymbol{v}=\boldsymbol{u} \cdot \boldsymbol{L}^{T} \boldsymbol{v}=\boldsymbol{v} \cdot \boldsymbol{L} \boldsymbol{u}
$$

Combining these two last relations, one finds

$$
\boldsymbol{Q u} \cdot \boldsymbol{Q} \boldsymbol{v}=\boldsymbol{u} \cdot \boldsymbol{Q}^{T} \boldsymbol{Q} \boldsymbol{v}=\boldsymbol{u} \cdot \boldsymbol{v}
$$

and therefore

$$
\boldsymbol{Q}^{T} \boldsymbol{Q}=\boldsymbol{I}
$$

Then

$$
\boldsymbol{Q}^{-T} \boldsymbol{Q}^{T} \boldsymbol{Q} \boldsymbol{Q}^{T}=\boldsymbol{Q}^{-T} \boldsymbol{I} \boldsymbol{Q}^{T}=\boldsymbol{Q}^{-T} \boldsymbol{Q}^{T}
$$

one concludes

$$
\boldsymbol{Q} \boldsymbol{Q}^{T}=\boldsymbol{I} .
$$

Multiplying left and right this equation by $\boldsymbol{Q}^{-1}$, one writes

$$
\boldsymbol{Q}^{T} \boldsymbol{Q} \boldsymbol{Q}^{-1}=\boldsymbol{I} \boldsymbol{Q}^{-1}
$$

and consequently

$$
Q^{T}=Q^{-1} .
$$

## Solution 1.15

The antisymmetric tensor $\boldsymbol{L}^{A}$ has the following matrix

$$
\left[L^{A}\right]=\left(\begin{array}{ccc}
0 & -\omega_{3} & \omega_{2} \\
\omega_{3} & 0 & -\omega_{1} \\
-\omega_{2} & \omega_{1} & 0
\end{array}\right)
$$

The first invariant is the sum of the diagonal elements (cf. (B1.121))

$$
I_{1}=0 .
$$

For $I_{2}$ and $I_{3}$, one obtains easily

$$
I_{2}=\omega_{1}^{2}+\omega_{2}^{2}+\omega_{3}^{2}
$$

and

$$
I_{3}=0 .
$$

The characteristic equation (B1.120) gives

$$
\lambda^{3}+\left(\omega_{1}^{2}+\omega_{2}^{2}+\omega_{3}^{2}\right) \lambda=0
$$

or

$$
\left(\lambda^{2}+\left(\omega_{1}^{2}+\omega_{2}^{2}+\omega_{3}^{2}\right)\right) \lambda=0 .
$$

The eigenvalues are the roots of this equation. One has

$$
\lambda_{1}=0
$$

and

$$
\lambda_{2,3}= \pm i \sqrt{\omega_{1}^{2}+\omega_{2}^{2}+\omega_{3}^{2}} .
$$

The system ( $\mathbf{B} 1.111$ ) produces the equations

$$
\begin{array}{r}
0-\omega_{3} n_{2}+\omega_{2} n_{3}=0 \\
\omega_{3} n_{1}+0-\omega_{1} n_{3}=0 .
\end{array}
$$

One extracts

$$
\begin{aligned}
& \frac{n_{2}}{n_{3}}=\frac{\omega_{2}}{\omega_{3}} \\
& \frac{n_{1}}{n_{3}}=\frac{\omega_{1}}{\omega_{3}} .
\end{aligned}
$$

Since the vector $\boldsymbol{n}$ is a unit vector

$$
n_{1}^{2}+n_{2}^{2}+n_{3}^{2}=1
$$

one finds

$$
n_{3}^{2}\left(\left(\frac{\omega_{1}}{\omega_{3}}\right)^{2}+\left(\frac{\omega_{2}}{\omega_{3}}\right)^{2}+1\right)=1
$$

Finally, one obtains

$$
\begin{aligned}
& n_{1}=\omega_{1} \\
& n_{2}=\omega_{2} \\
& n_{3}=\omega_{3}
\end{aligned}
$$

## Solution 1.16

By (B1.109) and (B1.59), one has successively

$$
\boldsymbol{L} \boldsymbol{L} \boldsymbol{n}_{i}=\lambda_{i} \boldsymbol{L} \boldsymbol{n}_{i}
$$

$$
\begin{gathered}
\boldsymbol{L}^{2} \boldsymbol{n}_{i}=\lambda_{i} \lambda_{i} \boldsymbol{n}_{i}=\lambda_{i}^{2} \boldsymbol{n}_{i} \\
\boldsymbol{L}^{3} \boldsymbol{n}_{i}=\lambda_{i}^{2} \boldsymbol{L} \boldsymbol{n}_{i}=\lambda_{i}^{3} \boldsymbol{n}_{i}
\end{gathered}
$$

Each term of the characteristic equation can be rewritten using the previous relations

$$
\begin{array}{r}
\lambda_{i}^{3} \boldsymbol{n}_{i}=\boldsymbol{L}^{3} \boldsymbol{n}_{i} \\
-I_{1} \lambda_{i}^{2} \boldsymbol{n}_{i}=-I_{1} \boldsymbol{L}^{2} \boldsymbol{n}_{i} \\
I_{2} \lambda_{i} \boldsymbol{n}_{i}=I_{2} \boldsymbol{L} \boldsymbol{n}_{i} \\
-I_{3} \boldsymbol{n}_{i}=-I_{3} \boldsymbol{I} \boldsymbol{n}_{i}
\end{array}
$$

By adding one finds consecutively

$$
\begin{gathered}
\lambda_{i}^{3} \boldsymbol{n}_{i}-I_{1} \lambda_{i}^{2} \boldsymbol{n}_{i}+I_{2} \lambda_{i} \boldsymbol{n}_{i}-I_{3} \boldsymbol{n}_{i}=\boldsymbol{L}^{3} \boldsymbol{n}_{i}-I_{1} \boldsymbol{L}^{2} \boldsymbol{n}_{i}+I_{2} \boldsymbol{L} \boldsymbol{n}_{i}-I_{3} \boldsymbol{I} \boldsymbol{n}_{i} \\
\left(\lambda_{i}^{3}-I_{1} \lambda_{i}^{2}+I_{2} \lambda_{i}-I_{3}\right) \boldsymbol{n}_{i}=\left(\boldsymbol{L}^{3}-I_{1} \boldsymbol{L}^{2}+I_{2} \boldsymbol{L}-I_{3} \boldsymbol{I}\right) \boldsymbol{n}_{i}
\end{gathered}
$$

Thus one has

$$
\boldsymbol{L}^{3}-I_{1} \boldsymbol{L}^{2}+I_{2} \boldsymbol{L}-I_{3} \boldsymbol{I}=0
$$

## Solution 1.17

Multiplying (B1.123) by $\boldsymbol{T}^{-1}$, one obtains

$$
\boldsymbol{T}^{3} \boldsymbol{T}^{-1}-I_{1} \boldsymbol{T}^{2} \boldsymbol{T}^{-1}+I_{2} \boldsymbol{T} \boldsymbol{T}^{-1}-I_{3} \boldsymbol{I} \boldsymbol{T}^{-1}=\boldsymbol{T}^{2}-I_{1} \boldsymbol{T}+I_{2} \boldsymbol{I}-I_{3} \boldsymbol{T}^{-1}=\mathbf{0}
$$

or

$$
\boldsymbol{T}^{2}=I_{1} \boldsymbol{T}-I_{2} \boldsymbol{I}+I_{3} \boldsymbol{T}^{-1} .
$$

By (B1.139) and (B1.140), one has

$$
\begin{aligned}
\boldsymbol{L}=\boldsymbol{f}(\boldsymbol{T}) & =\varphi_{0} \boldsymbol{I}+\varphi_{1} \boldsymbol{T}+\varphi_{2}\left(I_{1} \boldsymbol{T}-I_{2} \boldsymbol{I}+I_{3} \boldsymbol{T}^{-1}\right)= \\
& =\left(\varphi_{0}-\varphi_{2} I_{2}\right) \boldsymbol{I}+\left(\varphi_{1}+\varphi_{2} I_{1}\right) \boldsymbol{T}+\varphi_{2} I_{3} \boldsymbol{T}^{-1}
\end{aligned}
$$

Setting up

$$
\begin{aligned}
& \alpha_{0}=\varphi_{0}-\varphi_{2} I_{2} \\
& \alpha_{1}=\varphi_{1}+\varphi_{2} I_{1} \\
& \alpha_{2}=\varphi_{2} I_{3}
\end{aligned}
$$

one obtains relation (B1.245).

## Solution 1.18

Let the matrix $[A]$ be of the order 3 such that

$$
[A]=\left(\begin{array}{lll}
a_{11} & a_{12} & a_{13}  \tag{1.4}\\
a_{21} & a_{22} & a_{23} \\
a_{31} & a_{32} & a_{33}
\end{array}\right)
$$

Its determinant is the scalar given by the relation

$$
\begin{aligned}
\operatorname{det}[A] & =a_{11} a_{22} a_{33}+a_{12} a_{23} a_{31}+a_{13} a_{21} a_{32} \\
& -a_{13} a_{22} a_{31}-a_{11} a_{23} a_{32}-a_{12} a_{21} a_{33}
\end{aligned}
$$

1) The product $\varepsilon_{i j k} \varepsilon_{l m n} a_{i l} a_{j m} a_{k n}$ generates 36 non vanishing terms (among the 81 possible) that can be grouped in six independent components as the one above. Thus the equation

$$
\operatorname{det}[A]=\frac{1}{6} \varepsilon_{i j k} \varepsilon_{l m n} a_{i l} a_{j m} a_{k n}
$$

is the determinant of $[A]$.
2) By definition, the inverse matrix of $[A]$ is

$$
\begin{equation*}
[A]^{-1}=\frac{[M]^{T}}{\operatorname{det}[A]} \tag{1.5}
\end{equation*}
$$

where $[M]^{T}$ is the transpose of the cofactor matrix of $[A]$ with $\operatorname{det}[A] \neq 0$. For matrix $[A]$ given by (1.4) the cofactor matrix has
the following elements

$$
\begin{array}{ll}
M_{11}=(-1)^{1+1}\left|\begin{array}{ll}
a_{22} & a_{23} \\
a_{32} & a_{33}
\end{array}\right| & M_{12}=(-1)^{1+2}\left|\begin{array}{ll}
a_{21} & a_{23} \\
a_{31} & a_{33}
\end{array}\right| \\
M_{13}=(-1)^{1+3}\left|\begin{array}{ll}
a_{21} & a_{22} \\
a_{31} & a_{32}
\end{array}\right| & M_{21}=(-1)^{2+1}\left|\begin{array}{ll}
a_{12} & a_{13} \\
a_{32} & a_{33}
\end{array}\right| \\
M_{22}=(-1)^{2+2}\left|\begin{array}{ll}
a_{11} & a_{13} \\
a_{31} & a_{33}
\end{array}\right| & M_{23}=(-1)^{2+3}\left|\begin{array}{ll}
a_{11} & a_{12} \\
a_{31} & a_{32}
\end{array}\right| \\
M_{31}=(-1)^{3+1}\left|\begin{array}{ll}
a_{12} & a_{13} \\
a_{22} & a_{23}
\end{array}\right| & M_{32}=(-1)^{3+2}\left|\begin{array}{ll}
a_{11} & a_{13} \\
a_{21} & a_{23}
\end{array}\right| \\
M_{33}=(-1)^{3+3}\left|\begin{array}{ll}
a_{11} & a_{12} \\
a_{21} & a_{22}
\end{array}\right| .
\end{array}
$$

Thus one can write

$$
\begin{aligned}
& {[M]=} \\
& {\left[\begin{array}{ccc}
\left(a_{22} a_{33}-a_{23} a_{32}\right) & -\left(a_{21} a_{33}-a_{23} a_{31}\right) & \left(a_{21} a_{32}-a_{22} a_{31}\right) \\
-\left(a_{12} a_{33}-a_{13} a_{32}\right) & \left(a_{11} a_{33}-a_{13} a_{31}\right) & -\left(a_{11} a_{32}-a_{12} a_{31}\right) \\
\left(a_{12} a_{23}-a_{13} a_{22}\right) & -\left(a_{11} a_{23}-a_{13} a_{21}\right) & \left(a_{11} a_{22}-a_{12} a_{21}\right)
\end{array}\right]}
\end{aligned}
$$

These elements are expressed in index notation using the permutation symbol

$$
M_{i j}=\frac{1}{2} \varepsilon_{i k l} \varepsilon_{j m n} a_{k m} a_{l n}
$$

One can verify easily that expression $\varepsilon_{i k l} \varepsilon_{j m n}$ produces only four non zero terms or two terms that appear twice. This makes the factor one half in the last relation necessary. For example

$$
\begin{aligned}
M_{13} & =\left[\varepsilon_{123} \varepsilon_{312} a_{21} a_{32}+\varepsilon_{123} \varepsilon_{321} a_{22} a_{31}\right. \\
& \left.+\varepsilon_{132} \varepsilon_{312} a_{31} a_{22}+\varepsilon_{132} \varepsilon_{321} a_{32} a_{21}\right] / 2 \\
& =2\left(a_{21} a_{32}-a_{22} a_{31}\right) / 2
\end{aligned}
$$

The transpose of the cofactor matrix $[M]$ is called the adjoint of $[A]$. It is expressed as

$$
\begin{equation*}
\left([M]^{T}\right)_{i j}=\frac{1}{2} \varepsilon_{j k l} \varepsilon_{i m n} a_{k m} a_{l n} \tag{1.6}
\end{equation*}
$$

Using (1.6) in (1.5), one obtains in index form the elements of the inverse matrix

$$
\left([A]^{-1}\right)_{i j}=\frac{1}{2 \operatorname{det}[A]} \varepsilon_{j k l} \varepsilon_{i m n} a_{k m} a_{l n}
$$

## Solution 1.19

A series of algebraic manipulations leads to the result

$$
\begin{aligned}
\nabla^{2}(f g) & =\frac{\partial^{2}(f g)}{\partial x_{1}^{2}}+\frac{\partial^{2}(f g)}{\partial x_{2}^{2}}+\frac{\partial^{2}(f g)}{\partial x_{3}^{2}}= \\
& =\frac{\partial}{\partial x_{1}}\left(\frac{\partial(f g)}{\partial x_{1}}\right)+\frac{\partial}{\partial x_{2}}\left(\frac{\partial(f g)}{\partial x_{2}}\right)+\frac{\partial}{\partial x_{3}}\left(\frac{\partial(f g)}{\partial x_{3}}\right)= \\
& =\frac{\partial}{\partial x_{1}}\left(\frac{\partial f}{\partial x_{1}} g+f \frac{\partial g}{\partial x_{1}}\right)+\frac{\partial}{\partial x_{2}}\left(\frac{\partial f}{\partial x_{2}} g+f \frac{\partial g}{\partial x_{2}}\right) \\
& +\frac{\partial}{\partial x_{3}}\left(\frac{\partial f}{\partial x_{3}} g+f \frac{\partial g}{\partial x_{3}}\right)= \\
& =\frac{\partial^{2} f}{\partial x_{1}^{2}} g+\frac{\partial f}{\partial x_{1}} \frac{\partial g}{\partial x_{1}}+\frac{\partial f}{\partial x_{1}} \frac{\partial g}{\partial x_{1}}+\frac{\partial^{2} g}{\partial x_{1}^{2}} f+ \\
& +\frac{\partial^{2} f}{\partial x_{2}^{2}} g+\frac{\partial f}{\partial x_{2}} \frac{\partial g}{\partial x_{2}}+\frac{\partial f}{\partial x_{2}} \frac{\partial g}{\partial x_{2}}+\frac{\partial^{2} g}{\partial x_{2}^{2}} f+ \\
& +\frac{\partial^{2} f}{\partial x_{3}^{2}} g+\frac{\partial f}{\partial x_{3}} \frac{\partial g}{\partial x_{3}}+\frac{\partial f}{\partial x_{3}} \frac{\partial g}{\partial x_{3}}+\frac{\partial^{2} g}{\partial x_{3}^{2}} f= \\
& =\left(\frac{\partial^{2} f}{\partial x_{1}^{2}}+\frac{\partial^{2} f}{\partial x_{2}^{2}}+\frac{\partial^{2} f}{\partial x_{3}^{2}}\right) g+\left(\frac{\partial^{2} g}{\partial x_{1}^{2}}+\frac{\partial^{2} g}{\partial x_{2}^{2}}+\frac{\partial^{2} g}{\partial x_{3}^{2}}\right) f \\
& +2 \frac{\partial f}{\partial x_{1}} \frac{\partial g}{\partial x_{1}}+2 \frac{\partial f}{\partial x_{2}} \frac{\partial g}{\partial x_{2}}+2 \frac{\partial f}{\partial x_{3}} \frac{\partial g}{\partial x_{3}}= \\
& =f \nabla^{2} g+g \nabla^{2} f+2 \nabla f \cdot \nabla g
\end{aligned}
$$

## Chapitre 2

## Kinematics of continuous media

## Solution 2.1

With relation (B2.8), one writes

$$
\boldsymbol{U}(\boldsymbol{X}, t)=\boldsymbol{x}-\boldsymbol{X}=-\frac{1}{2} X_{1} \boldsymbol{e}_{1}
$$

and by (B2.9), one has

$$
\boldsymbol{u}(\boldsymbol{x}, t)=-x_{1} \boldsymbol{e}_{1}
$$

## Solution 2.2

By (B2.133), one has

$$
x_{1}=X_{1}+k X_{2} \quad x_{2}=X_{2} \quad x_{3}=X_{3} .
$$

Recalling the definition of the deformation gradient tensor (B2.67)

$$
F_{i j}=\frac{\partial x_{i}}{\partial X_{j}},
$$

one calculates successively

$$
\boldsymbol{F}=\left(\begin{array}{ccc}
1 & k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

the right Cauchy-Green deformation tensor

$$
\boldsymbol{C}=\boldsymbol{F}^{T} \boldsymbol{F}=\left(\begin{array}{ccc}
1 & 0 & 0 \\
k & 1 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{ccc}
1 & k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)=\left(\begin{array}{ccc}
1 & k & 0 \\
k & k^{2}+1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

the left Cauchy-Green deformation tensor

$$
\boldsymbol{c}=\boldsymbol{F} \boldsymbol{F}^{T}=\left(\begin{array}{ccc}
1 & k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{ccc}
1 & 0 & 0 \\
k & 1 & 0 \\
0 & 0 & 1
\end{array}\right)=\left(\begin{array}{ccc}
1+k^{2} & k & 0 \\
k & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

the Green-Lagrange deformation tensor

$$
\boldsymbol{E}=\frac{1}{2}(\boldsymbol{C}-\boldsymbol{I})=\frac{1}{2}\left(\begin{array}{ccc}
0 & k & 0 \\
k & k^{2} & 0 \\
0 & 0 & 0
\end{array}\right)
$$

and the Euler-Almansi deformation tensor

$$
\begin{gathered}
\boldsymbol{e}=\frac{1}{2}\left(\boldsymbol{I}-\boldsymbol{c}^{-1}\right) \\
\boldsymbol{c}^{-1}=\boldsymbol{F}^{-T} \boldsymbol{F}^{-1} \\
\boldsymbol{F}^{-1}=\left(\begin{array}{ccc}
1 & -k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
\boldsymbol{F}^{-T}=\left(\begin{array}{ccc}
1 & 0 & 0 \\
-k & 1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
\boldsymbol{c}^{-1}=\boldsymbol{F}^{-T} \boldsymbol{F}^{-1}=\left(\begin{array}{ccc}
1 & -k & 0 \\
-k & k^{2}+1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
\boldsymbol{e}=\frac{1}{2}\left(\begin{array}{ccc}
0 & k & 0 \\
k & -k^{2} & 0 \\
0 & 0 & 0
\end{array}\right) .
\end{gathered}
$$

We detail below the computation of $\boldsymbol{F}^{-1}$ by the adjoint and cofactors method

$$
\begin{gathered}
A_{i j}=(-1)^{i+j} M_{i j} \\
A_{11}=(-1)^{2}\left|\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right|=1 \\
A_{22}=(-1)^{4}\left|\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right|=1 \\
A_{33}=(-1)^{6}\left|\begin{array}{ll}
1 & k \\
0 & 1
\end{array}\right|=1
\end{gathered}
$$

$$
\begin{aligned}
& A_{12}=(-1)^{3}\left|\begin{array}{ll}
0 & 0 \\
0 & 1
\end{array}\right|=0 \\
& A_{13}=(-1)^{4}\left|\begin{array}{ll}
0 & 1 \\
0 & 0
\end{array}\right|=0 \\
& A_{21}=(-1)^{3}\left|\begin{array}{cc}
k & 0 \\
0 & 1
\end{array}\right|=-k \\
& A_{23}=(-1)^{5}\left|\begin{array}{ll}
1 & k \\
0 & 0
\end{array}\right|=0 \\
& A_{31}=(-1)^{4}\left|\begin{array}{cc}
k & 0 \\
1 & 0
\end{array}\right|=0 \\
& A_{32}=(-1)^{5}\left|\begin{array}{ll}
1 & 0 \\
0 & 0
\end{array}\right|=0 \\
& \boldsymbol{A}=\left(\begin{array}{ccc}
1 & 0 & 0 \\
-k & 1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
& \boldsymbol{A}^{T}=\left(\begin{array}{ccc}
1 & -k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
& |\boldsymbol{M}|=\left|\begin{array}{lll}
1 & k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right|=1 \\
& \boldsymbol{F}^{-1}=\frac{\boldsymbol{A}^{T}}{|\boldsymbol{M}|}=\left(\begin{array}{ccc}
1 & -k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
& \boldsymbol{F} \cdot \boldsymbol{F}^{-1}=\left(\begin{array}{ccc}
1 & k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{ccc}
1 & -k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)=\left(\begin{array}{ccc}
1 & -k+k & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)=\boldsymbol{I}
\end{aligned}
$$

Solution 2.3

$$
\begin{gathered}
\boldsymbol{M}=\left(\begin{array}{ccc}
m & 0 & 0 \\
0 & m & 0 \\
0 & 0 & m
\end{array}\right) \\
\boldsymbol{M}^{-1}=\left(\begin{array}{ccc}
m^{-1} & 0 & 0 \\
0 & m^{-1} & 0 \\
0 & 0 & m^{-1}
\end{array}\right)
\end{gathered}
$$

As

$$
F_{i j}=\frac{\partial x_{i}}{\partial X_{j}}
$$

one obtains with (B2.130)

$$
\boldsymbol{F}=\left(\begin{array}{ccc}
m & 0 & 0 \\
0 & m & 0 \\
0 & 0 & m
\end{array}\right)
$$

and the right Cauchy-Green deformation tensor as

$$
\boldsymbol{C}=\boldsymbol{F}^{T} \boldsymbol{F}=\left(\begin{array}{ccc}
m & 0 & 0 \\
0 & m & 0 \\
0 & 0 & m
\end{array}\right)\left(\begin{array}{ccc}
m & 0 & 0 \\
0 & m & 0 \\
0 & 0 & m
\end{array}\right)=\left(\begin{array}{ccc}
m^{2} & 0 & 0 \\
0 & m^{2} & 0 \\
0 & 0 & m^{2}
\end{array}\right)
$$

Let us remark that $\boldsymbol{C}=\boldsymbol{c}$ as tensors $\boldsymbol{F}$ and $\boldsymbol{F}^{T}$ are diagonal. The following tensors can be easily expressed :
the Green-Lagrange deformation tensor

$$
\boldsymbol{E}=\frac{1}{2}(\boldsymbol{C}-\boldsymbol{I})=\frac{1}{2}\left(\begin{array}{ccc}
m^{2}-1 & 0 & 0 \\
0 & m^{2}-1 & 0 \\
0 & 0 & m^{2}-1
\end{array}\right)
$$

the inverse of the left Cauchy-Green tensor

$$
\boldsymbol{c}^{-1}=\boldsymbol{F}^{-T} \boldsymbol{F}^{-1}=\left(\begin{array}{ccc}
m^{-2} & 0 & 0 \\
0 & m^{-2} & 0 \\
0 & 0 & m^{-2}
\end{array}\right)
$$

the Euler-Almansi deformation tensor

$$
\boldsymbol{e}=\frac{1}{2}\left(\boldsymbol{I}-\boldsymbol{C}^{-1}\right)=\frac{1}{2}\left(\begin{array}{ccc}
1-m^{-2} & 0 & 0 \\
0 & 1-m^{-2} & 0 \\
0 & 0 & 1-m^{-2}
\end{array}\right)
$$

## Solution 2.4

1) The trajectory is the spatial curve describing the successive positions $\boldsymbol{x}$ of a particle $\boldsymbol{X}$ with respect to time $t$. One eliminates the variable $t$ in the relation $\boldsymbol{x}=\boldsymbol{\chi}(\boldsymbol{X}, t)$ in such a way as to obtain a system of implicit equations linking the positions $x_{i}$. Thus

$$
\begin{aligned}
\left(\frac{x_{1}-X_{1}}{a}\right)^{2} & +\left(\frac{x_{2}-X_{2}}{b}\right)^{2}=\cos ^{2} 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) \\
& +\sin ^{2} 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)=1 \\
x_{3} & =X_{3}
\end{aligned}
$$

The trajectory of a particle with given material coordinates $X_{1}, X_{2}, X_{3}$ is located on the plane $x_{3}=X_{3}$. In this plane, the trajectory is an ellipse whose centre is at point with coordinates $X_{1}, X_{2}, X_{3}$ and whose principal axes are oriented in the directions $\boldsymbol{e}_{1}$ and $\boldsymbol{e}_{2}$ and have $2 a$ and $2 b$ as respective lengths.
2) As the motion is given in the Lagrangian description, one finds the velocity and acceleration components by taking the partial derivative with respect to time with the $X_{j}$ held fixed. One finds

$$
\begin{aligned}
& V_{1}=\left.\frac{\partial x_{1}}{\partial t}\right|_{X_{j}}=-\frac{2 \pi a}{T} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) \\
& V_{2}=\left.\frac{\partial x_{2}}{\partial t}\right|_{X_{j}}=\frac{2 \pi b}{T} \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) \\
& V_{3}=\left.\frac{\partial x_{3}}{\partial t}\right|_{X_{j}}=0 \\
& A_{1}=\left.\frac{\partial V_{1}}{\partial t}\right|_{X_{j}}=-\frac{4 \pi^{2} a}{T^{2}} \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) \\
& A_{2}=\left.\frac{\partial V_{2}}{\partial t}\right|_{X_{j}}=-\frac{4 \pi^{2} b}{T^{2}} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) \\
& A_{3}=\left.\frac{\partial V_{3}}{\partial t}\right|_{X_{j}}=0 .
\end{aligned}
$$

3) The deformation gradient tensor is obtained by the relation $F_{i j}=$ $\frac{\partial x_{i}}{\partial X_{j}}$. Thus one has

$$
[F]=\left(\begin{array}{ccc}
1+\frac{2 \pi a}{L} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) & 0 & 0 \\
-\frac{2 \pi \hbar}{L} \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

In Lagrangian representation, one writes $\frac{D \boldsymbol{F}}{D t}=\left.\frac{\partial}{\partial t} F_{i j}\right|_{X_{k}}$. The matrix $[\dot{F}]$ gives

$$
[\dot{F}]=\left(\begin{array}{ccc}
\frac{4 \pi^{2} a}{L T} \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) & 0 & 0 \\
\frac{4 \pi^{2} b}{L T} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) & 0 & 0 \\
0 & 0 & 0
\end{array}\right)
$$

4) Using relation (B2.179), $\dot{\boldsymbol{F}}=\boldsymbol{L F} \boldsymbol{F}$, one has $\boldsymbol{L}=\dot{\boldsymbol{F}} \boldsymbol{F}^{-1}$. As on the one hand, $\operatorname{det}(\boldsymbol{F})=1+\frac{2 \pi a}{L} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)$ and on the other hand, the adjoint of $[F]$ is written by setting $2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)=\arg$

$$
\operatorname{adj}[F]=\left(\begin{array}{ccc}
1 & 0 & 0 \\
\frac{2 \pi b}{L} \cos \arg & 1+\frac{2 \pi a}{L} \sin \arg & 0 \\
0 & 0 & 1+\frac{2 \pi a}{L} \sin \arg
\end{array}\right)
$$

one obtains

$$
\left[F^{-1}\right]=\frac{1}{\operatorname{det}(\boldsymbol{F})} \operatorname{adj}[F]=\left(\begin{array}{ccc}
\frac{1}{1+\frac{2 \pi a}{L} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 & 0 \\
\frac{2 \pi b}{L} \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) \\
1+\frac{2 \pi a}{L} \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right) & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

One thus finds

$$
[L]=\left(\begin{array}{ccc}
\frac{4 \pi^{2} a \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 & 0 \\
\frac{4 \pi^{2} b \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 & 0 \\
0 & 0 & 0
\end{array}\right)
$$

5) the tensors $\boldsymbol{d}$ et $\dot{\boldsymbol{\omega}}$ are respectively the symmetric and antisymmetric parts of $\boldsymbol{L}$, cf. (B2.184). One finds

$$
[d]=\left(\begin{array}{ccc}
\frac{4 \pi^{2} a \cos 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & \frac{2 \pi^{2} b \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 \\
\frac{2 \pi^{2} b \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 & 0 \\
0 & 0 & 0
\end{array}\right)
$$

and

$$
[\dot{\boldsymbol{\omega}}]=\left(\begin{array}{ccc}
0 & \frac{-2 \pi^{2} b \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 \\
\frac{2 \pi^{2} b \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} & 0 & 0 \\
0 & 0 & 0
\end{array}\right) .
$$

The evaluation of the components of the vorticity vector is based on relation (B2.187). One calculates

$$
\begin{aligned}
\dot{\Omega}_{1} & =0 \\
\dot{\Omega}_{2} & =0 \\
\dot{\Omega}_{3}=\dot{\omega}_{21} & =\frac{2 \pi^{2} b \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)}{L T+2 \pi a T \sin 2 \pi\left(\frac{t}{T}-\frac{X_{1}}{L}\right)} .
\end{aligned}
$$

## Solution 2.5

1) With the definition of the deformation gradient tensor (B2.67), we write

$$
F_{i j}=\frac{\partial x_{i}}{\partial X_{j}}=\left(\begin{array}{ccc}
1 & a & 0 \\
a & 1 & 0 \\
0 & 0 & 1
\end{array}\right) .
$$

As tensor $\boldsymbol{F}$ is independent of $\boldsymbol{X}$, the deformation is homogeneous. As $\operatorname{det} \boldsymbol{F} \neq 1$, this deformation is not isochoric, as we will notice in
chapter 3, (cf. (B3.38)). For the transformation to be invertible, it is necessary to verify the inequalities (B2.69). This implies $-1<a<$ +1 .
2) One calculates successively :
the right Cauchy-Green deformation tensor

$$
\boldsymbol{C}=\boldsymbol{F}^{T} \boldsymbol{F}=\left(\begin{array}{ccc}
1+a^{2} & 2 a & 0 \\
2 a & 1+a^{2} & 0 \\
0 & 0 & 1
\end{array}\right)
$$

the Green-Lagrange deformation tensor

$$
\boldsymbol{E}=\frac{1}{2}(\boldsymbol{C}-\boldsymbol{I})=\left(\begin{array}{ccc}
\frac{a^{2}}{2} & a & 0 \\
a & \frac{a^{2}}{2} & 0 \\
0 & 0 & 0
\end{array}\right)
$$

the gradient of the displacement vector

$$
\boldsymbol{\nabla} \boldsymbol{u}=\boldsymbol{F}-\boldsymbol{I}=\left(\begin{array}{ccc}
0 & a & 0 \\
a & 0 & 0 \\
0 & 0 & 0
\end{array}\right)
$$

the infinitesimal deformation tensor

$$
\boldsymbol{\varepsilon}=\frac{1}{2}\left(\boldsymbol{\nabla} \boldsymbol{u}+\boldsymbol{\nabla} \boldsymbol{u}^{T}\right)=\left(\begin{array}{ccc}
0 & a & 0 \\
a & 0 & 0 \\
0 & 0 & 0
\end{array}\right) .
$$

With the hypothesis

$$
a \ll 1
$$

the tensors $\boldsymbol{C}, \boldsymbol{E}, \boldsymbol{\varepsilon}$ become

$$
\begin{gathered}
\boldsymbol{C}=\left(\begin{array}{ccc}
1+a^{2} & 2 a & 0 \\
2 a & 1+a^{2} & 0 \\
0 & 0 & 1
\end{array}\right) \simeq\left(\begin{array}{ccc}
1 & 2 a & 0 \\
2 a & 1 & 0 \\
0 & 0 & 1
\end{array}\right) \\
\boldsymbol{E}=\frac{1}{2}(\boldsymbol{C}-\boldsymbol{I})=\left(\begin{array}{lll}
0 & a & 0 \\
a & 0 & 0 \\
0 & 0 & 0
\end{array}\right)=\boldsymbol{\varepsilon} .
\end{gathered}
$$

3) The unit eigenvectors of $\boldsymbol{C}$ are given by the relation (B2.110). Let's express the vectors oriented in direction $x_{3}$ and the diagonals $A H$ and $D E$. One has

$$
\begin{aligned}
& \boldsymbol{A}_{1}=\alpha\left(\boldsymbol{x}_{1}+\boldsymbol{x}_{2}\right) \\
& \boldsymbol{A}_{2}=\beta\left(\boldsymbol{x}_{1}-\boldsymbol{x}_{2}\right) \\
& \boldsymbol{A}_{3}=\gamma \boldsymbol{x}_{3} .
\end{aligned}
$$

By (B2.120), one has

$$
\boldsymbol{C} \boldsymbol{A}_{i}=\lambda_{i}^{2} \boldsymbol{A}_{i} \quad \text { (no summation on } i \text { ) }
$$

and therefore, one evaluates successively

$$
\begin{aligned}
\boldsymbol{C A}_{1} & =\left(\begin{array}{ccc}
1+a^{2} & 2 a & 0 \\
2 a & 1+a^{2} & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
\alpha \\
\alpha \\
0
\end{array}\right)=(1+a)^{2}\left(\begin{array}{l}
\alpha \\
\alpha \\
0
\end{array}\right) \\
& =(1+a)^{2} \boldsymbol{A}_{1} \\
\boldsymbol{C A}_{2} & =\left(\begin{array}{ccc}
1+a^{2} & 2 a & 0 \\
2 a & 1+a^{2} & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
\beta \\
-\beta \\
0
\end{array}\right)=(1-a)^{2}\left(\begin{array}{c}
\beta \\
-\beta \\
0
\end{array}\right) \\
& =(1-a)^{2} \boldsymbol{A}_{2} \\
\boldsymbol{C A}_{3} & =\left(\begin{array}{ccc}
1+a^{2} & 2 a & 0 \\
2 a & 1+a^{2} & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{l}
0 \\
0 \\
\gamma
\end{array}\right)=\left(\begin{array}{l}
0 \\
0 \\
\gamma
\end{array}\right)=\boldsymbol{A}_{3} .
\end{aligned}
$$

The eignevaleurs of $\boldsymbol{C}$ are

$$
\begin{aligned}
& \lambda_{1}^{2}=(1+a)^{2} \\
& \lambda_{2}^{2}=(1-a)^{2} \\
& \lambda_{3}^{2}=1
\end{aligned}
$$

By the spectral representation $\boldsymbol{C}$, one writes

$$
\boldsymbol{C}=\lambda_{i}^{2}\left(\boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}\right)=\lambda_{1}^{2}\left(\boldsymbol{A}_{1} \otimes \boldsymbol{A}_{1}\right)+\lambda_{2}^{2}\left(\boldsymbol{A}_{2} \otimes \boldsymbol{A}_{2}\right)+\lambda_{3}^{2}\left(\boldsymbol{A}_{3} \otimes \boldsymbol{A}_{3}\right) .
$$

One has

$$
\begin{aligned}
& \boldsymbol{A}_{1} \otimes \boldsymbol{A}_{1}=\left(\begin{array}{c}
\alpha \\
\alpha \\
0
\end{array}\right) \otimes\left(\begin{array}{c}
\alpha \\
\alpha \\
0
\end{array}\right)=\left(\begin{array}{ccc}
\alpha^{2} & \alpha^{2} & 0 \\
\alpha^{2} & \alpha^{2} & 0 \\
0 & 0 & 0
\end{array}\right)=\left(\begin{array}{ccc}
\frac{1}{2} & \frac{1}{2} & 0 \\
\frac{1}{2} & \frac{1}{2} & 0 \\
0 & 0 & 0
\end{array}\right) \\
& \boldsymbol{A}_{2} \otimes \boldsymbol{A}_{2}=\left(\begin{array}{c}
\beta \\
-\beta \\
0
\end{array}\right) \otimes\left(\begin{array}{c}
\beta \\
-\beta \\
0
\end{array}\right)=\left(\begin{array}{ccc}
\beta^{2} & -\beta^{2} & 0 \\
-\beta^{2} & \beta^{2} & 0 \\
0 & 0 & 0
\end{array}\right)=\left(\begin{array}{ccc}
\frac{1}{2} & -\frac{1}{2} & 0 \\
-\frac{1}{2} & \frac{1}{2} & 0 \\
0 & 0 & 0
\end{array}\right) \\
& \boldsymbol{A}_{3} \otimes \boldsymbol{A}_{3}=\left(\begin{array}{l}
0 \\
0 \\
\gamma
\end{array}\right) \otimes\left(\begin{array}{l}
0 \\
0 \\
\gamma
\end{array}\right)=\left(\begin{array}{ccc}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & \gamma^{2}
\end{array}\right)=\left(\begin{array}{lll}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & 1
\end{array}\right) .
\end{aligned}
$$

One verifies

$$
\begin{aligned}
{[C] } & =(1+a)^{2} \times\left(\begin{array}{ccc}
\alpha^{2} & \alpha^{2} & 0 \\
\alpha^{2} & \alpha^{2} & 0 \\
0 & 0 & 0
\end{array}\right)+(1-a)^{2} \times\left(\begin{array}{ccc}
\beta^{2} & -\beta^{2} & 0 \\
-\beta^{2} & \beta^{2} & 0 \\
0 & 0 & 0
\end{array}\right) \\
& +1 \times\left(\begin{array}{lll}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & \gamma^{2}
\end{array}\right)= \\
& =(1+a)^{2} \times\left(\begin{array}{ccc}
\frac{1}{2} & \frac{1}{2} & 0 \\
\frac{1}{2} & \frac{1}{2} & 0 \\
0 & 0 & 0
\end{array}\right)+(1-a)^{2} \times\left(\begin{array}{ccc}
\frac{1}{2} & -\frac{1}{2} & 0 \\
-\frac{1}{2} & \frac{1}{2} & 0 \\
0 & 0 & 0
\end{array}\right) \\
& +1 \times\left(\begin{array}{lll}
0 & 0 & 0 \\
0 & 0 & 0 \\
0 & 0 & 1
\end{array}\right)=\left(\begin{array}{ccc}
1+a^{2} & 2 a & 0 \\
2 a & 1+a^{2} & 0 \\
0 & 0 & 1
\end{array}\right) .
\end{aligned}
$$

By (B.2.109), one obtains easily

$$
\begin{align*}
& {[U]=\left(\begin{array}{ccc}
1 & a & 0 \\
a & 1 & 0 \\
0 & 0 & 1
\end{array}\right)=[F] .} \\
& \boldsymbol{R}=\boldsymbol{F} \boldsymbol{U}^{-1}=\boldsymbol{F} \boldsymbol{F}^{-1}=\boldsymbol{I} .
\end{align*}
$$

## Solution 2.6

For the right Cauchy-Green tensor, one carries out the algebra

$$
\begin{gathered}
\boldsymbol{C}=\boldsymbol{F}^{T} \boldsymbol{F} ; \boldsymbol{C}^{*}=\boldsymbol{F}^{* T} \boldsymbol{F}^{*} ; \boldsymbol{F}^{*}=\boldsymbol{Q} \boldsymbol{F} ; \boldsymbol{F}^{* T}=\boldsymbol{F}^{T} \boldsymbol{Q}^{T}=\boldsymbol{F}^{T} \boldsymbol{Q}^{-1} \\
\boldsymbol{C}^{*}=\boldsymbol{F}^{T} \boldsymbol{Q}^{T} \boldsymbol{Q} \boldsymbol{F}=\boldsymbol{F}^{T} \boldsymbol{I} \boldsymbol{F}=\boldsymbol{F}^{T} \boldsymbol{F}=\boldsymbol{C} .
\end{gathered}
$$

For the Green-Lagrange tensor, one finds

$$
\begin{aligned}
2 \boldsymbol{E} & =\boldsymbol{C}-\boldsymbol{I} ; \boldsymbol{C}^{*}=\boldsymbol{C} \\
2 \boldsymbol{E}^{*} & =2 \boldsymbol{E} \Rightarrow \boldsymbol{E}^{*}=\boldsymbol{E} .
\end{aligned}
$$

For the left Cauchy-Green tensor, one obtains

$$
\begin{gathered}
\boldsymbol{c}=\boldsymbol{F} \boldsymbol{F}^{T} ; \boldsymbol{c}^{*}=\boldsymbol{F}^{*} \boldsymbol{F}^{* T} ; \boldsymbol{F}^{*}=\boldsymbol{Q} \boldsymbol{F} ; \boldsymbol{F}^{* T}=\boldsymbol{F}^{T} \boldsymbol{Q}^{T}=\boldsymbol{F}^{T} \boldsymbol{Q}^{-1} \\
\boldsymbol{c}^{*}=\boldsymbol{F}^{*} \boldsymbol{F}^{* T}=\boldsymbol{Q} \boldsymbol{F} \boldsymbol{F}^{T} \boldsymbol{Q}^{T}=\boldsymbol{Q} \boldsymbol{c} \boldsymbol{Q}^{T} .
\end{gathered}
$$

As a consequence, $\boldsymbol{c}$ is spatially objective. For the Euler-Almansi tensor, one verifies

$$
\begin{gathered}
2 \boldsymbol{e}^{*}=\boldsymbol{I}-\boldsymbol{c}^{*-1} ; \boldsymbol{Q}^{-T}=\boldsymbol{Q} ; \boldsymbol{Q}^{-1}=\boldsymbol{Q}^{T} \\
2 \boldsymbol{e}^{*}=\boldsymbol{I}-\boldsymbol{c}^{*-1}=\boldsymbol{Q} \boldsymbol{I} \boldsymbol{Q}^{T}-\boldsymbol{Q}^{-T} \boldsymbol{c}^{-1} \boldsymbol{Q}^{-1}=2 \boldsymbol{Q} \boldsymbol{e} \boldsymbol{Q}^{T}
\end{gathered}
$$

## Solution 2.7

With (B2.205) and the orthogonality of $\boldsymbol{R}$, one has

$$
\boldsymbol{F}^{*}=\boldsymbol{Q} \boldsymbol{F} \quad \text { and } \quad \boldsymbol{R}^{T} \boldsymbol{R}=\boldsymbol{I}
$$

By the polar decomposition theorem (B1.132), one writes

$$
\begin{gathered}
\boldsymbol{F}=\boldsymbol{R} \boldsymbol{U}=\boldsymbol{V} \boldsymbol{R} ; \boldsymbol{F}^{*}=\boldsymbol{R}^{*} \boldsymbol{U}^{*}=\boldsymbol{V}^{*} \boldsymbol{R}^{*}=\boldsymbol{Q} \boldsymbol{F} \\
\boldsymbol{R}^{*} \boldsymbol{U}^{*}=\boldsymbol{Q} \boldsymbol{R} \boldsymbol{U} \\
\boldsymbol{R}^{*}=\boldsymbol{Q} \boldsymbol{R} \boldsymbol{U} \boldsymbol{U}^{*-1}
\end{gathered}
$$

This relation is trivially satisfied if we set $\boldsymbol{R}^{*}=\boldsymbol{Q R}$ and thus $\boldsymbol{U}^{*}=\boldsymbol{U}$. Similarly one has successively

$$
\boldsymbol{V}^{*} \boldsymbol{R}^{*}=\boldsymbol{Q} \boldsymbol{V} \boldsymbol{R}
$$

and thus,

$$
\boldsymbol{V}^{*}=\boldsymbol{Q} \boldsymbol{V} \boldsymbol{R} \boldsymbol{R}^{*-1}=\boldsymbol{Q} \boldsymbol{V}\left(\boldsymbol{R} \boldsymbol{R}^{T} \boldsymbol{Q}^{T}\right)=\boldsymbol{Q} \boldsymbol{V} \boldsymbol{Q}^{T}
$$

## Solution 2.8

With the help of $(\mathbf{B} 2.88),(\mathbf{B} 2.91),(\mathbf{B} 2.179)$ and $(\mathbf{B} 2.180)$, one has

$$
2 \boldsymbol{E}=\boldsymbol{C}-\boldsymbol{I} ; \boldsymbol{C}=\boldsymbol{F}^{T} \boldsymbol{F} ; \dot{\boldsymbol{F}}=\boldsymbol{L} \boldsymbol{F} ; 2 \boldsymbol{d}=\boldsymbol{L}+\boldsymbol{L}^{T} .
$$

Therefore, one finds successively

$$
\begin{gathered}
\dot{\boldsymbol{E}}=\frac{\dot{\boldsymbol{C}}}{2} \\
\dot{\boldsymbol{C}}=\boldsymbol{F}^{T} \dot{\boldsymbol{F}}+\dot{\boldsymbol{F}}^{T} \boldsymbol{F}=\boldsymbol{F}^{T} \boldsymbol{L} \boldsymbol{F}+\boldsymbol{F}^{T} \boldsymbol{L}^{T} \boldsymbol{F}=\boldsymbol{F}^{T}\left(\boldsymbol{L}+\boldsymbol{L}^{T}\right) \boldsymbol{F}=2 \boldsymbol{F}^{T} \boldsymbol{d} \boldsymbol{F} \\
\dot{\boldsymbol{E}}=\boldsymbol{F}^{T} \boldsymbol{d} \boldsymbol{F}
\end{gathered}
$$

## Solution 2.9

The motion described by the relations

$$
x_{1}=\lambda_{1} X_{1}, \quad x_{2}=\lambda_{2} X_{2}, \quad x_{3}=\lambda_{3} X_{3}
$$

leads to the deformation gradient tensor defined by (B2.67)

$$
F_{i j}=\frac{\partial x_{i}}{\partial X_{j}}
$$

as the diagonal tensor

$$
F_{i j}=\left(\begin{array}{ccc}
\lambda_{1} & 0 & 0 \\
0 & \lambda_{2} & 0 \\
0 & 0 & \lambda_{3}
\end{array}\right) .
$$

The matrix of the right Cauchy-Green deformation tensor is

$$
\begin{aligned}
{[C]=[F]^{T}[F] } & =\left(\begin{array}{ccc}
\lambda_{1} & 0 & 0 \\
0 & \lambda_{2} & 0 \\
0 & 0 & \lambda_{3}
\end{array}\right)\left(\begin{array}{ccc}
\lambda_{1} & 0 & 0 \\
0 & \lambda_{2} & 0 \\
0 & 0 & \lambda_{3}
\end{array}\right) \\
& =\left(\begin{array}{ccc}
\lambda_{1}^{2} & 0 & 0 \\
0 & \lambda_{2}^{2} & 0 \\
0 & 0 & \lambda_{3}^{2}
\end{array}\right)
\end{aligned}
$$

and the matrix of the right Green-Lagrange deformation tensor given by

$$
[E]=\frac{1}{2}([C]-[I])=\frac{1}{2}\left(\begin{array}{ccc}
\lambda_{1}^{2}-1 & 0 & 0 \\
0 & \lambda_{2}^{2}-1 & 0 \\
0 & 0 & \lambda_{3}^{2}-1
\end{array}\right) .
$$

By (B2.88)

$$
\boldsymbol{C}=\boldsymbol{U}^{2}
$$

one finds

$$
[U]=\left(\begin{array}{ccc}
\lambda_{1} & 0 & 0 \\
0 & \lambda_{2} & 0 \\
0 & 0 & \lambda_{3}
\end{array}\right)
$$

and thus,

$$
\left\{\begin{array}{l}
{[F]=[U]} \\
{[F]=[R][U]}
\end{array}\right.
$$

and

$$
[R]=[I] .
$$

## Solution 2.10

With (B2.91) and (B2.92), one has

$$
2 \boldsymbol{E}=\boldsymbol{F}^{T} \boldsymbol{F}-\boldsymbol{I} \quad \text { and } \quad 2 \boldsymbol{e}=\boldsymbol{I}-\boldsymbol{F}^{-T} \boldsymbol{F}^{-1} .
$$

The last relation is multiplied left by $\boldsymbol{F}^{-T} \boldsymbol{F}^{T}(=\boldsymbol{I})$ and right by $\boldsymbol{F} \boldsymbol{F}^{-1}$ to obtain successively

$$
\begin{aligned}
2 \boldsymbol{e} & =\boldsymbol{F}^{-T}\left(\boldsymbol{F}^{T}\left(\boldsymbol{I}-\boldsymbol{F}^{-T} \boldsymbol{F}^{-1}\right) \boldsymbol{F}\right) \boldsymbol{F}^{-1} \\
& =\boldsymbol{F}^{-T}\left(\boldsymbol{F}^{T} \boldsymbol{F}-\boldsymbol{F}^{T} \boldsymbol{F}^{-T} \boldsymbol{F}^{-1} \boldsymbol{F}\right) \boldsymbol{F}^{-1} \\
& =\boldsymbol{F}^{-T}\left(\boldsymbol{F}^{T} \boldsymbol{F}-\boldsymbol{I}\right) \boldsymbol{F}^{-1} \\
& =2 \boldsymbol{F}^{-T} \boldsymbol{E} \boldsymbol{F}^{-1} .
\end{aligned}
$$

By the polar decomposition theorem, one has

$$
\boldsymbol{F}^{T}=\boldsymbol{U}^{T} \boldsymbol{R}^{T}=\boldsymbol{U} \boldsymbol{R}^{T}
$$

By definition (B2.89), one gets

$$
\boldsymbol{c}=\boldsymbol{F} \boldsymbol{F}^{T}=\boldsymbol{R} \boldsymbol{U} \boldsymbol{U} \boldsymbol{R}^{T}=\boldsymbol{R} \boldsymbol{U}^{2} \boldsymbol{R}^{T}=\boldsymbol{R} \boldsymbol{C} \boldsymbol{R}^{T} .
$$

## Solution 2.11

We recall relation (B2.108)

$$
\boldsymbol{U} \boldsymbol{A}_{i}=\lambda_{i} \boldsymbol{A}_{i} .
$$

With (B2.112) one finds

$$
\boldsymbol{R} \boldsymbol{U} \boldsymbol{A}_{i}=\lambda_{i} \boldsymbol{R} \boldsymbol{A}_{i} ; \boldsymbol{F} \boldsymbol{A}_{i}=\lambda_{i} \boldsymbol{b}_{i} .
$$

With (B2.109)

$$
\boldsymbol{U}=\sum_{i=1}^{3} \lambda_{i} \boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}
$$

we find

$$
\boldsymbol{R} \boldsymbol{U}=\sum_{i=1}^{3} \lambda_{i}\left(\boldsymbol{R} \boldsymbol{A}_{i}\right) \otimes \boldsymbol{A}_{i} \quad \text { et } \quad \boldsymbol{F}=\sum_{i=1}^{3} \lambda_{i} \boldsymbol{b}_{i} \otimes \boldsymbol{A}_{i} .
$$

## Solution 2.12

Equation (B2.106) can be written as $\boldsymbol{F}^{T} \boldsymbol{n} d s=J \boldsymbol{N} d S$ or $F_{j i} n_{j} d s=$ $J N_{i} d S$.

Introduce (B2.147) in this relation to obtain

$$
\begin{equation*}
\left(\delta_{j i}+O(\varepsilon)\right) n_{j} d s=(1+O(\varepsilon)) N_{i} d S \tag{2.1}
\end{equation*}
$$

which yields

$$
\begin{equation*}
n_{i} d s \approx N_{i} d S \tag{2.2}
\end{equation*}
$$

2.13

Lengthy solution.
(a) Use (B2.77), (B2.70), (B2.88) and (B2.120) to obtain

$$
\begin{aligned}
F_{m i} F_{m j} & =\left(\delta_{m i}+\frac{\partial U_{m}}{\partial X_{i}}\right)\left(\delta_{m j}+\frac{\partial U_{m}}{\partial X_{j}}\right) \\
& =\delta_{m i} \delta_{m j}+\delta_{m i} \frac{\partial U_{m}}{\partial X_{j}}+\delta_{m j} \frac{\partial U_{m}}{\partial X_{i}}+\delta_{m i} \delta_{m j} \frac{\partial U_{m}}{\partial X_{j}} \frac{\partial U_{m}}{\partial X_{i}} \\
& =\delta_{i j}+\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)+\frac{\partial U_{i}}{\partial X_{j}} \frac{\partial U_{j}}{\partial X_{i}} \\
& =\delta_{i j}+\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)+O\left(\varepsilon^{2}\right)
\end{aligned}
$$

Then one evaluates

$$
\begin{aligned}
U_{i j} & =\left(F_{m i} F_{m j}\right)^{1 / 2}=\left(\delta_{i j}+\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)+\frac{\partial U_{i}}{\partial X_{j}} \frac{\partial U_{j}}{\partial X_{i}}\right)^{1 / 2} \\
& \approx \delta_{i j}+\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)+O\left(\varepsilon^{2}\right) \\
& \approx \delta_{i j}+\varepsilon_{i j}+O\left(\varepsilon^{2}\right) .
\end{aligned}
$$

In the last relation we used $(1+a)^{n} \approx 1+n a$ for $a \ll 1$. Therefore

$$
U \approx I+\varepsilon .
$$

(b) It is easy to show that

$$
U_{k j}^{-1}=\delta_{k j}-\frac{1}{2}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)+O\left(\varepsilon^{2}\right)
$$

is the inverse of $U_{i j}$, i.e. $U_{i k} U_{k j}^{-1} \approx \delta_{i j}$.
Then use (B2.73) as follows

$$
\begin{aligned}
R_{i j} & =F_{i k} U_{k j}^{-1}=\left(\delta_{i k}+\frac{\partial U_{i}}{\partial X_{k}}\right)\left(\delta_{k j}-\frac{1}{2}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)+O\left(\varepsilon^{2}\right)\right) \\
& =\delta_{i j}-\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)+O\left(\varepsilon^{2}\right)+\frac{\partial U_{i}}{\partial X_{j}} \\
& -\frac{1}{2} \frac{\partial U_{i}}{\partial X_{k}}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)+O\left(\varepsilon^{2}\right) \\
& =\delta_{i j}-\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}-2 \frac{\partial U_{i}}{\partial X_{j}}\right)+O\left(\varepsilon^{2}\right) \\
& -\frac{1}{2} \frac{\partial U_{i}}{\partial X_{k}}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)+O\left(\varepsilon^{2}\right) \\
& =\delta_{i j}+\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}-\frac{\partial U_{j}}{\partial X_{i}}\right)-\frac{1}{2} \frac{\partial U_{i}}{\partial X_{k}}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)+O\left(\varepsilon^{2}\right) \\
& =\delta_{i j}+\omega_{i j}+O\left(\varepsilon^{2}\right) .
\end{aligned}
$$

Therefore

$$
R \approx I+\omega
$$

Short version
(a) Use (B2.77), (B2.70), (B2.88) and (B2.120) to obtain

$$
\begin{aligned}
F_{m i} F_{m j} & =\left(\delta_{m i}+\frac{\partial U_{m}}{\partial X_{i}}\right)\left(\delta_{m j}+\frac{\partial U_{m}}{\partial X_{j}}\right) \\
& =\delta_{m i} \delta_{m j}+\delta_{m i} \frac{\partial U_{m}}{\partial X_{j}}+\delta_{m j} \frac{\partial U_{m}}{\partial X_{i}}+\delta_{m i} \delta_{m j} \frac{\partial U_{m}}{\partial X_{j}} \frac{\partial U_{m}}{\partial X_{i}} \\
& \approx \delta_{i j}+\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)
\end{aligned}
$$

Then we evaluate

$$
\begin{aligned}
U_{i j} & =\left(F_{m i} F_{m j}\right)^{1 / 2}=\left(\delta_{i j}+\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)\right)^{1 / 2} \\
& \approx \delta_{i j}+\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right) \\
& \approx \delta_{i j}+\varepsilon_{i j}
\end{aligned}
$$

In the last relation we used $(1+a)^{n} \approx 1+n a$ for $a \ll 1$. Therefore

$$
U \approx I+\varepsilon
$$

(b) It is easy to show that

$$
U_{k j}^{-1}=\delta_{k j}-\frac{1}{2}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)+O\left(\varepsilon^{2}\right)
$$

is the inverse of $U_{i j}$, i.e. $U_{i k} U_{k j}^{-1} \approx \delta_{i j}$.
Then use (B2.73) as follows

$$
\begin{aligned}
R_{i j} & =F_{i k} U_{k j}^{-1}=\left(\delta_{i k}+\frac{\partial U_{i}}{\partial X_{k}}\right)\left(\delta_{k j}-\frac{1}{2}\left(\frac{\partial U_{k}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{k}}\right)\right) \\
& \approx \delta_{i j}-\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}\right)+\frac{\partial U_{i}}{\partial X_{j}} \\
& \approx \delta_{i j}-\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}+\frac{\partial U_{j}}{\partial X_{i}}-2 \frac{\partial U_{i}}{\partial X_{j}}\right) \\
& =\delta_{i j}+\frac{1}{2}\left(\frac{\partial U_{i}}{\partial X_{j}}-\frac{\partial U_{j}}{\partial X_{i}}\right) \\
& =\delta_{i j}+\omega_{i j}
\end{aligned}
$$

to obtain

$$
\boldsymbol{R} \approx \boldsymbol{I}+\omega
$$

## Solution 2.14

Relation ( $\mathbf{B} 2.157$ ) is given by $\cos \gamma_{12} \cong 2 \varepsilon_{12}$, where $\gamma_{12}$ is the angle between two vectors initially (before deformation) orthogonal. The variables $\varepsilon_{i j}$ are the components of the infinitesimal deformation tensor. The components of vectors before and after the movement (or deformation) are given by the following relations with reference to figure 2.1


Figure 2.1 Modification of the angles between two vectors.

$$
\begin{align*}
& d \boldsymbol{X}:\left(d X_{1}, 0,0\right) \rightarrow d \boldsymbol{x}:\left(d x_{1}, d x_{2}, d x_{3}\right)  \tag{2.3}\\
& d \boldsymbol{Y}:\left(0, d Y_{2}, 0\right) \rightarrow d \boldsymbol{y}:\left(d y_{1}, d y_{2}, d y_{3}\right) . \tag{2.4}
\end{align*}
$$

According to the body motion, one has from (B2.8)

$$
\begin{equation*}
d x_{i}=d U_{i}+d X_{i} . \tag{2.5}
\end{equation*}
$$

We remplace $d U_{i}$ by the next expression (see top of p .90 of the book)

$$
d U_{i}=\varepsilon_{i j} d X_{j}+\omega_{i j} d X_{j}
$$

To simplify the algebra, we assume that the infinitesimal rotations vanish. Thus

$$
\omega_{i j}=0 \Rightarrow d U_{i}=\varepsilon_{i j} d X_{j}
$$

The relation (2.5) becomes, taking into account (2.3)

$$
d x_{i}=d X_{i}+\varepsilon_{i j} d X_{j}=\left(\delta_{i j}+\varepsilon_{i j}\right) d X_{j}=\left(\delta_{i 1}+\varepsilon_{i 1}\right) d X_{1}
$$

and then

$$
d x_{1}=\left(1+\varepsilon_{11}\right) d X_{1}, \quad d x_{2}=\varepsilon_{21} d X_{1}, \quad d x_{3}=\varepsilon_{31} d X_{1} .
$$

Similarly, we obtain for segment $d \boldsymbol{y}$

$$
d y_{i}=d Y_{i}+\varepsilon_{i j} d Y_{j}=\left(\delta_{i j}+\varepsilon_{i j}\right) d Y_{j}=\left(\delta_{i 2}+\varepsilon_{i 2}\right) d Y_{2}
$$

and then

$$
d y_{1}=\varepsilon_{12} d Y_{2}, \quad d y_{2}=\left(1+\varepsilon_{22}\right) d Y_{2}, \quad d y_{3}=\varepsilon_{32} d Y_{2} .
$$

By (B2.157), one writes

$$
\begin{aligned}
\cos \gamma_{12} & =\frac{d \boldsymbol{x} \cdot d \boldsymbol{y}}{\|d \boldsymbol{x}\|\|d \boldsymbol{y}\|} \\
& =\frac{\left(1+\varepsilon_{11}\right) \varepsilon_{12}+\left(1+\varepsilon_{22}\right) \varepsilon_{21}+\varepsilon_{31} \varepsilon_{32}}{\left[\left(1+\varepsilon_{11}\right)^{2}+\varepsilon_{21}^{2}+\varepsilon_{31}^{2}\right]^{1 / 2}\left[\varepsilon_{12}^{2}+\left(1+\varepsilon_{22}\right)^{2}+\varepsilon_{32}^{2}\right]^{1 / 2}} \\
& =\frac{\varepsilon_{12}+\varepsilon_{11} \varepsilon_{12}+\varepsilon_{21}+\varepsilon_{21} \varepsilon_{22}+\varepsilon_{31} \varepsilon_{32}}{\left[1+\varepsilon_{11}^{2}+2 \varepsilon_{11}+\varepsilon_{21}^{2}+\varepsilon_{31}^{2}\right]^{1 / 2}\left[\varepsilon_{12}^{2}+1+\varepsilon_{22}^{2}+2 \varepsilon_{22}+\varepsilon_{32}^{2}\right]^{1 / 2}} .
\end{aligned}
$$

Neglecting the terms of order greater than 1 (products of components $\varepsilon_{i j}$ ) and taking into account the symmetry $\varepsilon_{i j}=\varepsilon_{j i}$, one obtains ( $\mathbf{B} 2.157$ )

$$
\begin{aligned}
\cos \gamma_{12} & =\frac{d \boldsymbol{x} \cdot d \boldsymbol{y}}{\|d \boldsymbol{x}\|\|d \boldsymbol{y}\|} \approx \frac{2 \varepsilon_{12}}{\left[1+2 \varepsilon_{11}\right]^{1 / 2}\left[1+2 \varepsilon_{22}\right]^{1 / 2}} \approx \frac{2 \varepsilon_{12}}{\left(1+\varepsilon_{11}\right)\left(1+\varepsilon_{22}\right)} \\
& \approx 2 \varepsilon_{12}\left(1-\varepsilon_{11}\right)\left(1-\varepsilon_{22}\right)=2 \varepsilon_{12}\left(1-\varepsilon_{11}-\varepsilon_{22}-\varepsilon_{11} \varepsilon_{22}\right) \approx 2 \varepsilon_{12} .
\end{aligned}
$$

Thus we have

$$
\cos \gamma_{12}=2 \varepsilon_{12}
$$

and consequently

$$
\cos \gamma_{12}=\sin \varphi_{12} \approx \varphi_{12}=2 \varepsilon_{12},
$$

as $\gamma_{12}+\varphi_{12}=\pi / 2$.


Figure 2.2 Deformation of an infinitesimal element.

## Solution 2.15

Let $A B C D$ be an infinitesimal element with sides $d x_{1} d x_{2}$ given in figure 2.2 ( $\mathbf{B} 2.23$ in the book).

As angle $\theta_{1}$ is small, we have the approximation

$$
\tan \theta_{1} \approx \theta_{1}
$$

Inspecting figure 2.2 , we find

$$
\theta_{1}=\frac{\frac{\partial u_{2}}{\partial x_{1}} d x_{1}}{d x_{1}+\frac{\partial u_{1}}{\partial x_{1}} d x_{1}}
$$

and thus

$$
\theta_{1}=\frac{\partial u_{2}}{\partial x_{1}}
$$

By a similar line of reasoning, one has

$$
\theta_{2}=\frac{\partial u_{1}}{\partial x_{2}}
$$

We know that

$$
\theta_{1}+\theta_{2}=\phi_{12}=\frac{\pi}{2}-\gamma_{12}
$$

Consequently, one finds

$$
\cos \gamma_{12}=\sin \phi_{12}=\sin \left(\theta_{1}+\theta_{2}\right) \approx \theta_{1}+\theta_{2}=\frac{\partial u_{2}}{\partial x_{1}}+\frac{\partial u_{1}}{\partial x_{2}}=2 \varepsilon_{12} .
$$

By the definition of the relative extension (B2.153), we write

$$
\varepsilon_{11}=\frac{A^{\prime} B^{\prime}-A B}{A B}=\frac{A^{\prime} B^{\prime}-d x_{1}}{d x_{1}}
$$

and

$$
\varepsilon_{22}=\frac{A^{\prime} D^{\prime}-A D}{A D}=\frac{A^{\prime} D^{\prime}-d x_{2}}{d x_{2}} .
$$

Next we obtain the relations

$$
\begin{aligned}
\left(A^{\prime} B^{\prime}\right)^{2}=\left(d x_{1}\left(1+\varepsilon_{11}\right)\right)^{2} & =\left(d x_{1}+\frac{\partial u_{1}}{\partial x_{1}} d x_{1}\right)^{2}+\left(\frac{\partial u_{2}}{\partial x_{1}} d x_{1}\right)^{2} \\
d x_{1}^{2}\left(\varepsilon_{11}^{2}+2 \varepsilon_{11}+1\right) & =d x_{1}^{2}\left(1+2 \frac{\partial u_{1}}{\partial x_{1}}+\left(\frac{\partial u_{1}}{\partial x_{1}}\right)^{2}+\left(\frac{\partial u_{2}}{\partial x_{1}}\right)^{2}\right) \\
2 \varepsilon_{11} & \approx 2 \frac{\partial u_{1}}{\partial x_{1}}
\end{aligned}
$$

Thus

$$
\varepsilon_{11} \approx \frac{\partial u_{1}}{\partial x_{1}}
$$

and for $\left(A^{\prime} D^{\prime}\right)$ one has

$$
\varepsilon_{22} \approx \frac{\partial u_{2}}{\partial x_{2}}
$$

## Solution 2.16

By relation (B2.80), within the framework of small deformations, one has

$$
d s^{2}-d S^{2}=2 E_{i j} d X_{i} d X_{j} \approx 2 \varepsilon_{i j} d x_{i} d x_{j}
$$

1) 

$$
\begin{gathered}
d x_{1}=1, d x_{2}=d x_{3}=0 \\
d s^{2}-d S^{2}=2 \varepsilon_{11} d x_{1}^{2}=4.10^{-3} \\
d S=d x_{1}=1 \Rightarrow d s^{2}=1+4.10^{-3} \Rightarrow d s=1.002 \\
\Rightarrow d s-d S=1.002-1=0.002
\end{gathered}
$$

2) 

$$
\begin{gathered}
d x_{2}=1, d x_{1}=d x_{3}=0 \\
d s^{2}-d S^{2}=2 \varepsilon_{22} d x_{2}^{2}=4.10^{-3} \\
d S=d x_{2}=1 \Rightarrow d s^{2}=1+4.10^{-3} \Rightarrow d s=1.002 \\
\Rightarrow d s-d S=1.002-1=0.002
\end{gathered}
$$

3) 

$$
\begin{gathered}
d x_{1}=d x_{2}=1 \cdot \frac{\sqrt{2}}{2}, d x_{3}=0 \\
d s^{2}-d S^{2}=2 \varepsilon_{11} d x_{1}^{2}+2 \varepsilon_{22} d x_{2}^{2}+2 \cdot 2 \varepsilon_{12} d x_{1} d x+2 \\
=(2 \cdot 2+2 \cdot 2+2 \cdot 2 \cdot 1) \cdot 10^{-3} \frac{\sqrt{2}}{2} \frac{\sqrt{2}}{2}=6 \cdot 10^{-3} \\
d S=1 \Rightarrow d s^{2}=1+6 \cdot 10^{-3} \Rightarrow d s=1.003 \\
\Rightarrow d s-d S=1.003-1=0.003
\end{gathered}
$$

We may also use (B2.153) to get

$$
\varepsilon_{N}=\frac{d s-d S}{d S}=0.003 \Rightarrow d s-d S=0.003 \cdot d S=0.003
$$

## Solution 2.17

Let $P Q=d S$ and $p q=d s$ the lengths given on figure 2.3. By (B2.151), one has

$$
d s^{2}=d S^{2}+2 \varepsilon_{i j} d x_{i} d x_{j} .
$$

In two dimensions, one writes

$$
d s^{2}=d S^{2}+2 \varepsilon_{11} d x_{1}^{2}+2 \varepsilon_{22} d x_{2}^{2}+4 \varepsilon_{12} d x_{1} d x_{2} .
$$

This last relation yields

$$
\begin{equation*}
\frac{d s^{2}-d S^{2}}{d S^{2}}=2 \varepsilon_{11} \frac{d x_{1}^{2}}{d S^{2}}+2 \varepsilon_{22} \frac{d x_{2}^{2}}{d S^{2}}+4 \varepsilon_{12} \frac{d x_{1} d x_{2}}{d S^{2}} . \tag{2.6}
\end{equation*}
$$

Let us set

$$
\varepsilon_{N}=\frac{d s-d S}{d S} .
$$

The left hand side of (2.6) becomes

$$
\begin{align*}
\frac{d s^{2}-d S^{2}}{d S^{2}} & =\frac{d s^{2}}{d S^{2}}-1=\left(\frac{d s}{d S}-1+1\right)^{2}-1=\left(\frac{d s-d S}{d S}+1\right)^{2}-1 \\
& =\left(\varepsilon_{N}+1\right)^{2}-1=\varepsilon_{N}^{2}+2 \varepsilon_{N} \tag{2.7}
\end{align*}
$$



Figure 2.3 Deformation of a linear element.
In the case of small deformation, we allow that

$$
\begin{equation*}
\varepsilon_{N}^{2} \rightarrow 0 \tag{2.8}
\end{equation*}
$$

Referring to figure 2.3, we define

$$
\begin{equation*}
\cos \theta=\frac{d x_{1}}{d S} \quad \text { and } \quad \sin \theta=\frac{d x_{2}}{d S} \tag{2.9}
\end{equation*}
$$

Combining (2.6) and (2.7) and using (2.8) and (2.9), we eventually obtain

$$
\varepsilon_{N}=\varepsilon_{11} \cos ^{2} \theta+\varepsilon_{22} \sin ^{2} \theta+2 \varepsilon_{12} \cos \theta \sin \theta
$$

Utilizing trigonometric identities the last expression becomes

$$
\varepsilon_{N}=\frac{\varepsilon_{11}+\varepsilon_{22}}{2}+\frac{\varepsilon_{11}-\varepsilon_{22}}{2} \cos 2 \theta+2 \varepsilon_{12} \sin 2 \theta
$$

In a direction normal to $\boldsymbol{N}$, one has

$$
\begin{aligned}
& \varepsilon_{N+\pi / 2}=\frac{\varepsilon_{11}+\varepsilon_{22}}{2}+\frac{\varepsilon_{11}-\varepsilon_{22}}{2} \cos 2(\pi / 2+\theta)+2 \varepsilon_{12} \sin 2(\pi / 2+\theta) \\
& \varepsilon_{N+\pi / 2}=\frac{\varepsilon_{11}+\varepsilon_{22}}{2}-\frac{\varepsilon_{11}-\varepsilon_{22}}{2} \cos 2 \theta-2 \varepsilon_{12} \sin 2 \theta
\end{aligned}
$$

## Dynamics of continuous media

## Solution 3.1

Incompressibility (B3.45) requires $\boldsymbol{\nabla} \cdot \boldsymbol{v}=0$. Applying $\partial / \partial x_{i}$ to the velocity field, one has

$$
\begin{aligned}
\frac{\partial v_{i}}{\partial x_{i}} & =\frac{A r^{3} \delta_{i i}-3 A r^{2} x_{i} \frac{\partial r}{\partial x_{i}}}{r^{6}} \\
& =\frac{A r^{3} \delta_{i i}-3 A r^{2} x_{i} \frac{x_{i}}{r}}{r^{6}} \\
& =0
\end{aligned}
$$

since $x_{i} x_{i}=r^{2}$.
It is also possible to solve the problem by computing

$$
\begin{aligned}
\operatorname{div} \boldsymbol{v} & =\frac{\partial v_{i}}{\partial x_{i}}=A \frac{\partial}{\partial x_{i}}\left(x_{i}\left(x_{j} x_{j}\right)^{-3 / 2}\right) \\
& =A\left(\delta_{i i}\left(x_{j} x_{j}\right)^{-3 / 2}-\frac{3}{2} x_{i} \cdot 2\left(x_{j} x_{j}\right)^{-5 / 2} x_{j} \delta_{j i}\right) \\
& =A\left(\delta_{i i}\left(x_{j} x_{j}\right)^{-3 / 2}-\frac{3}{2} x_{j} \cdot 2\left(x_{j} x_{j}\right)^{-5 / 2} x_{j}\right)=0,
\end{aligned}
$$

as $\delta_{i i}=3$.

## Solution 3.2

The mass conservation equation (B3.53) gives

$$
\frac{D \rho}{D t}=-\rho \boldsymbol{\nabla} \cdot \boldsymbol{v}=-\rho \frac{\partial v_{i}}{\partial x_{i}}=-\rho\left(\frac{3}{1+t}\right) .
$$

Integrating from the initial time $t=0$ till the present time $t$, one obtains

$$
\int_{\rho_{0}}^{\rho} \frac{d \rho^{\prime}}{\rho^{\prime}}=-3 \int_{0}^{t} \frac{d t^{\prime}}{1+t^{\prime}}
$$

giving then

$$
\begin{equation*}
\frac{\rho}{\rho_{0}}=\frac{1}{(1+t)^{3}} . \tag{3.1}
\end{equation*}
$$

Let us calculate the trajectoires linked to the velocity field (B2.35). One has

$$
d x_{i}=v_{i} d t=\frac{x_{i}}{(1+t)} d t
$$

Integrating this relation from the initial time $t=0$ till the actual time $t$

$$
\int_{X_{i}}^{x_{i}} \frac{d x_{i}^{\prime}}{x_{i}^{\prime}}=\int_{0}^{t} \frac{d t^{\prime}}{1+t^{\prime}}
$$

yields

$$
\ln \frac{x_{i}}{X_{i}}=\ln (1+t)
$$

and thus

$$
\begin{align*}
x_{1}=X_{1}(1+t), x_{2} & =X_{2}(1+t), x_{3}=X_{3}(1+t) \\
x_{1} x_{2} x_{3} & =X_{1} X_{2} X_{3}(1+t)^{3} . \tag{3.2}
\end{align*}
$$

The combination of (3.1) and (3.2) results in the relation $\rho x_{1} x_{2} x_{3}=$ $\rho_{0} X_{1} X_{2} X_{3}$.

## Solution 3.3

The incompressibility equation in cylindrical coordinates (BA.2) is written as

$$
\frac{1}{r} \frac{\partial\left(r v_{r}\right)}{\partial r}+\frac{1}{r} \frac{\partial v_{\theta}}{\partial \theta}+\frac{\partial v_{z}}{\partial z}=0 .
$$

It is easily deduced that the given velocity field is incompressible.

## Solution 3.4

1) With the matrix $[\sigma]$

$$
[\sigma]=\left(\begin{array}{ccc}
0 & C x_{1} & 0 \\
C x_{1} & 0 & -C x_{2} \\
0 & -C x_{2} & 0
\end{array}\right)
$$

the static equilibrium equation (B3.126) gives

$$
\begin{aligned}
& f_{1}=-\sigma_{1 j, j} \\
&=0 \\
& f_{2}=-\sigma_{2 j, j}=-C \\
& f_{3}=-\sigma_{3 j, j}=C .
\end{aligned}
$$

2) The normal vector to the plane given by surface $f\left(x_{i}\right)=0$, at point $P$ of coordinates $\boldsymbol{x}_{P}=(4,-4,7)$ is defined by its gradient

$$
\begin{aligned}
\boldsymbol{n}\left(\boldsymbol{x}_{P}\right) & =\frac{\nabla f\left(\boldsymbol{x}_{P}\right)}{\left\|\nabla f\left(\boldsymbol{x}_{P}\right)\right\|} \\
& =(2,2,-1)^{T} \frac{1}{\left(2^{2}+2^{2}+1\right)^{\frac{1}{2}}}= \\
& =\frac{1}{3}(2,2,-1)^{T} .
\end{aligned}
$$

For the plane, the stress vector at point $P$ becomes

$$
\begin{aligned}
{[t]_{\text {plane }} } & =[\sigma][n]=\frac{1}{3}\left(\begin{array}{ccc}
0 & 4 C & 0 \\
4 C & 0 & 4 C \\
0 & 4 C & 0
\end{array}\right)\left(\begin{array}{c}
2 \\
2 \\
-1
\end{array}\right) \\
& =\frac{1}{3}\left(\begin{array}{c}
8 C \\
8 C-4 C \\
8 C
\end{array}\right)=\frac{1}{3}\left(\begin{array}{c}
8 C \\
4 C \\
8 C
\end{array}\right) .
\end{aligned}
$$

The normal vector on the sphere at point $P$ is

$$
\begin{aligned}
\boldsymbol{n}\left(\boldsymbol{x}_{P}\right) & =\frac{\nabla f\left(\boldsymbol{x}_{P}\right)}{\left\|\nabla f\left(\boldsymbol{x}_{P}\right)\right\|} \\
& =\left(2 x_{1}, 2 x_{2}, 2 x_{3}\right)^{T} \frac{1}{\left(\left(2 x_{1}\right)^{2}+\left(2 x_{2}\right)^{2}+\left(2 x_{3}\right)^{2}\right)^{\frac{1}{2}}}= \\
& =\left(x_{1}, x_{2}, x_{3}\right)^{T} \frac{1}{\left(\left(x_{1}\right)^{2}+\left(x_{2}\right)^{2}+\left(x_{3}\right)^{2}\right)^{\frac{1}{2}}}= \\
& =(4,-4,7)^{T} \frac{1}{(16+16+49)^{\frac{1}{2}}}= \\
& =\frac{1}{9}(4,-4,7)^{T} .
\end{aligned}
$$

The stress vector on the sphere at point $P$ is

$$
\begin{aligned}
{[t]_{S} } & =[\sigma][n]=\frac{1}{9}\left(\begin{array}{ccc}
0 & 4 C & 0 \\
4 C & 0 & 4 C \\
0 & 4 C & 0
\end{array}\right)\left(\begin{array}{c}
4 \\
-4 \\
7
\end{array}\right) \\
& =\frac{1}{9}\left(\begin{array}{c}
-16 C \\
16 C+28 C \\
-16 C
\end{array}\right)=\frac{1}{9}\left(\begin{array}{c}
-16 C \\
44 C \\
-16 C
\end{array}\right) .
\end{aligned}
$$

3) the principal stresses at point $P$ are the eigenvalues of tensor $\boldsymbol{\sigma}(P)$ obtained solving (B3.111) :

$$
\operatorname{det}(\boldsymbol{\sigma}(P)-\lambda \boldsymbol{I})=0,
$$

where $\lambda=\sigma$.
With the definition of invariants (B1.121), one obtains

$$
\left(\begin{array}{ccc}
0 & 4 C & 0 \\
4 C & 0 & 4 C \\
0 & 4 C & 0
\end{array}\right), I_{1}=0, I_{2}=-32 C^{2}, I_{3}=0
$$

with

$$
I_{2}=-4 \cdot 4 C^{2}-4 \cdot 4 C^{2}=-32 C^{2}
$$

The characteristic equation (B1.123) written for matrix $[\sigma]$ becomes

$$
\lambda^{3}-32 C^{2} \lambda=0
$$

or

$$
\lambda\left(\lambda^{2}-32 C^{2}\right)=0
$$

The first solution is $\lambda=\sigma_{2}=0$. The other solutions are given by

$$
\lambda^{2}=32 C^{2}
$$

or

$$
\sigma_{1,3}= \pm|C| \sqrt{32}
$$

For simplicity, the convention $\sigma_{1} \geqslant \sigma_{2} \geqslant \sigma_{3}$ is adopted.
The maximum shear stress is : $[\boldsymbol{\tau} \cdot \boldsymbol{\sigma} \boldsymbol{n}]_{\max }=\left(\sigma_{1}-\sigma_{3}\right) / 2=C \sqrt{32}$, where $\boldsymbol{\tau}$ is the tangent vector in $P$.
The deviatoric part of $\boldsymbol{\sigma}$ is by definition $\sigma_{i j}^{d}=\sigma_{i j}-\left(\sigma_{k k} / 3\right) \delta_{i j}$. As in our case, the trace of $\sigma$ is zero, the given stress tensor is identical to the deviatoric tensor and the principal deviatoric stresses are equal to the principal stresses.

## Solution 3.5

In absence of body forces, the equilibrium equation is written as

$$
\sigma_{i j, j}=0
$$

When applied to (B3.166), one has

$$
\begin{gathered}
\sigma_{11,1}+\sigma_{12,2}+\sigma_{13,3}=8 x_{1}+8 x_{2}-8 x_{1}-8 x_{2}=0 \\
\sigma_{21,1}+\sigma_{22,2}+\sigma_{23,3}=-\frac{x_{1}}{2}-8 x_{2}+\frac{x_{1}}{2}+8 x_{2}=0 \\
\sigma_{31,1}+\sigma_{32,2}+\sigma_{33,3}=0
\end{gathered}
$$

The given stress field satisfies equilibrium.

## Solution 3.6

The body $\mathcal{B}$ is in equilibrium if the total force and the total moment are both equal to zero.

## Equilibrium of forces

We write the equation with pressure $P$ in the component $F_{1}$ of the force. Then, we use the divergence theorem to obtain (as $P=c s t$ )

$$
F_{1}=-\int_{\partial \omega}\left(P n_{1}+0 n_{2}+0 n_{3}\right) d s=-\int_{\omega} \frac{\partial P}{\partial x_{1}} d v=0 .
$$

Similarly, one obtains $F_{2}=F_{3}=0$. Thus, the equilibrium of forces is satisfied.

## Equilibrium of moments

The moment with respect to the origin $O$ of the force generated by the pressure, at point $\boldsymbol{x}$ is

$$
\begin{gathered}
\boldsymbol{M}(O)=\int_{\partial \omega} \boldsymbol{O M} \times(-P \boldsymbol{n}) d s=-P \int_{\partial \omega} \boldsymbol{O} \boldsymbol{M} \times \boldsymbol{n} d s \\
\boldsymbol{O M} \times \boldsymbol{n}=\left|\begin{array}{lll}
\boldsymbol{e}_{1} & \boldsymbol{e}_{2} & \boldsymbol{e}_{3} \\
x_{1} & x_{2} & x_{3} \\
n_{1} & n_{2} & n_{3}
\end{array}\right| \\
=\boldsymbol{e}_{1}\left(x_{2} n_{3}-x_{3} n_{2}\right)-\boldsymbol{e}_{2}\left(x_{1} n_{3}-x_{3} n_{1}\right)+\boldsymbol{e}_{3}\left(x_{1} n_{2}-x_{2} n_{1}\right) .
\end{gathered}
$$

The first component of moment is expressed as

$$
\begin{aligned}
M_{1}(O) & =-P \boldsymbol{e}_{1} \int_{\partial \omega}\left(x_{2} n_{3}-x_{3} n_{2}\right) d s \\
& =-P \boldsymbol{e}_{1} \int_{\partial \omega}\left(\begin{array}{lll}
0 & -x_{3} & x_{2}
\end{array}\right)\left(\begin{array}{l}
n_{1} \\
n_{2} \\
n_{3}
\end{array}\right) d s .
\end{aligned}
$$

Appying the divergence theorem to this last expression, one obtains

$$
\begin{aligned}
M_{1}(O) & =-P \boldsymbol{e}_{1} \int_{\partial \omega}\left(\begin{array}{lll}
0 & -x_{3} & x_{2}
\end{array}\right)\left(\begin{array}{l}
n_{1} \\
n_{2} \\
n_{3}
\end{array}\right) d s \\
& =-P \boldsymbol{e}_{1} \int_{\omega}\left(\frac{\partial 0}{\partial x_{1}}-\frac{\partial x_{3}}{\partial x_{2}}+\frac{\partial x_{2}}{\partial x_{3}}\right) d s=0 .
\end{aligned}
$$

Similarly, one obtains $M_{2}(O)=M_{3}(O)=0$.
Thus the moment is also equal to zero. The solid body is in equilibrium.

## Solution 3.7

Equation (B3.111) gives

$$
\operatorname{det}([\sigma]-\lambda[I])=0
$$

Applied to (B3.167), one obtains

$$
\left|\left(\begin{array}{ccc}
p-\lambda & p & p \\
p & p-\lambda & p \\
p & p & p-\lambda
\end{array}\right)\right|=0
$$

or via (B1.120)

$$
\begin{equation*}
-\lambda^{3}+I_{1} \lambda^{2}-I_{2} \lambda+I_{3}=0 \tag{3.3}
\end{equation*}
$$

The invariants (B1.121) are

$$
\begin{gathered}
I_{1}=3 p \\
I_{2}=\left|\begin{array}{ll}
p & p \\
p & p
\end{array}\right|+\left|\begin{array}{ll}
p & p \\
p & p
\end{array}\right|+\left|\begin{array}{ll}
p & p \\
p & p
\end{array}\right|=0 \\
I_{3}=\left|\begin{array}{ccc}
p & p & p \\
p & p & p \\
p & p & p
\end{array}\right|=0
\end{gathered}
$$

Equation (3.3) becomes

$$
-\lambda^{3}+3 p \lambda^{2}=0
$$

yielding

$$
\left\{\begin{array}{l}
\sigma_{1}=3 p \\
\sigma_{2}=0 \\
\sigma_{3}=0
\end{array}\right.
$$

The resulting stress state is one of uniform traction ( p is supposed such that $p>0$ ).

Applying (B3.111) to (B3.168), one has

$$
\left|\begin{array}{ccc}
p-\lambda & p & p \\
p & p-\lambda & p \\
p & p & -2 p-\lambda
\end{array}\right|=0
$$

The invariants are $I_{1}=0, I_{2}=-6 p^{2}, I_{3}=0$. Equation (3.3) gives

$$
\lambda\left(6 p^{2}-\lambda^{2}\right)=0
$$

One finds

$$
\left\{\begin{array}{l}
\sigma_{1}=\sqrt{6} p \\
\sigma_{2}=0 \\
\sigma_{3}=-\sqrt{6} p
\end{array}\right.
$$

This is a simple shear stress state because the two principal stresses are equal and opposite.

Applying (B3.111) to (B3.169), one has

$$
\left|\begin{array}{ccc}
0-\lambda & p & p \\
p & 0-\lambda & p \\
p & p & 0-\lambda
\end{array}\right|=0
$$

The invariants are

$$
\begin{gathered}
I_{1}=0 \\
I_{2}=\left|\begin{array}{ll}
0 & p \\
p & 0
\end{array}\right|+\left|\begin{array}{cc}
0 & p \\
p & 0
\end{array}\right|+\left|\begin{array}{cc}
0 & p \\
p & 0
\end{array}\right|=-3 p^{2} \\
I_{3}=\left|\begin{array}{lll}
0 & p & p \\
p & 0 & p \\
p & p & 0
\end{array}\right|=2 p^{3}
\end{gathered}
$$

Equation (3.3) becomes

$$
-\lambda^{3}+3 p^{2} \lambda+2 p^{3}=0
$$

that can be decomposed as

$$
(\lambda-2 p)(\lambda+p)^{2}=0
$$

The principal stresses are

$$
\left\{\begin{array}{l}
\sigma_{1}=2 p \\
\sigma_{2}=\sigma_{3}=-p
\end{array}\right.
$$

The resulting stress state is a three-dimensional stress state.

## Solution 3.8

By the deviatoric tensor definition (B3.123), one has

$$
\begin{equation*}
s_{i j}=\sigma_{i j}-\frac{1}{3} \delta_{i j} \sigma_{k k}=\sigma_{i j}-\frac{1}{3} \delta_{i j} I_{1}(\boldsymbol{\sigma}) . \tag{3.4}
\end{equation*}
$$

The characteristic equation is given by ( $\mathbf{B} 1.123$ )

$$
s^{3}-I_{1}(s) s^{2}+I_{2}(s) s-I_{3}(s)=0
$$

With (3.4), one has for the first invariant

$$
I_{1}(s)=s_{i i}=\sigma_{i i}-\frac{1}{3} \delta_{i i} \sigma_{k k}=\sigma_{i i}-\frac{1}{3} 3 \sigma_{k k}=0 .
$$

Thus

$$
s^{3}+I_{2}(s) s-I_{3}(s)=0 .
$$

In the literature, the next form is used

$$
s^{3}-I_{2}(s) s-I_{3}(s)=0
$$

with

$$
I_{2}(\boldsymbol{s})=-\frac{1}{2}\left(s_{i i} s_{j j}-s_{i j} s_{j i}\right), \quad I_{3}(\boldsymbol{s})=\operatorname{det} \boldsymbol{s} .
$$

The first relation is none other than (B3.171).
One writes successively

$$
\begin{aligned}
I_{2}(\boldsymbol{s}) & =\frac{1}{2} s_{i j} s_{j i}=\frac{1}{2}\left(\sigma_{i j}-\frac{1}{3} \delta_{i j} \sigma_{k k}\right)\left(\sigma_{i j}-\frac{1}{3} \delta_{i j} \sigma_{k k}\right) \\
& =\frac{1}{2}\left(\sigma_{i j} \sigma_{i j}-\frac{1}{3} \delta_{i j} \sigma_{i j} \sigma_{k k}-\frac{1}{3} \delta_{i j} \sigma_{i j} \sigma_{k k}+\frac{1}{9} \delta_{i j} \delta_{i j} \sigma_{n n} \sigma_{k k}\right) \\
& =\frac{1}{2}\left(\sigma_{i j} \sigma_{i j}-\frac{1}{3} \sigma_{j j} \sigma_{k k}-\frac{1}{3} \sigma_{j j} \sigma_{k k}+\frac{1}{3} \sigma_{n n} \sigma_{k k}\right) \\
& =\frac{1}{2}\left(\sigma_{i j} \sigma_{i j}-\frac{1}{3} \sigma_{j j} \sigma_{k k}\right) .
\end{aligned}
$$

Using (B3.116) to replace $\sigma_{i j} \sigma_{i j}$ we obtain

$$
\begin{align*}
I_{2}(\boldsymbol{s}) & =\frac{1}{2}\left(-2 I_{2}(\boldsymbol{\sigma})+I_{1}^{2}(\boldsymbol{\sigma})-\frac{1}{3} I_{1}^{2}(\boldsymbol{\sigma})\right)=+\frac{1}{2}\left(-2 I_{2}(\boldsymbol{\sigma})+\frac{2}{3} I_{1}^{2}(\boldsymbol{\sigma})\right) \\
& =\frac{1}{3} I_{1}^{2}(\boldsymbol{\sigma})-I_{2}(\boldsymbol{\sigma}) . \tag{B3.171}
\end{align*}
$$

For the third invariant, we proceed as follows (cf. (B3.118))

$$
\begin{align*}
I_{3}(\boldsymbol{s}) & =s_{1} s_{2} s_{3}=\left(\sigma_{1}-\frac{1}{3} I_{1}(\boldsymbol{\sigma})\right)\left(\sigma_{2}-\frac{1}{3} I_{1}(\boldsymbol{\sigma})\right)\left(\sigma_{3}-\frac{1}{3} I_{1}(\boldsymbol{\sigma})\right) \\
& =\sigma_{1} \sigma_{2} \sigma_{3}-\frac{1}{3} I_{1}(\boldsymbol{\sigma})\left(\sigma_{1} \sigma_{2}+\sigma_{2} \sigma_{3}+\sigma_{3} \sigma_{1}\right) \\
& +\frac{1}{9} I_{1}^{2}(\boldsymbol{\sigma})\left(\sigma_{1}+\sigma_{2}+\sigma_{3}\right)-\frac{1}{27} I_{1}^{3}(\boldsymbol{\sigma}) \\
& =I_{3}(\boldsymbol{\sigma})-\frac{1}{3} I_{1}(\boldsymbol{\sigma}) I_{2}(\boldsymbol{\sigma})+\frac{1}{9} I_{1}^{3}(\boldsymbol{\sigma})-\frac{1}{27} I_{1}^{3}(\boldsymbol{\sigma}) \\
& =\frac{2}{27} I_{1}^{3}(\boldsymbol{\sigma})-\frac{1}{3} I_{1}(\boldsymbol{\sigma}) I_{2}(\boldsymbol{\sigma})+I_{3}(\boldsymbol{\sigma}) . \tag{B3.172}
\end{align*}
$$

## Solution 3.9

Through the divergence theorem the surface integral is converted into a volume integral

$$
\begin{aligned}
\int_{\partial \omega} P_{i j \ldots} \sigma_{p q} n_{q} d s & =\int_{\omega} \frac{\partial\left(P_{i j \ldots} \sigma_{p q}\right)}{\partial x_{q}} d v \\
& =\int_{\omega}\left[\sigma_{p q} P_{i j \ldots, q}+P_{i j \ldots} \sigma_{p q, q}\right] d v
\end{aligned}
$$

Replacing $\boldsymbol{\nabla} \cdot \boldsymbol{\sigma}$ in this last relation by its expression produced by the momentum conservation law (B3.96), one obtains (B3.173).

## Solution 3.10

The first Piola-Kirchhoff stress tensor is defined by the relation (B3.141) :

$$
\boldsymbol{P}=J \boldsymbol{\sigma} \boldsymbol{F}^{-T}
$$

Multiplying (B3.141) right by $\boldsymbol{F}^{T}$, one has

$$
\boldsymbol{P} \boldsymbol{F}^{T}=J \boldsymbol{\sigma} \boldsymbol{F}^{-T} \boldsymbol{F}^{T}=J \boldsymbol{\sigma}
$$

The transpose of (B3.141) gives

$$
\boldsymbol{P}^{T}=J \boldsymbol{F}^{-1} \boldsymbol{\sigma}^{T}
$$

Multiplying left this last equation by $\boldsymbol{F}$, one finds

$$
\boldsymbol{F} \boldsymbol{P}^{T}=J \boldsymbol{F} \boldsymbol{F}^{-1} \boldsymbol{\sigma}^{T}=J \boldsymbol{\sigma}^{T}=J \boldsymbol{\sigma}
$$

and thus

$$
\begin{equation*}
\boldsymbol{P} \boldsymbol{F}^{T}=\boldsymbol{F} \boldsymbol{P}^{T} \tag{B3.144}
\end{equation*}
$$

## Solution 3.11

Equations (B3.149) and (B2.205) give

$$
\boldsymbol{P}^{*}=\boldsymbol{Q P} ; \boldsymbol{F}^{*}=\boldsymbol{Q F}
$$

Equation (B3.152) yields

$$
\boldsymbol{S}=\boldsymbol{F}^{-1} \boldsymbol{P} \Rightarrow \boldsymbol{P}=\boldsymbol{F} \boldsymbol{S}
$$

Therefore, one has

$$
\boldsymbol{P}^{*}=\boldsymbol{F}^{*} \boldsymbol{S}^{*} \Rightarrow \boldsymbol{Q P}=\boldsymbol{Q} \boldsymbol{F} \boldsymbol{S}^{*} \Rightarrow \boldsymbol{Q}^{T} \boldsymbol{Q P}=\boldsymbol{P}=\boldsymbol{F} \boldsymbol{S}^{*}
$$

However, by definition,

$$
P=F S
$$

Consequently,

$$
\boldsymbol{F S}=\boldsymbol{F} \boldsymbol{S}^{*},
$$

that leads to the result

$$
\boldsymbol{S}=\boldsymbol{S}^{*}
$$

For the symmetry case, one has successively

$$
\boldsymbol{S}=J \boldsymbol{F}^{-1} \boldsymbol{\sigma} \boldsymbol{F}^{-T} \Rightarrow \boldsymbol{S}^{T}=J\left(\boldsymbol{F}^{-T}\right)^{T} \boldsymbol{\sigma}^{T}\left(\boldsymbol{F}^{-1}\right)^{T}=J \boldsymbol{F}^{-1} \boldsymbol{\sigma}^{T} \boldsymbol{F}^{-T}
$$

and since

$$
\boldsymbol{\sigma}=\boldsymbol{\sigma}^{T} \Rightarrow \boldsymbol{S}=\boldsymbol{S}^{T}
$$

## Energetics

## Solution 4.1

With Reynolds transport theorem (B3.23) and the continuity equation (B3.41), one has

$$
\begin{aligned}
\frac{d}{d t} \int_{w} \rho Q d v & =\int_{w}\left(\frac{D(\rho Q)}{D t}+\rho Q \frac{\partial v_{m}}{\partial x_{m}}\right) d v \\
& =\int_{w}\left(\frac{D \rho}{D t} Q+\rho \frac{D Q}{D t}+\rho Q \frac{\partial v_{m}}{\partial x_{m}}\right) d v \\
& =\int_{w}\left(\rho \frac{D Q}{D t}+Q\left(\frac{D \rho}{D t}+\rho \frac{\partial v_{m}}{\partial x_{m}}\right)\right) d v \\
& =\int_{w} \rho \frac{d Q}{D t} d v .
\end{aligned}
$$

Let us calculate the material derivative of the kinetic energy $E_{k}$

$$
\frac{D E_{k}}{D t}=\frac{d}{d t} \int_{\omega} \rho\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}\right) d v=\int_{\omega} \rho \frac{D}{D t}\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}\right) d v=\int_{\omega} \rho \boldsymbol{a} \cdot \boldsymbol{v} d v .
$$

Replacing $\rho \boldsymbol{a}$ by the expression from (B3.96) one obtains equation (B4.26).

## Solution 4.2

By the definition of the kinetic energy (B4.1) and the internal energy (B4.2), one has

$$
\frac{D}{D t}\left(E_{k}(t)+E_{\text {int }}(t)\right)=\frac{D}{D t} \int_{\omega} \rho\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right) d v .
$$

Applying Reynolds theorem (B3.23) we are left with the following development

$$
\begin{aligned}
\frac{D}{D t} \int_{\omega} \rho\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right) d v & = \\
\int_{\omega}\left[\frac{D}{D t}\left(\rho\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right)\right)\right. & \left.+\rho\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right) \boldsymbol{\nabla} \cdot \boldsymbol{v}\right] d v \\
=\int_{\omega}\left[\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right)\left(\frac{D \rho}{D t}+\rho \boldsymbol{\nabla} \cdot \boldsymbol{v}\right)\right. & \left.+\rho \frac{D}{D t}\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right)\right] d v
\end{aligned}
$$

The expression in the second term between parentheses of the last relation is the mass conservation (B3.41). As a consequence, this term vanishes. One finds

$$
\frac{D}{D t} \int_{\omega} \rho\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right) d v=\int_{\omega} \rho\left(\boldsymbol{v} \cdot \frac{D \boldsymbol{v}}{D t}+\frac{D u}{D t}\right) d v
$$

As $D \boldsymbol{v} / D t$ is equal to $\boldsymbol{a}$, the problem is solved.
If we keep the material derivative of $\boldsymbol{v}$, we may write

$$
\begin{aligned}
\int_{\omega} \rho\left(\boldsymbol{v} \cdot \frac{D \boldsymbol{v}}{D t}+\frac{D u}{D t}\right) d v & =\int_{\omega} \rho\left(\frac{D(\boldsymbol{v} \cdot \boldsymbol{v})}{2 D t}+\frac{D u}{D t}\right) d v \\
& =\int_{\omega} \rho \frac{D}{D t}\left(\frac{\boldsymbol{v} \cdot \boldsymbol{v}}{2}+u\right) d v
\end{aligned}
$$

This shows that the material derivative of the integral over a material volume of a quantity equal to $\rho$ times an expression is written in general form as the integral over a material volume of $\rho$ times the material derivative of this expression. This statement constitutes a general theorem.

## Solution 4.3

1) Inequation ( $\mathbf{B} 4.81$ ) is given by

$$
\frac{d}{d t} \int_{\omega} \rho s d v \geq \int_{\omega} \frac{r}{T} d v-\int_{\partial \omega} \frac{\boldsymbol{q} \cdot \boldsymbol{n}}{T} d s
$$

Using the result of the previous exercise, the left hand side becomes

$$
\frac{d}{d t} \int_{\omega} \rho s d v=\int_{\omega} \rho \frac{D s}{D t} d v
$$

By the divergence theorem (B1.228), the surface integral of (B4.81) becomes

$$
\int_{\partial \omega} \frac{q_{k} n_{k}}{T} d s=\int_{\omega} \frac{\partial}{\partial x_{k}}\left(\frac{q_{k}}{T}\right) d v
$$

Applying the localisation principle one obtains

$$
\begin{equation*}
\rho \frac{D s}{D t} \geq \frac{r}{T}-\operatorname{div}\left(\frac{\boldsymbol{q}}{T}\right) \tag{4.1}
\end{equation*}
$$

2) With the equations $(\mathbf{B} 4.23)$ and $(\mathbf{B} 4.25)$, one has

$$
\begin{equation*}
\rho \frac{D u}{D t}-\boldsymbol{\sigma}: \boldsymbol{d}+\operatorname{div} \boldsymbol{q}=r \tag{4.2}
\end{equation*}
$$

Employing the index notation, it is easily shown that

$$
\begin{equation*}
\operatorname{div}\left(\frac{\boldsymbol{q}}{T}\right)=\frac{1}{T} \operatorname{div} \boldsymbol{q}-\frac{\boldsymbol{q} \cdot \boldsymbol{\nabla} T}{T^{2}} \tag{4.3}
\end{equation*}
$$

Bringing (4.2) and (4.3) in (4.1), we find Clausius-Duhem inequality

$$
\begin{equation*}
\rho \frac{D s}{D t} \geq \frac{1}{T}\left(\rho \frac{D u}{D t}-\boldsymbol{\sigma}: \boldsymbol{d}\right)+\frac{1}{T^{2}} \boldsymbol{q} \cdot \boldsymbol{\nabla} T \tag{B4.83}
\end{equation*}
$$

3) If we introduce the Helmholtz specific free energy ,

$$
\begin{equation*}
f=u-T s \tag{B4.84}
\end{equation*}
$$

the Clausius-Duhem inequality $(\mathbf{B} 4.83)$ is rewritten as

$$
\begin{equation*}
\rho \frac{D f}{D t} \leq \operatorname{tr}(\boldsymbol{\sigma} \boldsymbol{L})-\rho s \frac{D T}{D t}-\frac{\boldsymbol{q} \cdot \boldsymbol{\nabla} T}{T} \tag{B4.85}
\end{equation*}
$$

## Solution 4.4

1) The principle of internal energy conservation is given by the relations (B4.23) and (B4.25)

$$
\begin{equation*}
\rho \frac{D u}{D t}=\boldsymbol{\sigma}: \boldsymbol{d}-\operatorname{div} \boldsymbol{q}+r \tag{4.4}
\end{equation*}
$$

For a perfect fluid, the term $\boldsymbol{\sigma}: \boldsymbol{d}$ becomes $-p \operatorname{tr} \boldsymbol{d}=-p \boldsymbol{\nabla} \cdot \boldsymbol{v}$.
2) Substituting the enthalpy definition in (4.4), one has

$$
\begin{equation*}
\rho \frac{D h}{D t}-\rho \frac{D}{D t}\left(\frac{p}{\rho}\right)=-p \frac{\partial v_{i}}{\partial x_{i}}-\frac{\partial q_{i}}{\partial x_{i}}+r \tag{4.5}
\end{equation*}
$$

The development of the term $\frac{D}{D t}\left(\frac{p}{\rho}\right)$ gives

$$
\frac{D}{D t}\left(\frac{p}{\rho}\right)=\frac{1}{\rho} \frac{D p}{D t}-\frac{p}{\rho^{2}} \frac{D \rho}{D t}
$$

Substituting this result in (4.5), one finds

$$
\begin{aligned}
\rho \frac{D h}{D t} & =\frac{D p}{D t}-\frac{p}{\rho} \frac{D \rho}{D t}-p \frac{\partial v_{i}}{\partial x_{i}}-\frac{\partial q_{i}}{\partial x_{i}}+r \\
& =\frac{D p}{D t}-\frac{p}{\rho}\left(\frac{D \rho}{D t}+\rho \frac{\partial v_{i}}{\partial x_{i}}\right)-\frac{\partial q_{i}}{\partial x_{i}}+r \\
& =\frac{D p}{D t}-\frac{\partial q_{i}}{\partial x_{i}}+r
\end{aligned}
$$

where we have used the mass conservation law (B3.41).
3) If moreover, the flow is adiabatic, i. e. $\boldsymbol{q}=\mathbf{0}$ and $r=0$, one has

$$
\rho \frac{D h}{D t}=\frac{D p}{D t}
$$

## Solution 4.5

As rigid body rotation implies by $(\mathbf{B 4 . 6 1 )} \boldsymbol{c}=\mathbf{0}$, it follows that $\dot{\boldsymbol{c}}=\mathbf{0}$. By (B2.211) and taking (B4.63) into account one has in

$$
\boldsymbol{v}^{*}=\dot{\boldsymbol{c}}+\dot{\boldsymbol{Q}} \boldsymbol{x}+\boldsymbol{Q} \boldsymbol{v}=\boldsymbol{Q} \boldsymbol{v}+\boldsymbol{\Omega} \boldsymbol{x}
$$

With (B2.60), one finds

$$
\begin{equation*}
\boldsymbol{v}^{*}=\boldsymbol{v}+\boldsymbol{\omega} \times \boldsymbol{x} \tag{4.6}
\end{equation*}
$$

The vector $\boldsymbol{\omega}$ is the dual vector of $\boldsymbol{\Omega}$, that expresses the angular velocity of rigid body rotation. We rewrite ( $\mathbf{B} 4.48$ )

$$
\begin{align*}
\rho \dot{u}-\boldsymbol{\sigma}: \nabla \boldsymbol{v}+\operatorname{div} \boldsymbol{q}-r & +\boldsymbol{v} \cdot(\rho \boldsymbol{a}-\boldsymbol{\operatorname { d i v }} \boldsymbol{\sigma}-\rho \boldsymbol{b}) \\
& +\left(\frac{1}{2} \boldsymbol{v} \cdot \boldsymbol{v}+u\right)(\dot{\rho}+\rho \operatorname{div} \boldsymbol{v})=0 \tag{4.7}
\end{align*}
$$

(where the notation $\dot{\rho}$ designates the material derivative in Lagrangian description), with starred quantities. We first replace $\boldsymbol{v}^{*}$ by its value (4.6) and from the resulting equation, we substract (4.7). Using the relations (B4.49)-(B4.53), (B2.212), (B2.213), one obtains

$$
\begin{aligned}
-\boldsymbol{\sigma}: \boldsymbol{\Omega} & +\frac{(\boldsymbol{\omega} \times \boldsymbol{x}) \cdot(\boldsymbol{\omega} \times \boldsymbol{x})}{2}(\dot{\rho}+\rho \operatorname{div} \boldsymbol{v})+(\boldsymbol{\omega} \times \boldsymbol{x}) \cdot \boldsymbol{v}(\dot{\rho}+\rho \operatorname{div} \boldsymbol{v}) \\
& +(\boldsymbol{\omega} \times \boldsymbol{x}) \cdot(\rho \boldsymbol{a}-\boldsymbol{\operatorname { d i v }} \boldsymbol{\sigma}-\rho \boldsymbol{b})=0
\end{aligned}
$$

This must be true for any rigid body rotation and thus we deduce the equations of mass and momentum conservation. The remaining term $\boldsymbol{\sigma}$ : $\boldsymbol{\Omega}$ must vanish. Due to the antisymmetric character of $\boldsymbol{\Omega}$, this imposes the symmetry of $\boldsymbol{\sigma}$.

# Constitutive equations : basic principles 

## Solution 5.1

The vector field $\boldsymbol{u}$ is spatially objective and satisfies relation (B2.197)

$$
\boldsymbol{u}^{*}=\boldsymbol{Q u} .
$$

One has on the other hand that

$$
(\nabla \boldsymbol{u})_{i j}^{*}=\frac{\partial u_{i}^{*}}{\partial x_{j}^{*}}=\frac{\partial u_{i}^{*}}{\partial x_{k}} \frac{\partial x_{k}}{\partial x_{j}^{*}} .
$$

By (B2.195), one finds

$$
\frac{\partial x_{j}^{*}}{\partial x_{k}}=Q_{j k}
$$

and its inverse $\partial x_{k} / \partial x_{j}^{*}$ is $Q_{k j}^{-1}=Q_{k j}^{T}$. Thus one has

$$
(\nabla \boldsymbol{u})_{i j}^{*}=\frac{\partial u_{i}^{*}}{\partial x_{j}^{*}}=Q_{i l} \frac{\partial u_{l}}{\partial x_{k}} Q_{k j}^{T}
$$

or

$$
\begin{equation*}
(\nabla \boldsymbol{u})^{*}=\boldsymbol{Q} \nabla \boldsymbol{u} \boldsymbol{Q}^{T} \tag{B5.61}
\end{equation*}
$$

## Solution 5.2

Through the relation (B2.213) and definitions (B2.181) and (B2.183) of $\boldsymbol{d}$ and $\dot{\boldsymbol{\omega}}$, respectively, and taking into account equation (B2.56)

$$
\dot{\boldsymbol{Q}} \boldsymbol{Q}^{T}+\boldsymbol{Q} \dot{\boldsymbol{Q}}^{T}=0,
$$

one writes successively

$$
\begin{aligned}
\left(\boldsymbol{L}^{*}\right)^{T} & =\boldsymbol{Q}\left(\boldsymbol{L}^{T} \boldsymbol{Q}^{T}+\dot{\boldsymbol{Q}}^{T}\right), \\
\boldsymbol{d}^{*} & =\frac{1}{2}\left(\boldsymbol{L}^{*}+\left(\boldsymbol{L}^{*}\right)^{T}\right)=\boldsymbol{Q} \frac{1}{2}\left(\boldsymbol{L}+\boldsymbol{L}^{T}\right) \boldsymbol{Q}^{T}+\frac{1}{2}\left(\dot{\boldsymbol{Q}} \boldsymbol{Q}^{T}+\boldsymbol{Q} \dot{\boldsymbol{Q}}^{T}\right) \\
& =\boldsymbol{Q} \boldsymbol{d} \boldsymbol{Q}^{T}, \\
\dot{\boldsymbol{\omega}}^{*} & =\frac{1}{2}\left(\boldsymbol{L}^{*}-\left(\boldsymbol{L}^{*}\right)^{T}\right)=\boldsymbol{Q} \frac{1}{2}\left(\boldsymbol{L}-\boldsymbol{L}^{T}\right) \boldsymbol{Q}^{T}+\frac{1}{2}\left(\dot{\boldsymbol{Q}} \boldsymbol{Q}^{T}-\boldsymbol{Q} \dot{\boldsymbol{Q}}^{T}\right) \\
& =\boldsymbol{Q} \dot{\boldsymbol{\omega}} \boldsymbol{Q}^{T}+\dot{\boldsymbol{Q}} \boldsymbol{Q}^{T} .
\end{aligned}
$$

Tensor $\boldsymbol{d}$ is spatially objective and tensor $\dot{\boldsymbol{\omega}}$ is not. Indeed, if two observers record the rotation rate of a continuous media, their observations differ by a quantity equal to their relative rotation rate.

## Solution 5.3

Note first that $D / D t^{*}=D / D t$. Tensor $\boldsymbol{T}$ being spatially objective, if we take the material derivative, one obtains successively

$$
\frac{D \boldsymbol{T}^{*}}{D t}=\frac{D\left(\boldsymbol{Q} \boldsymbol{T} \boldsymbol{Q}^{T}\right)}{D t}=\frac{D \boldsymbol{Q}}{D t} \boldsymbol{T} \boldsymbol{Q}^{T}+\boldsymbol{Q} \frac{D \boldsymbol{T}}{D t} \boldsymbol{Q}^{T}+\boldsymbol{Q} \boldsymbol{T} \frac{D \boldsymbol{Q}^{T}}{D t}
$$

This shows that the material derivative of a second order tensor is not spatially objective.

## Solution 5.4

With the help of results of problems 5.2 and 5.3 , and relation ( $\mathbf{B} 2.56$ ), we write

$$
\begin{aligned}
\dot{\boldsymbol{T}}^{*}+\boldsymbol{T}^{*} \dot{\boldsymbol{\omega}}^{*}-\dot{\boldsymbol{\omega}}^{*} \boldsymbol{T}^{*} & =\dot{\boldsymbol{Q}} \boldsymbol{T} \boldsymbol{Q}^{T}+\boldsymbol{Q} \dot{\boldsymbol{T}} \boldsymbol{Q}^{T}+\boldsymbol{Q} \boldsymbol{T} \dot{\boldsymbol{Q}}^{T} \\
+\boldsymbol{Q} \boldsymbol{T} \boldsymbol{Q}^{T}\left(\boldsymbol{Q} \dot{\boldsymbol{\omega}} \boldsymbol{Q}^{T}+\dot{\boldsymbol{Q}} \boldsymbol{Q}^{T}\right) & -\left(\boldsymbol{Q} \dot{\boldsymbol{\omega}} \boldsymbol{Q}^{T}+\dot{\boldsymbol{Q}} \boldsymbol{Q}^{T}\right) \boldsymbol{Q} \boldsymbol{T} \boldsymbol{Q}^{T} \\
& =\boldsymbol{Q} \dot{\boldsymbol{T}} \boldsymbol{Q}^{T}+\boldsymbol{Q} \boldsymbol{T} \dot{\boldsymbol{\omega}} \boldsymbol{Q}^{T}-\boldsymbol{Q} \dot{\boldsymbol{\omega}} \boldsymbol{T} \boldsymbol{Q}^{T} \\
& =\boldsymbol{Q}(\dot{\boldsymbol{T}}+\boldsymbol{T} \dot{\boldsymbol{\omega}}-\dot{\boldsymbol{\omega}} \boldsymbol{T}) \boldsymbol{Q}^{T}
\end{aligned}
$$

This shows that relation $(\mathbf{B} 5.62)$ is spatially objective.

## Solution 5.5

Equation (B5.64) is nothing else than problem 5.3 solved for $\boldsymbol{T}=\boldsymbol{d}$. Therefore, taking (B2.56) into account, from (B2.213) and its transpose, one has to within a factor 2

$$
\begin{aligned}
\dot{\boldsymbol{d}}^{*}+\boldsymbol{d}^{*} \boldsymbol{L}^{*}+\boldsymbol{L}^{T *} \boldsymbol{d}^{*} & =\dot{\boldsymbol{Q}} \boldsymbol{d} \boldsymbol{Q}^{T}+\boldsymbol{Q} \dot{\boldsymbol{d}} \boldsymbol{Q}^{T}+\boldsymbol{Q} \boldsymbol{d} \dot{\boldsymbol{Q}}^{T} \\
& +\boldsymbol{Q d} \boldsymbol{d} \boldsymbol{Q}^{T}(\boldsymbol{Q} \boldsymbol{L}+\dot{\boldsymbol{Q}}) \boldsymbol{Q}^{T}+\boldsymbol{Q}\left(\boldsymbol{L}^{T} \boldsymbol{Q}^{T}+\dot{\boldsymbol{Q}}^{T}\right) \boldsymbol{Q} \boldsymbol{d} \boldsymbol{Q}^{T} \\
& =\boldsymbol{Q} \dot{\boldsymbol{d}} \boldsymbol{Q}^{T}+\boldsymbol{Q} \boldsymbol{d} \boldsymbol{L} \boldsymbol{Q}^{T}+\boldsymbol{Q} \boldsymbol{L}^{T} \boldsymbol{d} \boldsymbol{Q}^{T}
\end{aligned}
$$

## Classical Constitutive Equations

## Solution 6.1

Equation (B2.88) gives

$$
\boldsymbol{C}=\boldsymbol{F}^{T} \boldsymbol{F}
$$

from which we obtain via (B2.179)

$$
\dot{\boldsymbol{C}}=\dot{\boldsymbol{F}}^{T} \boldsymbol{F}+\boldsymbol{F}^{T} \dot{\boldsymbol{F}}=\boldsymbol{F}^{T} \boldsymbol{L}^{T} \boldsymbol{F}+\boldsymbol{F}^{T} \boldsymbol{L} \boldsymbol{F} .
$$

As for the simple fluid $\boldsymbol{F}=\boldsymbol{I}$, one finds

$$
\dot{\boldsymbol{C}}=\boldsymbol{L}^{T}+\boldsymbol{L}=2 \boldsymbol{d} .
$$

## Solution 6.2

Equation ( $\mathbf{B 4 . 2 3 )}$ is

$$
\rho \frac{D u}{D t}=\boldsymbol{\sigma}:(\boldsymbol{\nabla} \boldsymbol{v})-\operatorname{div} \boldsymbol{q}+r .
$$

The term $\boldsymbol{\sigma}:(\boldsymbol{\nabla} \boldsymbol{v})$ with the constitutive equation (B6.14) becomes with the help of (B4.25)

$$
\boldsymbol{\sigma}: \boldsymbol{L}=\boldsymbol{\sigma}: \boldsymbol{d}=-p \operatorname{tr} \boldsymbol{d}+\lambda(\operatorname{tr} \boldsymbol{d})^{2}+2 \mu(\boldsymbol{d}: \boldsymbol{d}) .
$$

Thus, for the Newtonian viscous fluid, one obtains

$$
\begin{equation*}
\rho \frac{D u}{D t}=-p \operatorname{tr} \boldsymbol{d}+\lambda(\operatorname{tr} \boldsymbol{d})^{2}+2 \mu(\boldsymbol{d}: \boldsymbol{d})-\operatorname{div} \boldsymbol{q}+r . \tag{6.1}
\end{equation*}
$$

The perfect fluid is inviscid, i.e. $\lambda=\mu=0$. One finds

$$
\rho \frac{D u}{D t}=-p \operatorname{tr} \boldsymbol{d}-\operatorname{div} \boldsymbol{q}+r .
$$

If the perfect fluid is an ideal gas, then its internal energy is given by relation ( $\mathbf{B} 6.143$ ) and the previous equation becomes

$$
\rho c_{v} \frac{D T}{D t}=-p \operatorname{tr} \boldsymbol{d}-\operatorname{div} \boldsymbol{q}+r .
$$

## Solution 6.3

The Newtonian viscous incompressible fluid satisfies the constraint $\operatorname{div} \boldsymbol{v}=\operatorname{tr} \boldsymbol{d}=0$. In this case, relation (6.1) yields

$$
\rho \frac{D u}{D t}=2 \mu(\boldsymbol{d}: \boldsymbol{d})-\operatorname{div} \boldsymbol{q}+r .
$$

For the perfect fluid, one finds

$$
\rho \frac{D u}{D t}=-\operatorname{div} \boldsymbol{q}+r .
$$

## Solution 6.4

1) Relation (B2.108) shows that $\boldsymbol{U}$ has $\boldsymbol{A}_{i}$ as eigenvectors. By (B2.109), one writes the spectral decomposition

$$
\boldsymbol{U}=\sum_{i=1}^{3} \lambda_{i} \boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}
$$

With (B2.88), one has successively

$$
\begin{aligned}
\boldsymbol{C}=\boldsymbol{U} \boldsymbol{U} & =\sum_{i=1}^{3} \sum_{j=1}^{3} \lambda_{i}\left(\boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}\right) \lambda_{j}\left(\boldsymbol{A}_{j} \otimes \boldsymbol{A}_{j}\right)= \\
& =\sum_{i=1}^{3} \sum_{j=1}^{3} \lambda_{i} \lambda_{j}\left(\boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}\right)\left(\boldsymbol{A}_{j} \otimes \boldsymbol{A}_{j}\right)= \\
& =\sum_{i=1}^{3} \sum_{j=1}^{3} \lambda_{i} \lambda_{j} \delta_{i j}\left(\boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}\right)= \\
& =\sum_{i=1}^{3} \lambda_{i}^{2}\left(\boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}\right) .
\end{aligned}
$$

The isotropic hyperelastic material has for constitutive equation (B6.51)

$$
\boldsymbol{S}=2 \frac{\partial \widehat{\mathcal{W}}(\boldsymbol{C})}{\partial \boldsymbol{C}} .
$$

With (B6.64), one obtains relation (B6.67) for $\partial \widehat{\mathcal{W}}(\boldsymbol{C}) / \partial \boldsymbol{C}$ and finally, one has for (B6.68)

$$
\boldsymbol{S}=\sum_{i=1}^{3} \frac{1}{\lambda_{i}} \frac{\partial \phi}{\partial \lambda_{i}} \boldsymbol{A}_{i} \otimes \boldsymbol{A}_{i}
$$

which shows that $\boldsymbol{S}$ has $\boldsymbol{A}_{i}$ as eigenvectors.
2) Relation (B2.111) shows that $\boldsymbol{V}$ has $\boldsymbol{b}_{i}=\boldsymbol{R} \boldsymbol{A}_{i}$ as eigenvectors. By (B2.113), one writes the spectral decomposition

$$
\boldsymbol{V}=\sum_{i=1}^{3} \lambda_{i} \boldsymbol{b}_{i} \otimes \boldsymbol{b}_{i}
$$

With(B2.89), one has successively

$$
\begin{aligned}
\boldsymbol{c}=\boldsymbol{V} \boldsymbol{V} & =\sum_{i=1}^{3} \sum_{j=1}^{3} \lambda_{i}\left(\boldsymbol{b}_{i} \otimes \boldsymbol{b}_{i}\right) \lambda_{j}\left(\boldsymbol{b}_{j} \otimes \boldsymbol{b}_{j}\right)= \\
& =\sum_{i=1}^{3} \sum_{j=1}^{3} \lambda_{i} \lambda_{j}\left(\boldsymbol{b}_{i} \otimes \boldsymbol{b}_{i}\right)\left(\boldsymbol{b}_{j} \otimes \boldsymbol{b}_{j}\right)= \\
& =\sum_{i=1}^{3} \sum_{j=1}^{3} \lambda_{i} \lambda_{j} \delta_{i j}\left(\boldsymbol{b}_{i} \otimes \boldsymbol{b}_{i}\right)= \\
& =\sum_{i=1}^{3} \lambda_{i}^{2}\left(\boldsymbol{b}_{i} \otimes \boldsymbol{b}_{i}\right) .
\end{aligned}
$$

This proves that $\boldsymbol{c}$ has $\boldsymbol{b}_{i}$ as eigenvectors.
In section B6.5.1, after long developments, one obtains relation (B6.72)

$$
\boldsymbol{\sigma}=J^{-1}\left(\sum_{i=1}^{3} \lambda_{i} \frac{\partial \phi}{\partial \lambda_{i}} \boldsymbol{b}_{i} \otimes \boldsymbol{b}_{i}\right),
$$

which shows that $\boldsymbol{\sigma}$ has $\boldsymbol{b}_{i}$ as eigenvectors.

## Solution 6.5

Relation (B6.61) multiplied left by $\boldsymbol{F}$ and right by $\boldsymbol{F}^{T}$ gives

$$
\begin{equation*}
\frac{1}{2} \boldsymbol{F} \boldsymbol{S} \boldsymbol{F}^{T}=I_{3} \frac{\partial \Phi}{\partial I_{3}} \boldsymbol{F} \boldsymbol{C}^{-1} \boldsymbol{F}^{T}+\left(\frac{\partial \Phi}{\partial I_{1}}+I_{1} \frac{\partial \Phi}{\partial I_{2}}\right) \boldsymbol{F} \boldsymbol{I} \boldsymbol{F}^{T}-\frac{\partial \Phi}{\partial I_{2}} \boldsymbol{F} \boldsymbol{C} \boldsymbol{F}^{T} . \tag{6.2}
\end{equation*}
$$

With the help of (B2.88) and (B2.89), one has

$$
\boldsymbol{F} \boldsymbol{C}^{-1} \boldsymbol{F}^{T}=\boldsymbol{F} \boldsymbol{F}^{-1} \boldsymbol{F}^{-T} \boldsymbol{F}^{T}=\boldsymbol{I}
$$

and

$$
\boldsymbol{F C} \boldsymbol{F}^{T}=\boldsymbol{F} \boldsymbol{F}^{T} \boldsymbol{F} \boldsymbol{F}^{T}=\boldsymbol{c} \boldsymbol{c}=\boldsymbol{c}^{2}
$$

Equation (6.2) taking (B3.152) into account becomes (B6.63)

$$
\boldsymbol{\sigma}=2 J^{-1}\left(I_{3}(\boldsymbol{c}) \frac{\partial \Phi}{\partial I_{3}(\boldsymbol{c})} \boldsymbol{I}+\left(\frac{\partial \Phi}{\partial I_{1}(\boldsymbol{c})}+I_{1}(\boldsymbol{c}) \frac{\partial \Phi}{\partial I_{2}(\boldsymbol{c})}\right) \boldsymbol{c}-\frac{\partial \Phi}{\partial I_{2}(\boldsymbol{c})} \boldsymbol{c}^{2}\right) .
$$

## Solution 6.6

Relation (B6.61) is

$$
\boldsymbol{S}=2\left(I_{3} \frac{\partial \Phi}{\partial I_{3}} \boldsymbol{C}^{-1}+\left(\frac{\partial \Phi}{\partial I_{1}}+I_{1} \frac{\partial \Phi}{\partial I_{2}}\right) \boldsymbol{I}-\frac{\partial \Phi}{\partial I_{2}} \boldsymbol{C}\right)
$$

The Cayley-Hamilton equation ( $\mathbf{B} 1.123$ ) applied to $\boldsymbol{C}$ gives

$$
\boldsymbol{C}^{3}-I_{1} \boldsymbol{C}^{2}+I_{2} \boldsymbol{C}-I_{3} \boldsymbol{I}=\mathbf{0}
$$

and thus

$$
I_{3} \boldsymbol{C}^{-1}=\boldsymbol{C}^{2}-I_{1} \boldsymbol{C}+I_{2} \boldsymbol{I}
$$

Combining it with (B6.61) one finds

$$
\begin{aligned}
\frac{\boldsymbol{S}}{2} & =\frac{\partial \Phi}{\partial I_{3}}\left(\boldsymbol{C}^{2}-I_{1} \boldsymbol{C}+I_{2} \boldsymbol{I}\right)+\left(\frac{\partial \Phi}{\partial I_{1}}+I_{1} \frac{\partial \Phi}{\partial I_{2}}\right) \boldsymbol{I}-\frac{\partial \Phi}{\partial I_{2}} \boldsymbol{C} \\
& =\left(\frac{\partial \Phi}{\partial I_{1}}+I_{1} \frac{\partial \Phi}{\partial I_{2}}+I_{2} \frac{\partial \Phi}{\partial I_{3}}\right) \boldsymbol{I}-\left(I_{1} \frac{\partial \Phi}{\partial I_{3}}+\frac{\partial \Phi}{\partial I_{2}}\right) \boldsymbol{C}+\frac{\partial \Phi}{\partial I_{3}} \boldsymbol{C}^{2}
\end{aligned}
$$

Relation (B6.63) is given by

$$
\boldsymbol{\sigma}=2 J^{-1}\left(I_{3} \frac{\partial \Phi}{\partial I_{3}} \boldsymbol{I}+\left(\frac{\partial \Phi}{\partial I_{1}}+I_{1} \frac{\partial \Phi}{\partial I_{2}}\right) \boldsymbol{c}-\frac{\partial \Phi}{\partial I_{2}} \boldsymbol{c}^{2}\right)
$$

The Cayley-Hamilton equation (B1.123) applied to $\boldsymbol{c}$ gives

$$
\boldsymbol{c}^{3}-I_{1} \boldsymbol{c}^{2}+I_{2} \boldsymbol{c}-I_{3} \boldsymbol{I}=\mathbf{0}
$$

and thus

$$
\boldsymbol{c}^{2}=I_{1} \boldsymbol{c}-I_{2} \boldsymbol{I}+I_{3} \boldsymbol{c}^{-1}
$$

Combining it with (B6.63) one obtains

$$
\begin{aligned}
\frac{\boldsymbol{\sigma}}{2 J^{-1}} & =I_{3} \frac{\partial \Phi}{\partial I_{3}} \boldsymbol{I}+\left(\frac{\partial \Phi}{\partial I_{1}}+I_{1} \frac{\partial \Phi}{\partial I_{2}}\right) \boldsymbol{c}-I_{1} \frac{\partial \Phi}{\partial I_{2}} \boldsymbol{c}+I_{2} \frac{\partial \Phi}{\partial I_{2}} \boldsymbol{I}-I_{3} \frac{\partial \Phi}{\partial I_{2}} \boldsymbol{c}^{-1} \\
\boldsymbol{\sigma} & =2 J^{-1}\left(\left(I_{2} \frac{\partial \Phi}{\partial I_{2}}+I_{3} \frac{\partial \Phi}{\partial I_{3}}\right) \boldsymbol{I}+\frac{\partial \Phi}{\partial I_{1}} \boldsymbol{c}-I_{3} \frac{\partial \Phi}{\partial I_{2}} \boldsymbol{c}^{-1}\right)
\end{aligned}
$$

## Solution 6.7

According to (B6.80), we can write

$$
\begin{aligned}
\Phi\left(I_{1}, I_{2}, I_{3}\right) & =C_{000}+C_{100}\left(I_{1}-3\right)+C_{010}\left(I_{2}-3\right)+C_{001}\left(I_{3}-1\right) \\
& +C_{111}\left(I_{1}-3\right)\left(I_{2}-3\right)\left(I_{3}-1\right)+\ldots
\end{aligned}
$$

In the reference configuration, one has

$$
\lambda_{1}=\lambda_{2}=\lambda_{3}=1 \text { or } \boldsymbol{C}=\boldsymbol{I}
$$

and then

$$
I_{1}=3, I_{2}=3, I_{3}=1
$$

Consequently, one obtains

$$
\Phi(3,3,1)=C_{000}
$$

If $C_{000}=0$, then the energy is zero, at the reference configuration.
The partial derivatives of $\Phi$ with respect to the invariants are

$$
\begin{aligned}
& \frac{\partial \Phi}{\partial I_{1}}=C_{100}+C_{111}\left(I_{2}-3\right)\left(I_{3}-1\right)+\ldots \\
& \frac{\partial \Phi}{\partial I_{2}}=C_{010}+C_{111}\left(I_{1}-3\right)\left(I_{3}-1\right)+\ldots \\
& \frac{\partial \Phi}{\partial I_{3}}=C_{001}+C_{111}\left(I_{1}-3\right)\left(I_{2}-3\right)+\ldots
\end{aligned}
$$

Thus, at the reference configuration (B6.62) becomes

$$
\frac{\partial \Phi}{\partial I_{1}}+2 \frac{\partial \Phi}{\partial I_{2}}+\frac{\partial \Phi}{\partial I_{3}}=C_{100}+2 C_{010}+C_{001}=0 .
$$

## Solution 6.8

The first equality of (B6.59) gives with the help of (B1.144)

$$
\begin{aligned}
\frac{\partial I_{1}}{\partial \boldsymbol{C}} & =\sum_{i=1}^{3} \frac{\partial\left(\lambda_{1}^{2}+\lambda_{2}^{2}+\lambda_{3}^{2}\right)}{\partial \lambda_{i}^{2}}\left(\boldsymbol{n}_{i} \otimes \boldsymbol{n}_{i}\right) \\
& =\boldsymbol{n}_{1} \otimes \boldsymbol{n}_{1}+\boldsymbol{n}_{2} \otimes \boldsymbol{n}_{2}+\boldsymbol{n}_{3} \otimes \boldsymbol{n}_{3} \\
& =\sum_{1}^{3} \boldsymbol{n}_{i} \otimes \boldsymbol{n}_{i} \\
& =\boldsymbol{I} .
\end{aligned}
$$

The second equality of (B6.59) gives with the help of (B1.144)

$$
\begin{aligned}
\frac{\partial I_{2}}{\partial \boldsymbol{C}} & =\sum_{i=1}^{3} \frac{\partial\left(\lambda_{2}^{2} \lambda_{3}^{2}+\lambda_{3}^{2} \lambda_{1}^{2}+\lambda_{1}^{2} \lambda_{2}^{2}\right)}{\partial \lambda_{i}^{2}}\left(\boldsymbol{n}_{i} \otimes \boldsymbol{n}_{i}\right) \\
& =\left(\lambda_{2}^{2}+\lambda_{3}^{2}\right)\left(\boldsymbol{n}_{1} \otimes \boldsymbol{n}_{1}\right)+\left(\lambda_{3}^{2}+\lambda_{1}^{2}\right)\left(\boldsymbol{n}_{2} \otimes \boldsymbol{n}_{2}\right)+\left(\lambda_{1}^{2}+\lambda_{2}^{2}\right)\left(\boldsymbol{n}_{3} \otimes \boldsymbol{n}_{3}\right) \\
& =\left(\lambda_{1}^{2}+\lambda_{2}^{2}+\lambda_{3}^{2}\right)\left(\boldsymbol{n}_{1} \otimes \boldsymbol{n}_{1}\right)+\left(\lambda_{1}^{2}+\lambda_{2}^{2}+\lambda_{3}^{2}\right)\left(\boldsymbol{n}_{2} \otimes \boldsymbol{n}_{2}\right) \\
& +\left(\lambda_{1}^{2}+\lambda_{2}^{2}+\lambda_{3}^{2}\right)\left(\boldsymbol{n}_{3} \otimes \boldsymbol{n}_{3}\right) \\
& -\lambda_{1}^{2}\left(\boldsymbol{n}_{1} \otimes \boldsymbol{n}_{1}\right)-\lambda_{2}^{2}\left(\boldsymbol{n}_{2} \otimes \boldsymbol{n}_{2}\right)-\lambda_{3}^{2}\left(\boldsymbol{n}_{3} \otimes \boldsymbol{n}_{3}\right) \\
& =I_{1} \boldsymbol{I}-\boldsymbol{C}
\end{aligned}
$$

## Solution 6.9

The pressure in the inflated balloon is given by equation (B6.102)

$$
p_{i}(\lambda)=4 C_{10} \frac{e_{i}}{R} \frac{1}{\lambda}\left(1-\frac{1}{\lambda^{6}}\right) .
$$

The maximum pressure is obtained when the derivative of $p$ with respect to the stretch ratio $\lambda$ vanishes

$$
\frac{\partial p_{i}(\lambda)}{\partial \lambda}=0 .
$$

One has

$$
\frac{d}{d \lambda}\left(\frac{1}{\lambda}-\frac{1}{\lambda^{7}}\right)=-\frac{1}{\lambda^{2}}+\frac{7}{\lambda^{8}}=0
$$

and thus

$$
\begin{gathered}
\lambda^{6}=7 \Rightarrow \lambda=\sqrt[6]{7}=1.383 \\
p_{i}^{\max }=4 C_{10} \frac{e_{i}}{R} \frac{1}{1.383}\left(1-\frac{1}{7}\right)=2.479 \frac{C_{10} e_{i}}{R} .
\end{gathered}
$$

## Solution 6.10

The Ogden energy function is relation (B6.86)

$$
\phi\left(\lambda_{1}, \lambda_{2}, \lambda_{3}\right)=\sum_{i=1}^{N} \frac{\mu_{i}}{\alpha_{i}}\left(\lambda_{1}^{\alpha_{i}}+\lambda_{2}^{\alpha_{i}}+\lambda_{3}^{\alpha_{i}}-3\right) .
$$

The principal stresses are given by relation (B6.78)

$$
\sigma_{k}=-p+\lambda_{k} \frac{\partial \phi}{\partial \lambda_{k}}, \quad k=1,2,3 .
$$

Therefore, one writes

$$
\begin{aligned}
\sigma_{1} & =-p+\sum_{i=1}^{N} \mu_{i} \lambda_{1}^{\alpha_{i}} \\
\sigma_{2} & =-p+\sum_{i=1}^{N} \mu_{i} \lambda_{2}^{\alpha_{i}} \\
\sigma_{3} & =-p+\sum_{i=1}^{N} \mu_{i} \lambda_{3}^{\alpha_{i}}
\end{aligned}
$$

- Case of uniaxial stretch : $\sigma_{1}=\sigma, \sigma_{2}=\sigma_{3}=0$ and $\lambda_{1}=\lambda, \lambda_{2}=\lambda_{3}=$ $\lambda^{-1 / 2}$ (incompressibility). One obtains due to these conditions :

$$
\sigma=-\sum_{i=1}^{N} \mu_{i} \lambda^{-\frac{\alpha_{i}}{2}}+\sum_{i=1}^{N} \mu_{i} \lambda^{\alpha_{i}}=\sum_{i=1}^{N} \mu_{i}\left(\lambda^{\alpha_{i}}-\lambda^{\frac{-\alpha_{i}}{2}}\right)
$$

- Case of biaxial stretch : $\sigma_{1}, \sigma_{2} \neq 0, \sigma_{3}=0, \lambda_{3}=\lambda_{1}^{-1} \lambda_{2}^{-1}$ (incompressibility). One obtains due to these conditions :

$$
\begin{aligned}
\sigma_{1} & =-\sum_{i=1}^{N} \mu_{i}\left(\frac{1}{\lambda_{1} \lambda_{2}}\right)^{\alpha_{i}}+\sum_{i=1}^{N} \mu_{i} \lambda_{1}^{\alpha_{i}}=\sum_{i=1}^{N} \mu_{i}\left(\lambda_{1}^{\alpha_{i}}-\left(\frac{1}{\lambda_{1} \lambda_{2}}\right)^{\alpha_{i}}\right) \\
\sigma_{2} & =-\sum_{i=1}^{N} \mu_{i}\left(\frac{1}{\lambda_{1} \lambda_{2}}\right)^{\alpha_{i}}+\sum_{i=1}^{N} \mu_{i} \lambda_{2}^{\alpha_{i}}=\sum_{i=1}^{N} \mu_{i}\left(\lambda_{2}^{\alpha_{i}}-\left(\frac{1}{\lambda_{1} \lambda_{2}}\right)^{\alpha_{i}}\right)
\end{aligned}
$$

- Case of equibiaxial stretch : $\sigma_{1}=\sigma_{2}=\sigma, \sigma_{3}=0$ (particular case of biaxial stretch) and $\lambda_{1}=\lambda_{2}=\lambda, \lambda_{3}=\lambda^{-2}$ (by incompressibility). One obtains due to these conditions :

$$
\sigma=-\sum_{i=1}^{N} \mu_{i} \lambda^{-2 \alpha_{i}}+\sum_{i=1}^{N} \mu_{i} \lambda^{\alpha_{i}}=\sum_{i=1}^{N} \mu_{i}\left(\lambda^{\alpha_{i}}-\lambda^{-2 \alpha_{i}}\right)
$$

Applying the prescribed values of the terms, namely $N=3, \alpha_{1}=$ $1,3, \alpha_{2}=5, \alpha_{3}=-2, \mu_{1}=0,63 \mathrm{MPa}, \mu_{2}=0,0012 \mathrm{MPa}$ and $\mu_{3}=$ $-0,01 \mathrm{MPa}$, we can plot $\sigma_{1}, \sigma_{2}, \sigma_{3}$ as a function of their corresponding elongation.

Figure 6.1 shows the evolution of the uniaxial stress with respect to its elongation, while figure 6.2 exhibits the evolution of the equibiaxial stress.

## Solution 6.11

The stress tensor has only one non zero component, namely $\sigma_{11}=S$, where $S$ is the traction load per unit surface.

The free energy $f$ is given by ( $\mathbf{B 6} 6.159$ )

$$
\frac{1}{\rho} \sigma_{i j}=\frac{\partial f}{\partial \epsilon_{i j}}
$$



Figure 6.1 Uniaxial stress
and thus

$$
\begin{equation*}
\frac{S}{\rho}=\frac{\partial f}{\partial \epsilon_{11}} . \tag{6.3}
\end{equation*}
$$

By (B6.110), one has

$$
\epsilon_{11}=\frac{\lambda+\mu}{\mu(3 \lambda+2 \mu)} S=\frac{1}{E} S,
$$

where $E$ is the Young's modulus (B6.109). One writes (6.3) as follows

$$
\frac{S}{\rho}=\frac{\partial f}{\partial S} \frac{\partial S}{\partial \epsilon_{11}} .
$$

So

$$
\frac{\partial f}{\partial S}=\frac{S}{\rho E}
$$

and finally we have integrating :

$$
f=\frac{1}{\rho E} \frac{S^{2}}{2}+f_{0}
$$

where $f_{0}$ is the unstressed free energy in natural state.


Figure 6.2 Equibiaxial stress

## Solution 6.12

The development of $f(\boldsymbol{\epsilon}, T)$ in the neighborhood of $\boldsymbol{\epsilon}=\mathbf{0}, T=T_{0}$ is written as:
$\rho f=\rho f_{0}-\rho s_{0}\left(T-T_{0}\right)+\frac{\lambda}{2} \epsilon_{i i} \epsilon_{k k}+\mu \epsilon_{i j} \epsilon_{i j}$

$$
\begin{equation*}
+\epsilon_{i j} c_{i j}\left(T-T_{0}\right)-\frac{\rho c}{2 T_{0}}\left(T-T_{0}\right)^{2}, \tag{B6.165}
\end{equation*}
$$

where we have eliminated all terms of order greater than 2 , and where the coefficients $f_{0}, s_{0}, c_{i j}$ and $c$ are still to be determined. The factors $\rho$ and $\frac{\rho}{T_{0}}$ were added to simplify subsequent steps.

For an isotropic material, $c_{i j}$ must isotropic, of the form $a \delta_{i j}$ with $a$ a scalar. Taking this scalar as $a=-(3 \lambda+2 \mu) \alpha$, with $\alpha$ yet to be determined, one has

$$
\begin{equation*}
c_{i j} \epsilon_{i j}\left(T-T_{0}\right)=-(3 \lambda+2 \mu) \alpha \epsilon_{k k}\left(T-T_{0}\right) . \tag{B6.166}
\end{equation*}
$$

Furthermore, combining (B6.165) and (B6.166) one obtains by (B6.159) the next relation

$$
\begin{aligned}
\sigma_{i j} & =\rho \frac{\partial f}{\partial \epsilon_{i j}} \\
& =\lambda \epsilon_{k k} \delta_{i j}+2 \mu \epsilon_{i j}-(3 \lambda+2 \mu) \alpha\left(T-T_{0}\right) \delta_{i j} .
\end{aligned}
$$

Inverting to obtain $\boldsymbol{\epsilon}$ as a function of $\boldsymbol{\sigma}$, one has

$$
\boldsymbol{\epsilon}=\frac{1}{2 \mu}\left\{\boldsymbol{\sigma}+\left[2 \mu \alpha\left(T-T_{0}\right)-\frac{\lambda}{3 \lambda+2 \mu} \operatorname{tr} \boldsymbol{\sigma}\right] \boldsymbol{I}\right\}
$$

where we have used the relation

$$
\epsilon_{k k}=\frac{1}{3 \lambda+2 \mu} \sigma_{k k}+3 \alpha\left(T-T_{0}\right)
$$

As a reminder, $\alpha$ is the thermal expansion coefficient and has dimensions of the inverse of temperature. If we consider the case of free dilatation without exterior stresses, then $\boldsymbol{\sigma}=\mathbf{0}$ and one has

$$
\boldsymbol{\epsilon}=\alpha\left(T-T_{0}\right) \boldsymbol{I}
$$

As $T=T\left(x_{1}\right), \epsilon_{11}$ is the only non zero component and consequently

$$
\epsilon_{11}=\epsilon_{11}\left(x_{1}\right)=\alpha\left(T_{1}-T_{0}\right) \frac{x_{1}}{L}
$$

## Solution 6.13

To solve this problem, we will use the solution of the second part of exercise 6.2. Introducing the Fourier conduction law $\boldsymbol{q}=-k \boldsymbol{\nabla} T$ in (6.2), one obtains

$$
\rho c_{v} \frac{D T}{D t}=-p \operatorname{tr} \boldsymbol{d}+\operatorname{div}(k \nabla T)+r
$$

By the mass conservation equation (B3.41), one has the equality :

$$
\operatorname{tr} \boldsymbol{d}=-\frac{1}{\rho} \frac{D \rho}{D t}
$$

and the energy equation becomes

$$
\rho c_{v} \frac{D T}{D t}=\frac{p}{\rho} \frac{D \rho}{D t}+\operatorname{div}(k \nabla T)+r
$$

Using the state equation (B6.136), we transform the previous equation in the relation :

$$
\rho c_{v} \frac{D T}{D t}=\frac{D p}{D t}-\rho R \frac{D T}{D t}+\operatorname{div}(k \nabla T)+r .
$$

Finally, one can write taking (B6.141) into account

$$
\rho c_{p} \frac{D T}{D t}=\frac{D p}{D t}+\operatorname{div}(k \nabla T)+r
$$

## Solution 6.14

1) The relations (B6.175)-(B6.176) lead to write

$$
\begin{equation*}
\sigma_{i j}^{d}+\sigma_{0} \delta_{i j}=3 \lambda \varepsilon_{0} \delta_{i j}+2 \mu\left(\varepsilon_{i j}^{d}+\varepsilon_{0} \delta_{i j}\right) . \tag{6.4}
\end{equation*}
$$

Let us recall that the deviatoric tensors have a zero trace

$$
\operatorname{tr} \sigma_{i j}^{d}=\operatorname{tr} \varepsilon_{i j}^{d}=0 .
$$

Therefore, computing the trace of (6.4), one obtains

$$
3 \sigma_{0}=3 \lambda \varepsilon_{0} \cdot 3+2 \mu \varepsilon_{0} \cdot 3
$$

and

$$
\sigma_{0}=3 \lambda \varepsilon_{0}+2 \mu \varepsilon_{0}=(3 \lambda+2 \mu) \varepsilon_{0} .
$$

The definition (B6.119)

$$
K=\frac{3 \lambda+2 \mu}{3}
$$

gives

$$
\sigma_{0}=3 K \varepsilon_{0} .
$$

We rewrite (6.4) successively

$$
\begin{aligned}
\sigma_{i j}^{d}+3 K \varepsilon_{0} \delta_{i j} & =3 \lambda \varepsilon_{0} \delta_{i j}+2 \mu \varepsilon_{i j}^{d}+2 \mu \varepsilon_{0} \delta_{i j} \\
& =(3 \lambda+2 \mu) \varepsilon_{0} \delta_{i j}+2 \mu \varepsilon_{i j}^{d} \\
& =3 K \varepsilon_{0} \delta_{i j}+2 \mu \varepsilon_{i j}^{d} .
\end{aligned}
$$

One finds

$$
\sigma_{i j}^{d}=2 \mu \varepsilon_{i j}^{d} .
$$

2) Let us recall that for a second order symmetric tensor $\boldsymbol{L}$, one has $\boldsymbol{L} \boldsymbol{n}=\lambda \boldsymbol{n}$, where $\lambda$ is the eigenvalue of $\boldsymbol{L}$ and $\boldsymbol{n}$ the corresponding eigenvector (sec. 1.3.8).
For the deviatoric stress tensor $\sigma_{i j}^{d}$, one has

$$
\begin{equation*}
\sigma_{i j}^{d} n_{j}=\lambda n_{i} \tag{6.5}
\end{equation*}
$$

We modify (6.5) as follows

$$
\sigma_{i j}^{d} n_{j}+\sigma_{0} n_{i}=\sigma_{0} n_{i}+\lambda n_{i}=\left(\lambda+\sigma_{0}\right) n_{i}
$$

With the help of (B6.175), one writes

$$
\sigma_{i j}^{d} n_{j}+\sigma_{0} n_{i}=\sigma_{i j}^{d} n_{j}+\sigma_{0} \delta_{i j} n_{j}=\left(\sigma_{i j}^{d}+\sigma_{0} \delta_{i j}\right) n_{j}=\sigma_{i j} n_{j} .
$$

And we obtain $\sigma_{i j} n_{j}=\left(\lambda+\sigma_{0}\right) n_{i}$. This shows that $\sigma_{i j}^{d}$ and $\sigma_{i j}$ have the same eigenvectors.
As regards the displacements, we proceed in a similar fashion. Using (B6.177) in (6.5) gives

$$
\begin{equation*}
\varepsilon_{i j}^{d} n_{j}=\frac{\lambda}{2 \mu} n_{i} . \tag{6.6}
\end{equation*}
$$

This shows that $\sigma_{i j}^{d}$ and $\varepsilon_{i j}^{d}$ have the same eigenvectors and consequently, the same principal directions. Using (B6.176) to rewrite (6.6) leads to the relation

$$
\left(\varepsilon_{i j}-\varepsilon_{0} \delta_{i j}\right) n_{j}=\frac{\lambda}{2 \mu} n_{i}
$$

or

$$
\varepsilon_{i j} n_{j}=\left(\varepsilon_{0}+\frac{\lambda}{2 \mu}\right) n_{i}
$$

Comparing this last relation with (6.6), one concludes that $\varepsilon_{i j}$ and $\varepsilon_{i j}^{d}$ have the same eigenvectors. Finally, as
$\sigma_{i j}^{d}$ and $\sigma_{i j}$ have the same eigenvectors $n_{i}$,
$\sigma_{i j}^{d}$ and $\varepsilon_{i j}^{d}$ have the same eigenvectors $n_{i}$,
$\varepsilon_{i j}$ and $\varepsilon_{i j}^{d}$ have the same eigenvectors $n_{i}$,
we conclude that $\varepsilon_{i j}$ and $\sigma_{i j}$ have the same eigenvectors $n_{i}$ and consequently, the same principal directions.
3) The potential strain energy is defined by the next relation

$$
W(\varepsilon)=\frac{1}{2} \varepsilon_{i j} \sigma_{i j} .
$$

One thus has

$$
W(\boldsymbol{\varepsilon})=\frac{1}{2} \varepsilon_{i j} \sigma_{i j}=\frac{1}{2} \varepsilon_{i j}\left(\lambda \varepsilon_{k k} \delta_{i j}+2 \mu \varepsilon_{i j}\right)=\frac{1}{2} \lambda \varepsilon_{k k}^{2}+\mu \varepsilon_{i j} \varepsilon_{i j} .
$$

With the help of (B6.176), one writes

$$
\begin{align*}
W(\varepsilon) & =\frac{1}{2} \lambda\left(3 \varepsilon_{0}\right)^{2}+\mu\left(\varepsilon_{i j}^{d}+\varepsilon_{0} \delta_{i j}\right)\left(\varepsilon_{i j}^{d}+\varepsilon_{0} \delta_{i j}\right) \\
& =\frac{9}{2} \lambda\left(\varepsilon_{0}\right)^{2}+\mu\left(\varepsilon_{i j}^{d} \varepsilon_{i j}^{d}+3 \varepsilon_{0}^{2}\right) \\
& =\frac{9}{2} \lambda\left(\varepsilon_{0}\right)^{2}+3 \mu\left(\varepsilon_{0}\right)^{2}+\mu \varepsilon_{i j}^{d} \varepsilon_{i j}^{d} \\
& =\frac{9}{2} \frac{3 \lambda+2 \mu}{3} \varepsilon_{0}^{2}+\mu \varepsilon_{i j}^{d} \varepsilon_{i j}^{d} \\
& =\frac{9}{2} K \varepsilon_{0}^{2}+\mu \varepsilon_{i j}^{d} \varepsilon_{i j}^{d} . \tag{B6.178}
\end{align*}
$$

4) For the stability condition

$$
W(\varepsilon)>0
$$

to be satisfied, as relation (B6.178) is composed of two squares, the coefficients must be such that

$$
K>0 \quad \text { et } \quad \mu>0 .
$$

## Solution 6.15

1) With (B6.106),(B6.180) and (B6.181), one writes

$$
\begin{aligned}
\varepsilon_{i j} & =-\frac{\lambda \delta_{i j}}{2 \mu(3 \lambda+2 \mu)} \sigma_{n n}+\frac{\sigma_{i j}}{2 \mu} \\
& =-\frac{\lambda \delta_{i j}}{2 \mu(3 \lambda+2 \mu)} \sigma_{n n}+\frac{\sigma \delta_{i j}}{2 \mu} \\
& =\left(-\frac{3 \lambda}{2 \mu(3 \lambda+2 \mu)}+\frac{1}{2 \mu}\right) \sigma \delta_{i j} \\
& =\frac{-3 \lambda+3 \lambda+2 \mu}{2 \mu(3 \lambda+2 \mu)} \sigma \delta_{i j} \\
& =\frac{\sigma}{3 K} \delta_{i j} \\
& =\varepsilon \delta_{i j} .
\end{aligned}
$$

2) By Hooke's law (B6.104) and (B6.182), one has

$$
\begin{aligned}
\sigma_{i j} & =\lambda \varepsilon_{k k} \delta_{i j}+2 \mu \varepsilon_{i j} \\
& =\frac{\lambda \gamma}{2}\left(m_{k} n_{k}+m_{k} n_{k}\right) \delta_{i j}+\mu \gamma\left(m_{i} n_{j}+m_{j} n_{i}\right) \\
& =0+\mu \gamma\left(m_{i} n_{j}+m_{j} n_{i}\right) \\
& =\tau\left(m_{i} n_{j}+m_{j} n_{i}\right) .
\end{aligned}
$$

3) With (B6.106) and (B6.184), one obtains

$$
\begin{aligned}
\varepsilon_{i j} & =-\frac{\lambda \delta_{i j}}{2 \mu(3 \lambda+2 \mu)} \sigma n_{m} n_{m}+\frac{1}{2 \mu} \sigma n_{i} n_{j} \\
& =-\frac{\lambda \sigma}{2 \mu(3 \lambda+2 \mu)}\left(\delta_{i j}-n_{i} n_{j}+n_{i} n_{j}\right)+\frac{1}{2 \mu} \sigma n_{i} n_{j} \\
& =\left(\frac{-\lambda+3 \lambda+2 \mu}{2 \mu(3 \lambda+2 \mu)}\right) \sigma n_{i} n_{j}-\frac{\lambda \sigma}{2 \mu(3 \lambda+2 \mu)}\left(\delta_{i j}-n_{i} n_{j}\right) \\
& =\left(\frac{(\lambda+\mu) \sigma}{\mu(3 \lambda+2 \mu)}\right) n_{i} n_{j}-\frac{\lambda \sigma}{2 \mu(3 \lambda+2 \mu)}\left(\delta_{i j}-n_{i} n_{j}\right) \\
& =\varepsilon_{n} n_{i} n_{j}+\varepsilon_{T}\left(\delta_{i j}-n_{i} n_{j}\right) .
\end{aligned}
$$

## Introduction to Solid Mechanics

## Solution 7.1

With (B7.18), equations (B7.21)-(B7.23) yield

$$
\begin{aligned}
\sigma_{11} & =\frac{E}{(1+\nu)(1-2 \nu)}\left(\frac{\partial u_{1}}{\partial x_{1}}(1-\nu)+\nu \frac{\partial u_{2}}{\partial x_{2}}\right) \\
\sigma_{22} & =\frac{E}{(1+\nu)(1-2 \nu)}\left(\frac{\partial u_{2}}{\partial x_{2}}(1-\nu)+\nu \frac{\partial u_{1}}{\partial x_{1}}\right) \\
\sigma_{12} & =\frac{E}{2(1+\nu)}\left(\frac{\partial u_{1}}{\partial x_{2}}+\frac{\partial u_{2}}{\partial x_{1}}\right) .
\end{aligned}
$$

Let us evaluate the partial derivatives of the components of the stress tensor with respect to the space variables

$$
\begin{aligned}
\frac{\partial \sigma_{11}}{\partial x_{1}} & =\frac{E}{(1+\nu)(1-2 \nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}(1-\nu)+\nu \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right) \\
\frac{\partial \sigma_{22}}{\partial x_{2}} & =\frac{E}{(1+\nu)(1-2 \nu)}\left(\frac{\partial^{2} u_{2}}{\partial x_{2}^{2}}(1-\nu)+\nu \frac{\partial^{2} u_{1}}{\partial x_{1} \partial x_{2}}\right) \\
\frac{\partial \sigma_{12}}{\partial x_{1}} & =\frac{E}{2(1+\nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{1} \partial x_{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1}^{2}}\right) \\
\frac{\partial \sigma_{12}}{\partial x_{2}} & =\frac{E}{2(1+\nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right) .
\end{aligned}
$$

Inserting these derivatives in the equilibrium equations given by (B7.20)

$$
\begin{aligned}
& \frac{\partial \sigma_{11}}{\partial x_{1}}+\frac{\partial \sigma_{12}}{\partial x_{2}}+f_{1}=0 \\
& \frac{\partial \sigma_{22}}{\partial x_{2}}+\frac{\partial \sigma_{12}}{\partial x_{1}}+f_{2}=0
\end{aligned}
$$

one finds successively

$$
\begin{aligned}
& \frac{\partial \sigma_{11}}{\partial x_{1}}+\frac{\partial \sigma_{12}}{\partial x_{2}}+f_{1}=0 \\
& \frac{E}{(1+\nu)(1-2 \nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}(1-\nu)+\nu \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right) \\
& +\frac{E}{2(1+\nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right)+f_{1}=0 \\
& \frac{2}{1-2 \nu}\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}(1-\nu)+\nu \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right)+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{f_{1}}{\mu}=0 \\
& \frac{2(1-\nu)}{1-2 \nu} \frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{2 \nu}{1-2 \nu} \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{f_{1}}{\mu}=0 \\
& \frac{2(1-\nu)}{1-2 \nu} \frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{1}{1-2 \nu} \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{f_{1}}{\mu}=0 \\
& \frac{1}{1-2 \nu} \frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{1}{1-2 \nu} \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{f_{1}}{\mu}=0 \\
& \mu\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}\right)+\frac{\mu}{1-2 \nu} \frac{\partial}{\partial x_{1}}\left(\frac{\partial u_{1}}{\partial x_{1}}+\frac{\partial u_{2}}{\partial x_{2}}\right)+f_{1}=0 \\
& \mu\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}\right)+(\lambda+\mu) \frac{\partial}{\partial x_{1}}\left(\frac{\partial u_{1}}{\partial x_{1}}+\frac{\partial u_{2}}{\partial x_{2}}\right)+f_{1}=0
\end{aligned}
$$

A similar reasoning leads to (B7.313).

## Solution 7.2

With (B7.41), the equations (B7.43) are written as

$$
\begin{aligned}
\sigma_{11} & =\frac{E}{1-\nu^{2}}\left(\frac{\partial u_{1}}{\partial x_{1}}+\nu \frac{\partial u_{2}}{\partial x_{2}}\right) \\
\sigma_{22} & =\frac{E}{1-\nu^{2}}\left(\frac{\partial u_{2}}{\partial x_{2}}+\nu \frac{\partial u_{1}}{\partial x_{1}}\right) \\
\sigma_{12} & =\frac{E}{2(1+\nu)}\left(\frac{\partial u_{1}}{\partial x_{2}}+\frac{\partial u_{2}}{\partial x_{1}}\right) .
\end{aligned}
$$

Let us evaluate the partial derivatives of the components of the stress tensor with respect to the space variables

$$
\begin{aligned}
\frac{\partial \sigma_{11}}{\partial x_{1}} & =\frac{E}{1-\nu^{2}}\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\nu \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right) \\
\frac{\partial \sigma_{22}}{\partial x_{2}} & =\frac{E}{1-\nu^{2}}\left(\frac{\partial^{2} u_{2}}{\partial x_{2}^{2}}+\nu \frac{\partial^{2} u_{1}}{\partial x_{1} \partial x_{2}}\right) \\
\frac{\partial \sigma_{12}}{\partial x_{1}} & =\frac{E}{2(1+\nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{1} \partial x_{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1}^{2}}\right) \\
\frac{\partial \sigma_{12}}{\partial x_{2}} & =\frac{E}{2(1+\nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right) .
\end{aligned}
$$

The equilibrium equations are given by (B7.20)

$$
\begin{aligned}
& \frac{\partial \sigma_{11}}{\partial x_{1}}+\frac{\partial \sigma_{12}}{\partial x_{2}}+f_{1}=0 \\
& \frac{\partial \sigma_{22}}{\partial x_{2}}+\frac{\partial \sigma_{12}}{\partial x_{1}}+f_{2}=0 .
\end{aligned}
$$

Thus the first equilibrium equation becomes successively

$$
\begin{aligned}
& \frac{E}{1-\nu^{2}}\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\nu \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right)+\frac{E}{2(1+\nu)}\left(\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right)+f_{1}=0 \\
& \frac{2}{1-\nu}\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\nu \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}\right)+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{f_{1}}{\mu}=0 \\
& \frac{2}{1-\nu} \frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{2 \nu}{1-\nu} \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{f_{1}}{\mu}=0 \\
& \frac{2}{1-\nu} \frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}-\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}+\frac{1+\nu}{1-\nu} \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+\frac{f_{1}}{\mu}=0 \\
& \mu\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}\right)+\mu \frac{1+\nu}{1-\nu} \frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\mu \frac{1+\nu}{1-\nu} \frac{\partial^{2} u_{2}}{\partial x_{1} \partial x_{2}}+f_{1}=0 \\
& \mu\left(\frac{\partial^{2} u_{1}}{\partial x_{1}^{2}}+\frac{\partial^{2} u_{1}}{\partial x_{2}^{2}}\right)+\frac{E}{2(1-\nu)} \frac{\partial}{\partial x_{1}}\left(\frac{\partial u_{1}}{\partial x_{1}}+\frac{\partial u_{2}}{\partial x_{2}}\right)+f_{1}=0 .
\end{aligned}
$$

A similar reasoning leads to (B7.315).

## Solution 7.3

The Navier equations (B7.6) are written as

$$
\begin{equation*}
(\lambda+\mu) u_{k, k i}+\mu u_{i, j j}+f_{i}=0 . \tag{7.1}
\end{equation*}
$$

With the $u_{i}$ expression given by (B7.315), one has successively

$$
\mu u_{i, j j}=\frac{\lambda+2 \mu}{\lambda+\mu} g_{i, m m j j}-g_{n, n i j j}
$$

as well as

$$
(\lambda+\mu) u_{k}=\frac{\lambda+2 \mu}{\mu} g_{k, m m}-\frac{\lambda+\mu}{\mu} g_{n, n k} .
$$

Taking twice the derivative of this last relation, one obtains

$$
\begin{aligned}
(\lambda+\mu) u_{k, k i} & =\frac{\lambda+2 \mu}{\mu} g_{k, m m k i}-\frac{\lambda+\mu}{\mu} g_{n, n k k i}= \\
& =\frac{\lambda+2 \mu}{\mu} g_{k, k m m i}-\frac{\lambda+\mu}{\mu} g_{k, k m m i} \\
& =g_{k, k m m i} .
\end{aligned}
$$

Inserting these expressions in the Navier equation (7.1), one has

$$
\begin{aligned}
g_{k, k m m i} & +\frac{\lambda+2 \mu}{\lambda+\mu} g_{i, m m j j}-g_{n, n i j j} \\
& =g_{k, k m m i}+\frac{\lambda+2 \mu}{\lambda+\mu} g_{i, m m j j}-g_{k, k m m i} \\
& =\frac{\lambda+2 \mu}{\lambda+\mu} g_{i, m m j j} \\
& =0
\end{aligned}
$$

if $g_{i, m m n n}=0$.

## Solution 7.4

1) We use identity (B1.238)

$$
\boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{u}=\boldsymbol{\nabla}(\boldsymbol{\nabla} \cdot \boldsymbol{u})-\nabla^{2} \boldsymbol{u}
$$

that gives

$$
\boldsymbol{\nabla}(\boldsymbol{\nabla} \cdot \boldsymbol{u})=\nabla^{2} \boldsymbol{u}+\boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{u}
$$

We introduce this last relation in (B7.7) that becomes

$$
\mu \nabla^{2} \boldsymbol{u}+(\lambda+\mu)\left(\nabla^{2} \boldsymbol{u}+\boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{u}\right)=0 .
$$

Therefore

$$
(\lambda+2 \mu) \nabla^{2} \boldsymbol{u}+(\lambda+\mu) \boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{u}=0
$$

2) From relation (B7.7), one writes

$$
\frac{\mu}{\lambda+\mu} \nabla^{2} \boldsymbol{u}+\nabla(\boldsymbol{\nabla} \cdot \boldsymbol{u})=0
$$

The elasticity constants are linked together. By (B6.109), $2 \nu=$ $\lambda /(\lambda+\mu)$. Thus $\mu /(\lambda+\mu)=1-2 \nu$ and then the result.
3) We use relation (B1.238) in the Navier equation (B7.7)

$$
\mu(\nabla(\boldsymbol{\nabla} \cdot \boldsymbol{u})-\boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{u})+(\lambda+\mu) \nabla(\boldsymbol{\nabla} \cdot \boldsymbol{u})=0
$$

from which we get easily

$$
(\lambda+2 \mu) \nabla(\boldsymbol{\nabla} \cdot \boldsymbol{u})-\mu \boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{u}=0
$$

## Solution 7.5

Taking the divergence of the relation

$$
\begin{equation*}
(\lambda+2 \mu) \nabla^{2} \boldsymbol{u}=\rho \frac{\partial^{2} \boldsymbol{u}}{\partial t^{2}}, \tag{B7.209}
\end{equation*}
$$

one obtains

$$
(\lambda+2 \mu) \operatorname{div} \nabla^{2} \boldsymbol{u}=\rho \frac{\partial^{2}(\operatorname{div} \boldsymbol{u})}{\partial t^{2}}
$$

Afterwards we use (B1.191) to obtain

$$
(\lambda+2 \mu) \nabla^{2}(\operatorname{div} \boldsymbol{u})=\rho \frac{\partial^{2}(\operatorname{div} \boldsymbol{u})}{\partial t^{2}}
$$

## Solution 7.6

With div $\boldsymbol{u}=\varepsilon_{i i}=0$, the motion equations (B7.202) become

$$
\mu \nabla^{2} \boldsymbol{u}=\rho \frac{\partial^{2} \boldsymbol{u}}{\partial t^{2}}
$$

We take the curl of this last relation

$$
\mu \nabla \times\left(\nabla^{2} \boldsymbol{u}\right)=\rho \nabla \times \frac{\partial^{2} \boldsymbol{u}}{\partial t^{2}}
$$

Using (B1.237), one finds

$$
\begin{equation*}
\mu \nabla^{2}(\boldsymbol{\nabla} \times \boldsymbol{u})=\rho \frac{\partial^{2}(\boldsymbol{\nabla} \times \boldsymbol{u})}{\partial t^{2}} \tag{B7.319}
\end{equation*}
$$

## Solution 7.7

The function $\Phi$ is a solution of the problem if it satisfies the biharmonic equation (B7.38). To this end, let us first calculate its Laplacian. One finds (cf. (BA.27))

$$
\begin{aligned}
\nabla^{2} \Phi & =\left(\frac{\partial^{2}}{\partial r^{2}}+\frac{1}{r} \frac{\partial}{\partial r}+\frac{1}{r^{2}} \frac{\partial^{2}}{\partial \theta^{2}}\right) \Phi \\
& =\frac{1}{r^{2}} \frac{\partial^{2} \Phi}{\partial \theta^{2}} \\
& =-\frac{4 B}{r^{2}} \sin 2 \theta
\end{aligned}
$$

Afterwards, one evaluates the double Laplacian. One has

$$
\begin{aligned}
\nabla^{4} \Phi & =\left(\frac{\partial^{2}}{\partial r^{2}}+\frac{1}{r} \frac{\partial}{\partial r}+\frac{1}{r^{2}} \frac{\partial^{2}}{\partial \theta^{2}}\right)\left(-\frac{4 B}{r^{2}} \sin 2 \theta\right) \\
& =-\frac{24 B}{r^{4}} \sin 2 \theta+\frac{1}{r} \frac{8 B}{r^{3}} \sin 2 \theta+\frac{1}{r^{2}} \frac{4 B}{r^{2}} \cdot 2 \cdot 2 \sin 2 \theta \\
& =0
\end{aligned}
$$

The stress components are given by the relations (BA.28)-(BA.30)

$$
\begin{aligned}
\sigma_{r r} & =\frac{1}{r} \frac{\partial \Phi}{\partial r}+\frac{1}{r^{2}} \frac{\partial^{2} \Phi}{\partial \theta^{2}}=-\frac{4 B}{r^{2}} \sin 2 \theta \\
\sigma_{\theta \theta} & =\frac{\partial^{2} \Phi}{\partial r^{2}}=0 \\
\sigma_{r \theta} & =\frac{1}{r^{2}} \frac{\partial \Phi}{\partial \theta}-\frac{1}{r} \frac{\partial^{2} \Phi}{\partial \theta \partial r}=\frac{1}{r^{2}}(A+2 B \cos 2 \theta)
\end{aligned}
$$

The equilibrium of the corner is given by the relation

$$
\int_{-\alpha}^{\alpha}\left(\sigma_{r \theta} r d \theta\right) r-M=0
$$

and thus

$$
\begin{equation*}
\int_{-\alpha}^{\alpha}(A+2 B \cos 2 \theta) d \theta=M \tag{7.2}
\end{equation*}
$$

To the condition (7.2), we must add the one expressing that the corner edges are free and not subjected to any shear stress. One writes

$$
\begin{equation*}
\sigma_{r \theta}(\theta=\alpha)=A+2 B \cos 2 \alpha=0 \tag{7.3}
\end{equation*}
$$

We thus have $A=-2 B \cos 2 \alpha$. Replacing $A$ by this value in (7.2) and integrating, we find

$$
B=\frac{M}{2 \sin 2 \alpha-4 \alpha \cos 2 \alpha}
$$

The radial stress becomes

$$
\sigma_{r r}=-\frac{4 M}{2 \sin 2 \alpha-4 \alpha \cos 2 \alpha} \frac{\sin 2 \theta}{r^{2}}=-\frac{2 C}{r^{2}} \sin 2 \theta
$$

with

$$
C=-\frac{M}{\sin 2 \alpha-2 \alpha \cos 2 \alpha} .
$$

The shear stress is

$$
\sigma_{r \theta}=-\frac{M}{r^{2}} \frac{\cos 2 \alpha-\cos 2 \theta}{\sin 2 \alpha-2 \alpha \cos 2 \alpha} .
$$

## Solution 7.8

(a) One verifies easily that the stress function (B7.322) satisfies the biharmonic equation. From the expressions in the appendix $\mathbf{B}(A .28)$ B(A.30), one obtains

$$
\sigma_{r r}=\frac{2 C \cos \theta}{r}, \quad \sigma_{\theta \theta}=\sigma_{r \theta}=0
$$

To find the $C$ constant, one considers the equilibrium of the slab at a distance $r$ from the origin. Thus

$$
P=C \int_{-\alpha}^{\alpha} \frac{2 \cos \theta}{r} r \cos \theta d \theta=2 C \int_{-\alpha}^{\alpha} \cos ^{2} \theta d \theta .
$$

Therefore

$$
P=2 C\left[\frac{\theta}{2}+\frac{1}{4} \sin 2 \theta\right]_{-\alpha}^{\alpha}=C[2 \alpha+\sin 2 \alpha]
$$

Thus, one finds

$$
C=\frac{P}{2 \alpha+\sin 2 \alpha} .
$$

The stress takes the form

$$
\sigma_{r r}=\frac{2 P \cos \theta}{r(2 \alpha+\sin 2 \alpha)}
$$

(b) Setting $\alpha=\pi / 2$, one obtains the stresses for a plate subject to a linear load

$$
\sigma_{r r}=\frac{2 P}{\pi} \frac{\cos \theta}{r}, \quad \sigma_{\theta \theta}=\sigma_{r \theta}=0
$$

Let us note that the difference of this result with (B7.163) comes from the fact that the directions of $\sigma_{r r}$ are opposite in both problems.

## Introduction to Newtonian Fluid Mechanics

## Solution 8.1

The Navier-Stokes equations are treated in section B8.4. The mass conservation law is given by (B3.44) or (B8.9)

$$
\frac{\partial \rho}{\partial t}+\frac{\partial \rho v_{i}}{\partial x_{i}}=0 .
$$

The momentum conservation equation (B3.96) gives

$$
\frac{\partial \sigma_{i j}}{\partial x_{j}}+\rho b_{i}=\rho \frac{D v_{i}}{D t} .
$$

The Newtonian viscous fluid constitutive equation is written as (B6.14)

$$
\sigma_{i j}=-p \delta_{i j}+\lambda d_{k k} \delta_{i j}+2 \mu d_{i j} .
$$

Inserting (B6.14) in (B3.96), one obtains

$$
\rho \frac{D v_{i}}{D t}=\frac{\partial}{\partial x_{j}}\left(-p \delta_{i j}+\lambda d_{k k} \delta_{i j}\right)+\frac{\partial}{\partial x_{j}}\left(2 \mu d_{i j}\right)+\rho b_{i}
$$

and thus the relation (B8.10)

$$
\rho \frac{D v_{i}}{D t}=-\frac{\partial p}{\partial x_{i}}+\frac{\partial}{\partial x_{i}}\left(\lambda d_{k k}\right)+\frac{\partial}{\partial x_{j}}\left(2 \mu d_{i j}\right)+\rho b_{i}
$$

as

$$
\frac{\partial}{\partial x_{j}}\left(-p+\lambda d_{k k}\right) \delta_{i j}=\frac{\partial}{\partial x_{i}}\left(-p+\lambda d_{k k}\right) .
$$

The energy equation (B8.7), taking the Fourier conduction law into account (B6.123), becomes

$$
\rho c_{v} \frac{D T}{D t}=\frac{p}{\rho} \frac{D \rho}{D t}+\lambda(t r \boldsymbol{d})^{2}+2 \mu \boldsymbol{d}: \boldsymbol{d}+\operatorname{div}(k \boldsymbol{\nabla} T)+r .
$$

If the coefficients $\lambda$ and $\mu$ are constant, the momentum equation gives

$$
\rho \frac{D v_{i}}{D t}=-\frac{\partial p}{\partial x_{i}}+\lambda \frac{\partial}{\partial x_{i}}\left(d_{k k}\right)+\mu \Delta v_{i}+\mu \frac{\partial}{\partial x_{i}}\left(\frac{\partial v_{j}}{\partial x_{j}}\right)+\rho b_{i}
$$

If the flow is incompressible, it results that div $\boldsymbol{v}=\operatorname{tr} \boldsymbol{d}=0$ and the previous equation simplifies itself to

$$
\rho \frac{D v_{i}}{D t}=-\frac{\partial p}{\partial x_{i}}+\mu \Delta v_{i}+\rho b_{i}
$$

## Solution 8.2

The solutions obtained in section B8.7.1, i.e. (B8.51) and (B8.63), for the plane Couette and Poiseuille flows, respectively, result from linear differential equations. As the non-linear terms of the Navier-Stokes equations do not intervene in this problem, one invokes the principle of linear superposition and the solution of the combined plane Couette-Poiseuille flow is written as

$$
v_{1}=-\frac{h^{2}}{2 \mu} \frac{d P}{d x_{1}} \frac{x_{2}}{h}\left(1-\frac{x_{2}}{h}\right)+\frac{U x_{2}}{h} .
$$

The shear stress is

$$
\sigma_{12}=\mu \frac{d v_{1}}{d x_{2}}=-\frac{h}{2} \frac{d P}{d x_{1}}\left(1-\frac{2 x_{2}}{h}\right)+\frac{\mu U}{h} .
$$

Finally, the flow rate is

$$
Q=\int_{0}^{h} v_{1} d x_{2}=-\frac{h^{3}}{12 \mu} \frac{d P}{d x_{1}}+\frac{U h}{2} .
$$

## Solution 8.3

From the geometrical point of view, this flow occurs between two concentric cylinders as in the circular Couette flow. The inner cylinder of radius $R_{1}$ and the outer one of radius $R_{2}$ have a rate of angular rotation $\omega_{1}$ and $\omega_{2}$, respectively. The viscous fluid between the cylinders is also subject to an axial pressure gradient. As the flow is in steady
state $(\partial / \partial t=0)$ and presents a symmetry of revolution $(\partial / \partial \theta=0)$, the velocity profile depends only on $r$. One has

$$
\begin{equation*}
v_{r}=v_{r}(r), v_{\theta}=v_{\theta}(r), v_{z}=v_{z}(r), p=p(r, z) \tag{8.1}
\end{equation*}
$$

As the fluid sticks to the wall, the boundary conditions are

$$
\begin{align*}
& v_{r}\left(R_{1}\right)=v_{r}\left(R_{2}\right)=0, \quad v_{\theta}\left(R_{1}\right)=\omega_{1} R_{1}, v_{\theta}\left(R_{2}\right)=\omega_{2} R_{2} \\
& v_{z}\left(R_{1}\right)=v_{z}\left(R_{2}\right)=0 \tag{8.2}
\end{align*}
$$

By a similar reasoning as the one for the circular Couette flow, it is possible to show that the component $v_{r}$ vanishes everywhere (cf. (B8.98) and (B8.99)). The Navier-Stokes equations in cylindrical coordinates (A.32)-(A.34) become

$$
\begin{align*}
\frac{1}{\rho} \frac{\partial p}{\partial r} & =\frac{v_{\theta}^{2}}{r}  \tag{8.3}\\
\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial v_{\theta}}{\partial r}\right)-\frac{v_{\theta}^{2}}{r} & =0  \tag{8.4}\\
-\frac{\partial p}{\partial z}+\mu\left(\frac{\partial^{2} v_{z}}{\partial r^{2}}+\frac{1}{r} \frac{\partial v_{z}}{\partial r}\right) & =0 \tag{8.5}
\end{align*}
$$

The Couette solution ( $\mathbf{B} 8.102$ ) remains valid

$$
\begin{equation*}
v_{\theta}(r)=A r+\frac{B}{r}=\frac{\omega_{2} R_{2}^{2}-\omega_{1} R_{1}^{2}}{R_{2}^{2}-R_{1}^{2}} r-\frac{\left(\omega_{2}-\omega_{1}\right) R_{1}^{2} R_{2}^{2}}{R_{2}^{2}-R_{1}^{2}} \frac{1}{r} \tag{8.6}
\end{equation*}
$$

Relation (8.3) gives

$$
\begin{equation*}
p=\rho \int_{R_{1}}^{r} \frac{v_{\theta}^{2}}{r^{\prime}} d r^{\prime}+f(z) \tag{8.7}
\end{equation*}
$$

where $v_{\theta}$ is the Couette solution and $f(z)$ is an undetermined function of $z$. Introducing (8.7) in (8.5), one finds

$$
\begin{equation*}
-\frac{d f}{d z}+\mu \frac{1}{r} \frac{d}{d r}\left(r \frac{d v_{z}}{d r}\right)=0 \tag{8.8}
\end{equation*}
$$

As $f$ does depend only on $z$ and $v_{z}$ is only function of $r$, one has

$$
\frac{d f}{d z}=\mu \frac{1}{r} \frac{d}{d r}\left(r \frac{d v_{z}}{d r}\right)=-C
$$

where $C$ is a constant. Here we will refer to the development described in pages B295 and B296 for the integration of $v_{z}$. The solutions are written by taking the boundary conditions (8.2) into account

$$
\begin{align*}
f(z) & =-C z+D  \tag{8.9}\\
v_{z}(r) & =\frac{C}{4 \mu}\left[-r^{2}+\frac{R_{2}^{2}-R_{1}^{2}}{\ln \left(R_{2} / R_{1}\right)} \ln r+\frac{R_{1}^{2} \ln R_{2}-R_{2}^{2} \ln R_{1}}{\ln \left(R_{2} / R_{1}\right)}\right] \tag{8.10}
\end{align*}
$$

The factor $D$ is a constant. The pressure field is given by

$$
\begin{equation*}
p(r, z)=\rho \int_{R_{1}}^{r} \frac{v_{\theta}^{2}}{r^{\prime}} d r^{\prime}-C z+D \tag{8.11}
\end{equation*}
$$

Pressure is known up to a constant $D$, that will set the reference pressure; the pressure gradient $-C$ acts in the direction of the axis and finally, the first term of the right hand side of (8.11) balances the centrifugal force of the rotating fluid. Note that the axial velocity does not depend on the angular velocity of the cylinders, while the azimuthal velocity $v_{\theta}$ is independent of the pressure gradient.

## Solution 8.4

We will refer to the spherical coordinates $(r, \theta, \varphi)$ as in figure $\mathbf{B} 8.20$. The rotation axis of the sphere with the angular velocity $\boldsymbol{\Omega}=\Omega \boldsymbol{e}_{x_{3}}$ is axis $x_{3}$. As a consequence of the problem's symmetries, the velocity field has only a single component such that

$$
\begin{equation*}
\boldsymbol{v}=v_{\varphi}(r, \theta) \boldsymbol{e}_{\varphi} . \tag{8.12}
\end{equation*}
$$

We solve the Stokes equations with the boundary conditions

$$
\begin{align*}
\boldsymbol{v} & =0 \quad \text { in } \quad r=\infty  \tag{8.13}\\
v_{\varphi} & =\Omega R \sin \theta \quad \text { in } \quad r=R . \tag{8.14}
\end{align*}
$$

The form of the boundary conditions (8.14) suggest to search the solution under the form

$$
\begin{equation*}
v_{\varphi}=\Omega R f(r) \sin \theta, \quad p=p_{\infty} \tag{8.15}
\end{equation*}
$$

One verifies that the mass conservation equation (BB.30) is trivially verified by (8.15). The pressure gradient does not intervene in (BB.33) because of axial symmetry $(\partial / \partial \varphi=0)$. One has

$$
\begin{equation*}
\Delta v_{\varphi}-\frac{v_{\varphi}}{r^{2} \sin ^{2} \theta}=\Omega R \sin \theta\left(f^{\prime \prime}+\frac{2 f^{\prime}}{r}-\frac{2 f}{r^{2}}\right)=0 . \tag{8.16}
\end{equation*}
$$

The $f$ solution is written as $f(r)=\sum_{n=-\infty}^{+\infty} C_{n} r^{n}$. This gives

$$
\begin{equation*}
f(r)=C_{1} r+\frac{C_{2}}{r^{2}} . \tag{8.17}
\end{equation*}
$$

The boundary conditions (8.13) and (8.14) impose $C_{1}=0$ and $C_{2}=R^{2}$, respectively. The velocity field around the rotating sphere is

$$
v_{\varphi}=\Omega \frac{R^{3}}{r^{2}} \sin \theta \boldsymbol{e}_{\varphi}
$$

## Solution 8.5

The boundary conditions are

$$
\begin{array}{lll}
v_{\varphi}=\Omega_{1} R_{1} \sin \theta & \text { in } \quad r=R_{1} \\
v_{\varphi}=\Omega_{2} R_{2} \sin \theta & \text { in } \quad r=R_{2} . \tag{8.19}
\end{array}
$$

The considerations of the previous exercise remain valid for the velocity profile search under the form (8.15)

$$
\begin{equation*}
v_{\varphi}=f(r) \sin \theta, \quad p=p_{\infty} . \tag{8.20}
\end{equation*}
$$

We will note that we do not use anymore the factor $\Omega R$, as now we have two radii and two angular velocities to take care of. The equation to solve is thus

$$
\begin{equation*}
\Delta v_{\varphi}-\frac{v_{\varphi}}{r^{2} \sin ^{2} \theta}=\sin \theta\left(f^{\prime \prime}+\frac{2 f^{\prime}}{r}-\frac{2 f}{r^{2}}\right)=0 \tag{8.21}
\end{equation*}
$$

whose solution is written as

$$
\begin{equation*}
f(r)=C_{1} r+\frac{C_{2}}{r^{2}} . \tag{8.22}
\end{equation*}
$$

The imposition of the boundary conditions (8.18) and (8.19) yields

$$
\begin{equation*}
C_{1}=\frac{\Omega_{2} R_{2}^{3}-\Omega_{1} R_{1}^{3}}{R_{2}^{3}-R_{1}^{3}}, \quad C_{2}=\frac{\Omega_{1}-\Omega_{2}}{R_{2}^{3}-R_{1}^{3}} R_{1}^{3} R_{2}^{3} \tag{8.23}
\end{equation*}
$$

## Solution 8.6

We consider the streamline $S A$ from the free surface $S$ toward the orifice $O r$ of the enclosure (cf. figure 8.1) and we apply to it the Bernoulli theorem (B8.223) to obtain

$$
p_{S}+\frac{\rho v_{S}^{2}}{2}+\rho \chi_{S}=p_{A}+\frac{\rho v_{A}^{2}}{2}+\rho \chi_{A} .
$$



Figure 8.1 Enclosure with free surface and orifice.
By (B8.193), one obtains

$$
-g=-\frac{\partial \chi}{\partial x_{3}},
$$

and thus $\chi=g x_{3}+C$. At the free surface the pressure is that of ambiant air; it is the same situation at the orifice. Therefore $p_{S}=p_{A}=p_{\text {air }}$. If we set the origin of the $x_{3}$ axis at the level of the orifice, the contribution of $\rho \chi$ is equal to $C$. For the sake of simplicity, we set $C=0$, while at the free surface $x_{3}=h, \rho \chi_{S}=\rho g h$. On the free surface, the velocity $v_{S}$ is zero (this is especially true when the enclosure is large) and setting $v_{A}=v$ one has

$$
\begin{equation*}
\rho g h=\frac{\rho}{2} v^{2} . \tag{8.24}
\end{equation*}
$$

This gives the sought relation, which is known as Torricelli formula.

## Solution 8.7

From the Navier-Stokes equations (B8.17) without body forces and assuming a velocity field of the form

$$
\begin{equation*}
v_{1}=v_{1}\left(x_{2}, x_{3}\right), \quad v_{2}=v_{3}=0, \tag{8.25}
\end{equation*}
$$

the only relation that gives a non zero contribution is the one related to the $v_{1}$ velocity component. One finds

$$
\begin{equation*}
0=-\frac{\partial p}{\partial x_{1}}+\mu\left(\frac{\partial^{2} v_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} v_{1}}{\partial x_{3}^{2}}\right), \tag{8.26}
\end{equation*}
$$

or

$$
\begin{equation*}
\frac{1}{\mu} \frac{\partial p}{\partial x_{1}}=\frac{\partial^{2} v_{1}}{\partial x_{2}^{2}}+\frac{\partial^{2} v_{1}}{\partial x_{3}^{2}}=C \tag{8.27}
\end{equation*}
$$

as $p=p\left(x_{1}\right)$ and $v_{1}=v_{1}\left(x_{2}, x_{3}\right)$.
On the elliptical wall $v_{1}=0$ and thus $A+B=0$, or $B=-A$. We calculate the second order derivatives of the velocity that are injected in (8.27). One has

$$
2 A\left(\frac{1}{a^{2}}+\frac{1}{b^{2}}\right)=\frac{1}{\mu} \frac{\partial p}{\partial x_{1}}=C
$$

From which we get $A$ and $B=-A$. One obtains a generic solution, that will become particular for a given pressure gradient.

## Solution 8.8



Figure 8.2 Flow between two concentric cylinders, one fixed and the other moving with the velocity $U$.

We work in cylindrical coordinates with the $z$ axis in the direction of the axes of both cylinders (cf. figure 8.2). The only non zero velocity component is clearly $v_{z}$. Moreover $v_{z}=v_{z}(r)$.

The flow is kinematically forced by the displacement of the inner cylinder. No pressure gradient is involved in the fluid motion.

Equation (A.34) gives

$$
\begin{equation*}
\frac{1}{r} \frac{d}{d r}\left(r \frac{d v_{z}}{d r}\right)=0 \tag{8.28}
\end{equation*}
$$

Integrating (8.28), one finds

$$
\begin{equation*}
v_{z}=C_{1} \ln r+C_{2} . \tag{8.29}
\end{equation*}
$$

The boundary conditions are

$$
\begin{align*}
& v_{z}\left(r=R_{1}\right)=U  \tag{8.30}\\
& v_{z}\left(r=R_{2}\right)=0 . \tag{8.31}
\end{align*}
$$

Imposing (8.30) and (8.31) to (8.29), one obtains the velocity field

$$
v_{z}=\frac{U}{\ln \frac{R_{1}}{R_{2}}} \ln \frac{r}{R_{2}}
$$

The only non zero component of the stress tensor is $\sigma_{r z}$ equal to

$$
\begin{equation*}
\sigma_{r z}=\mu\left(\frac{\partial v_{z}}{\partial r}+\frac{\partial v_{r}}{\partial z}\right)=\mu \frac{U}{\ln \frac{R_{1}}{R_{2}}} \frac{1}{r} \tag{8.32}
\end{equation*}
$$

The friction force per unit length that acts on the moving cylinder is given by the integral

$$
\left.\int_{0}^{1} \sigma_{r z}\right|_{r=R 1} 2 \pi R_{1} d z=2 \pi \mu \frac{U}{\ln \frac{R_{1}}{R_{2}}}
$$

